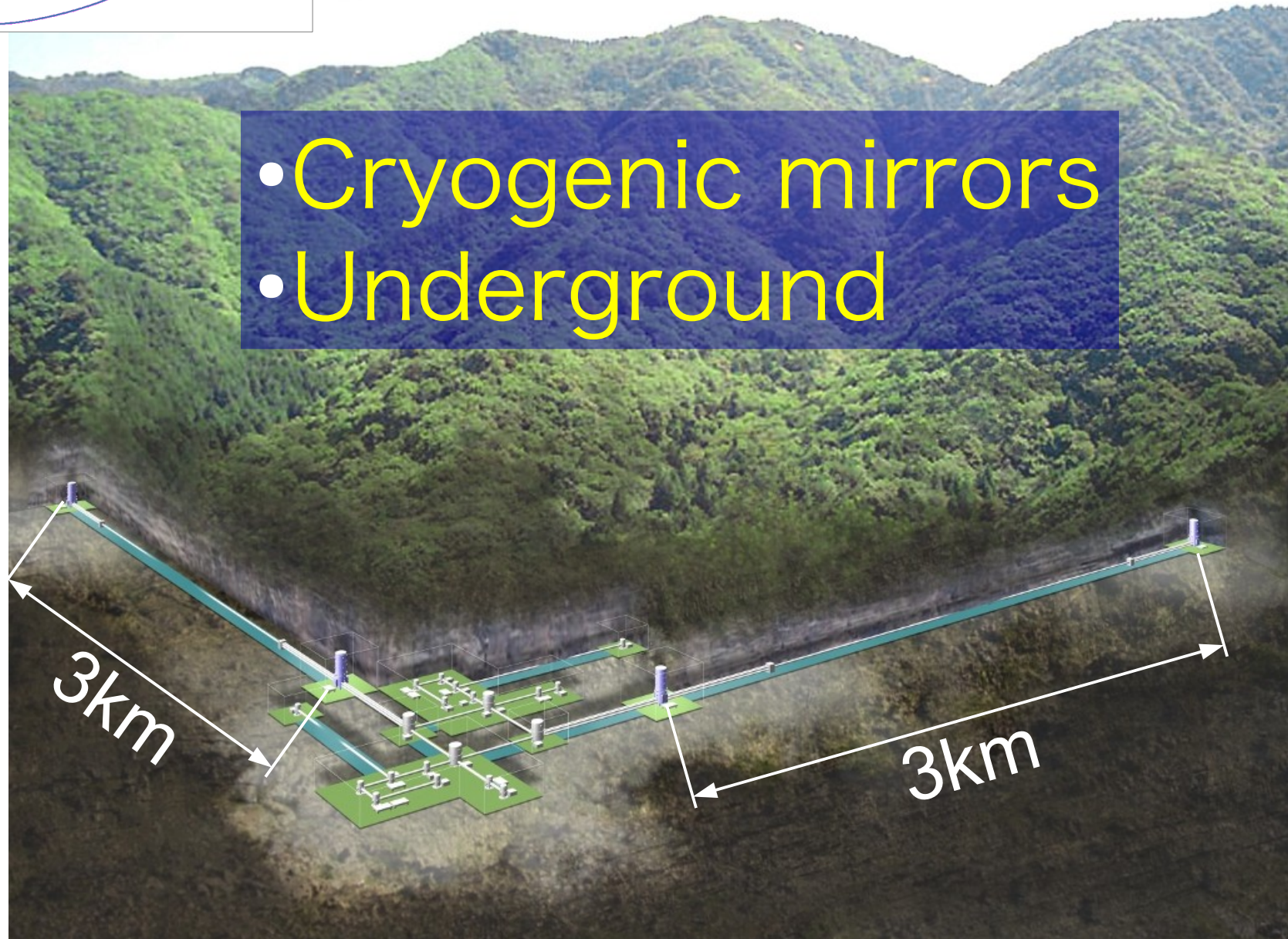


R&D for the gravitational wave detector KAGRA

Yoichi Aso for the KAGRA collaboration
National Astronomical Observatory of Japan
2016/7/13 @ GR21, Columbia University

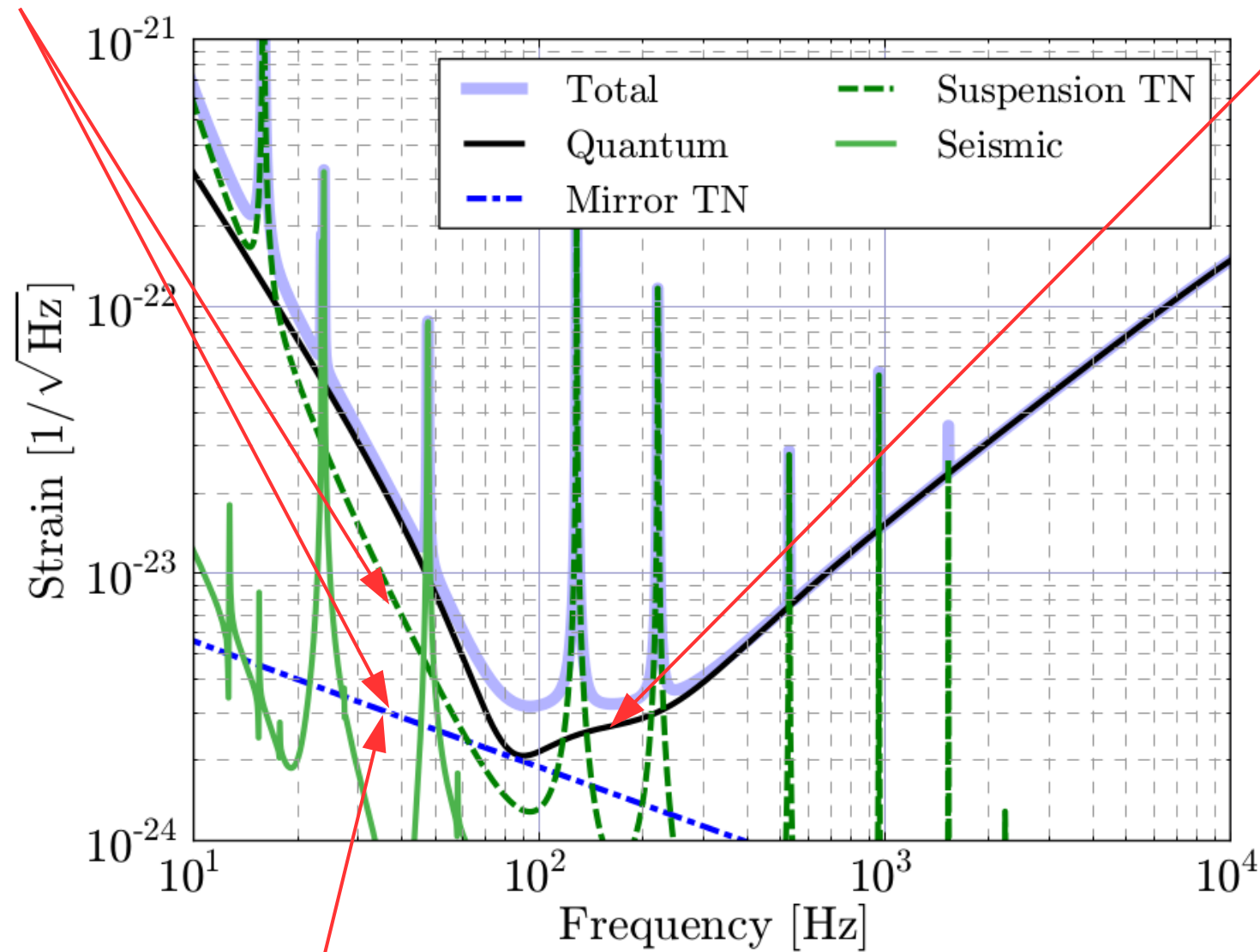


- Cryogenic mirrors
- Underground



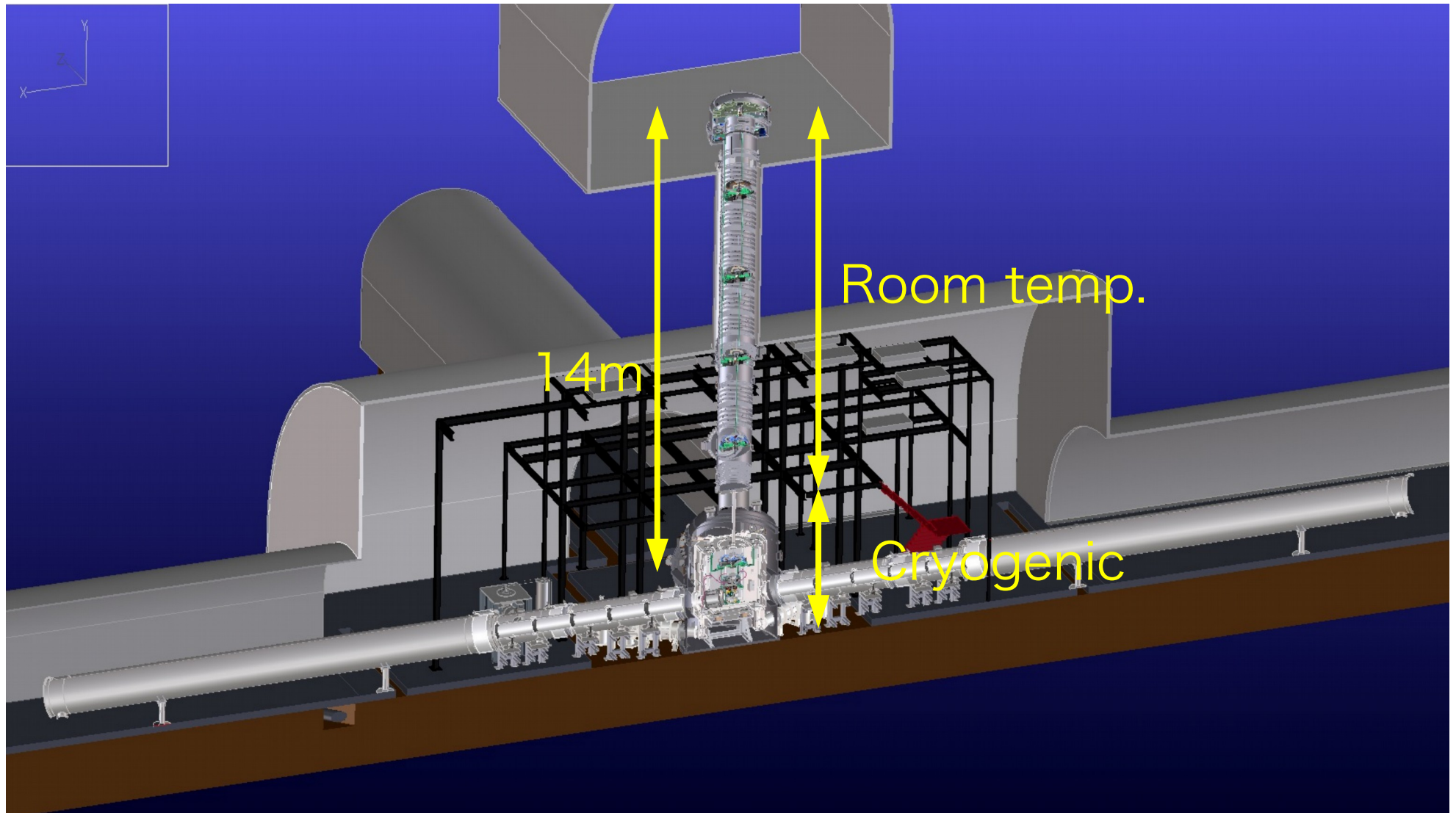
Cryogenic Suspension Mirror Absorption Measurements

300m Filter Cavity

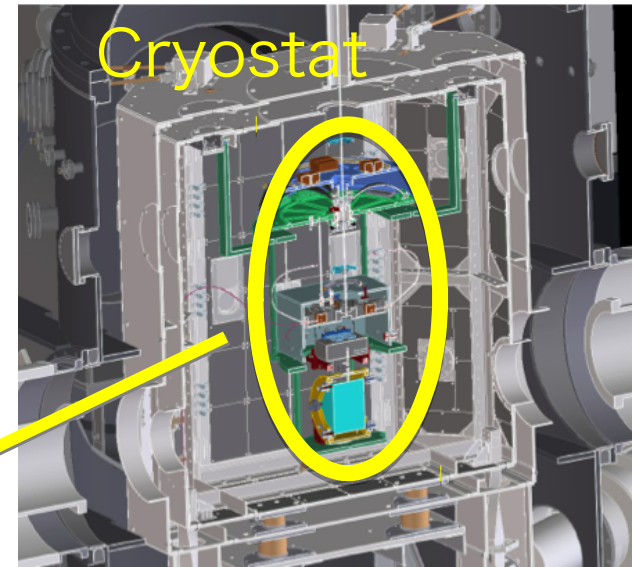
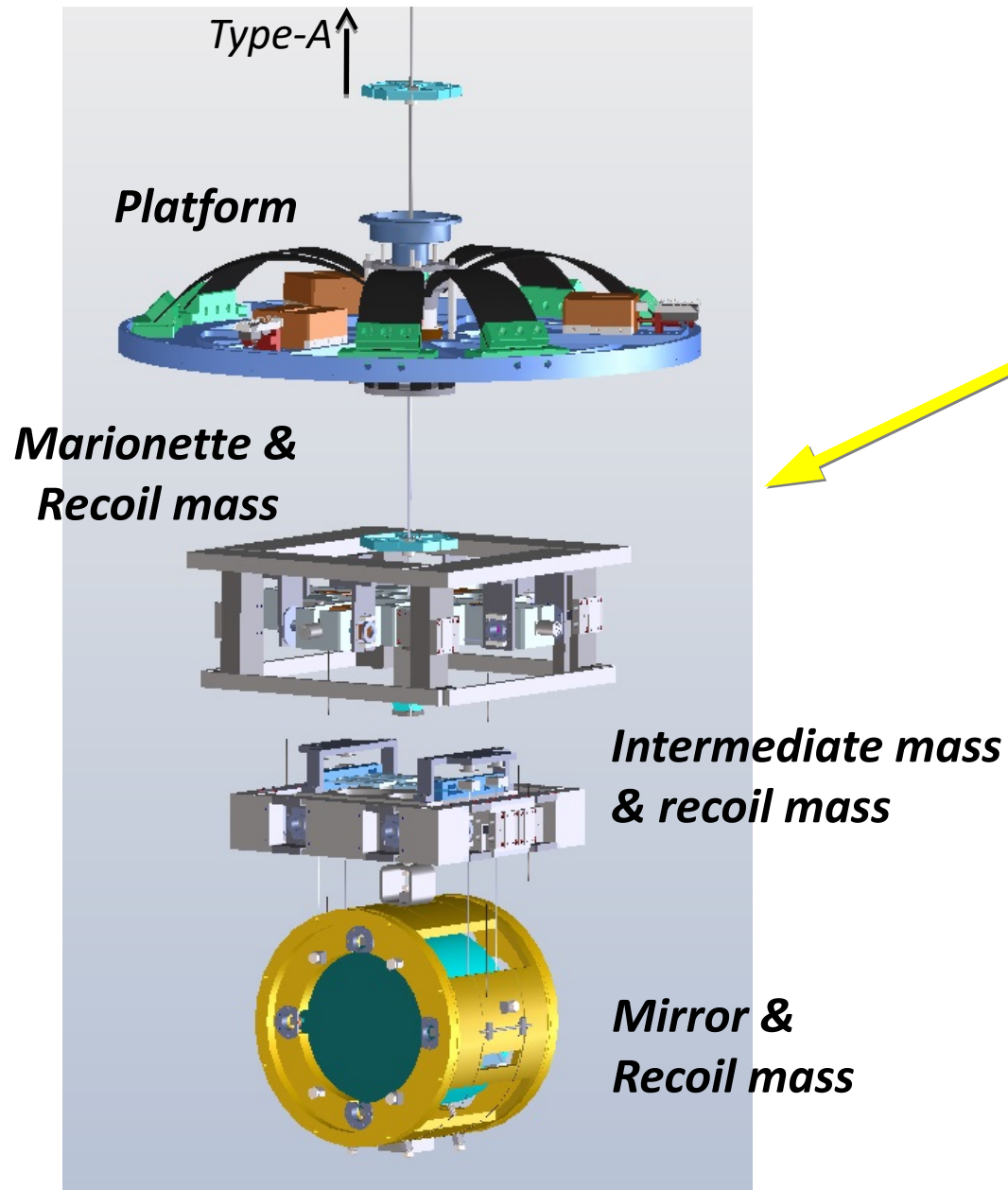


Crystalline Coatings
Cryogenic Thermal Noise Measurements

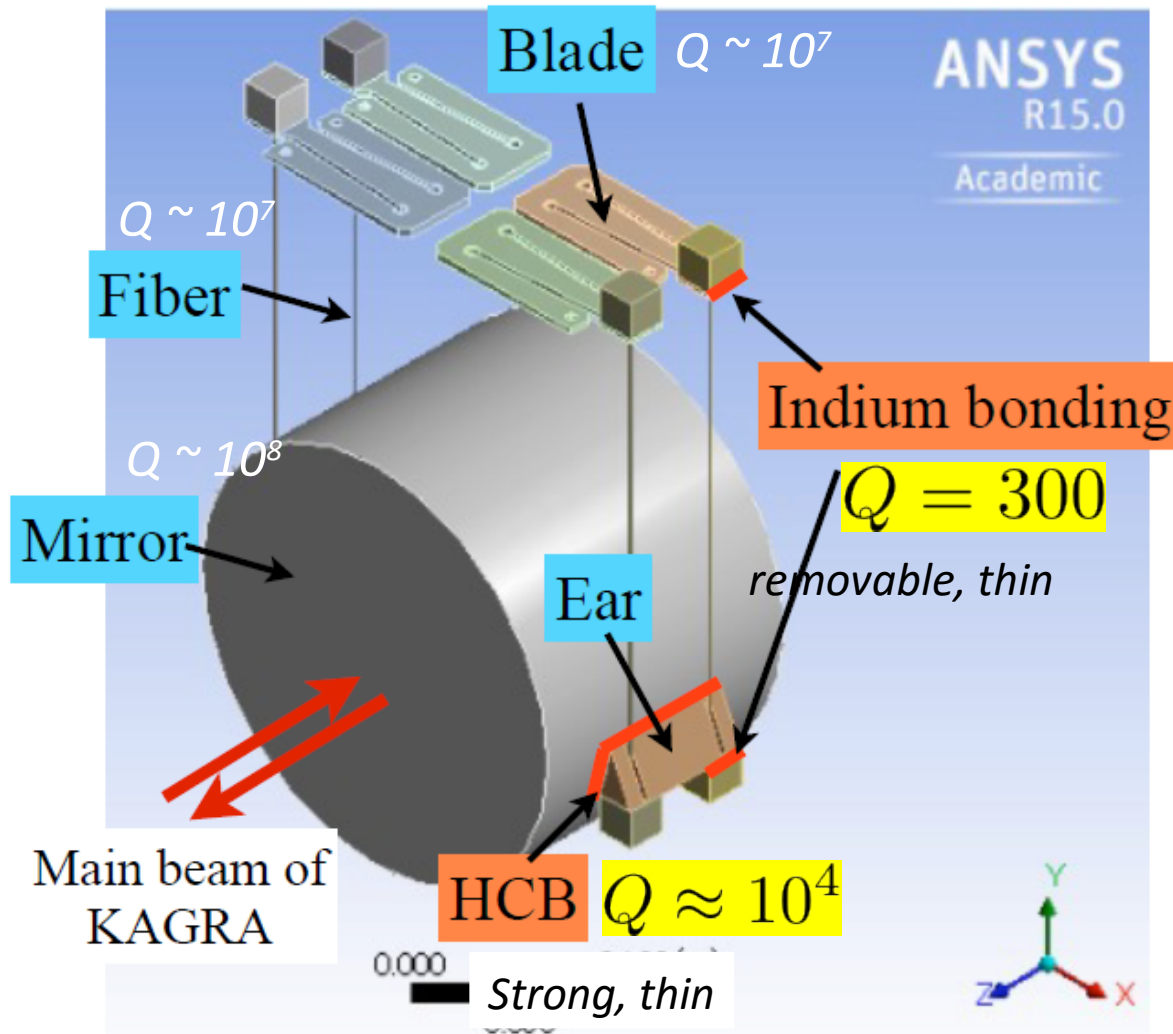
Cryogenic Suspensions



Cryogenic Payload (KEK and ICRR)



Bonding techniques



Semi-monolithic sapphire suspension must be realized to be sufficiently small thermal noises.

HCB and Indium bonding on sapphire satisfies requirements of mechanical Q.

Practical way to do these bonding is under testing.

Installation into KAGRA: 2017 Summer

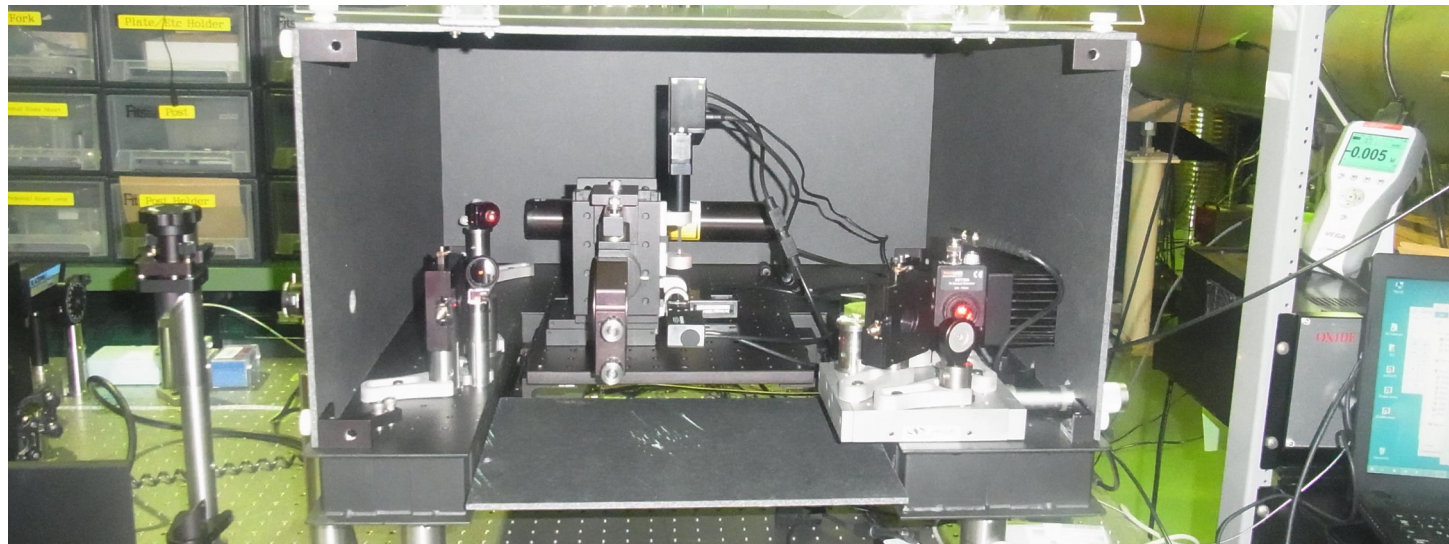
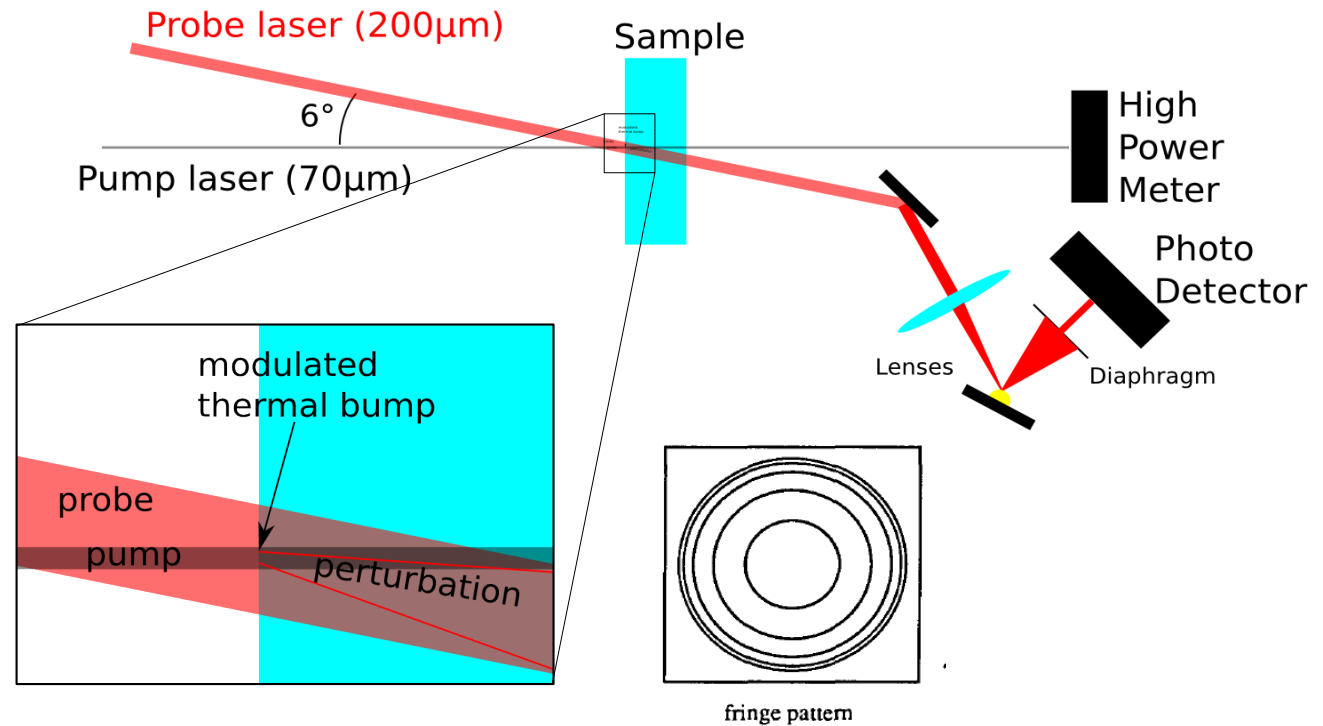
Characterizing mirror absorption

M. Marchio, D. Tatsumi, R. Flaminio

Photo-thermal
Common-path
Interferometry

Sensitivity

- 1 ppm/cm substrate
- 0.1 ppm in coating



Current status

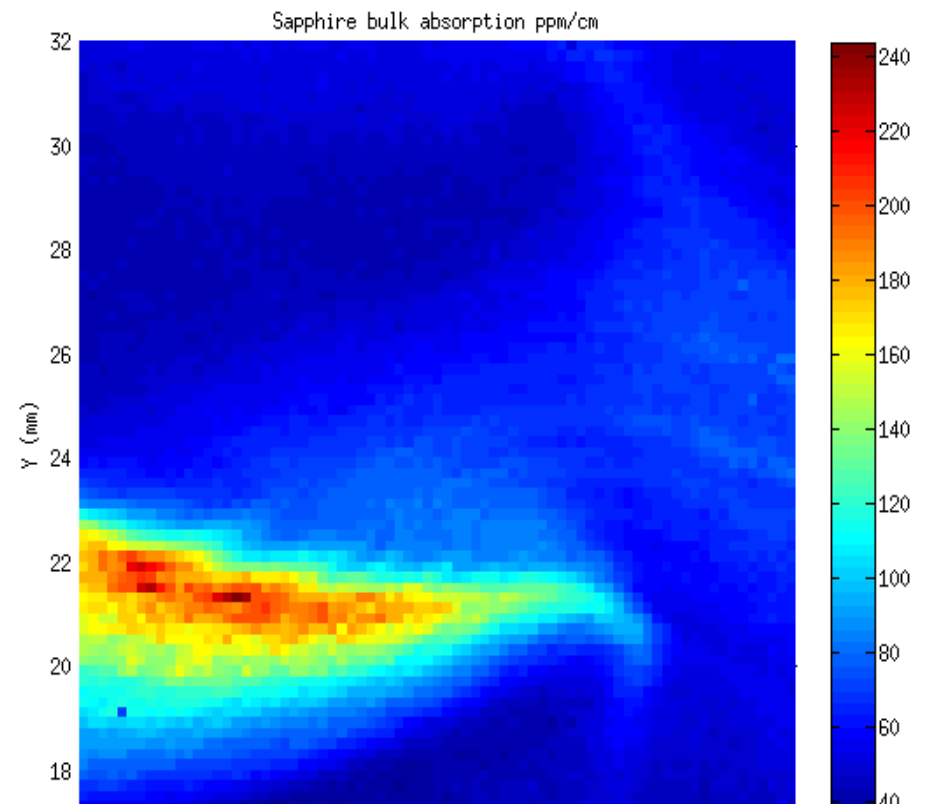
- System calibration/verification measurements: done
- Large sample holder is ready

Measurement of sapphire samples

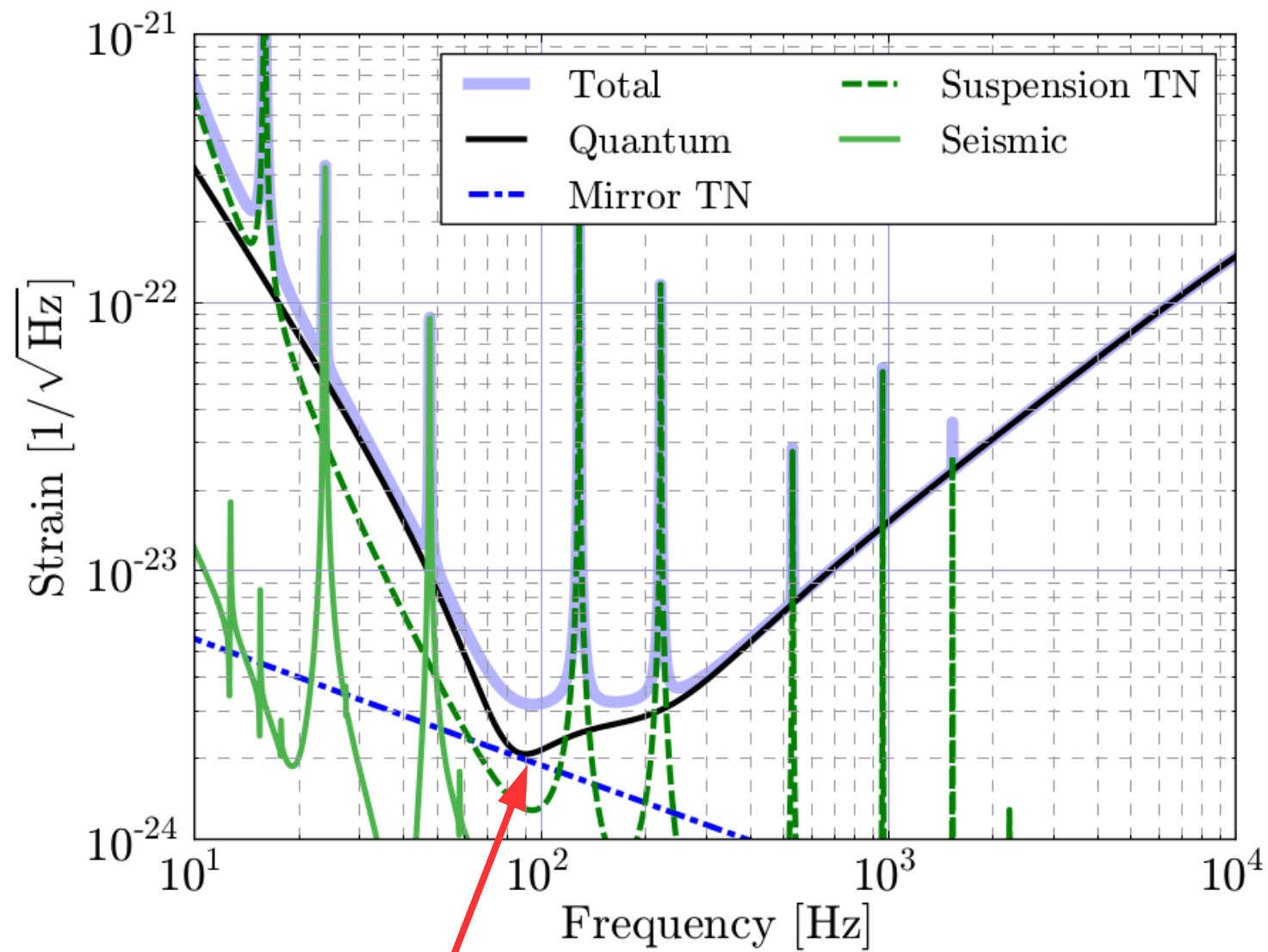
A small sample

- Absorption ranging 30-240ppm
- Strong inhomogeneity

Preparing for large samples



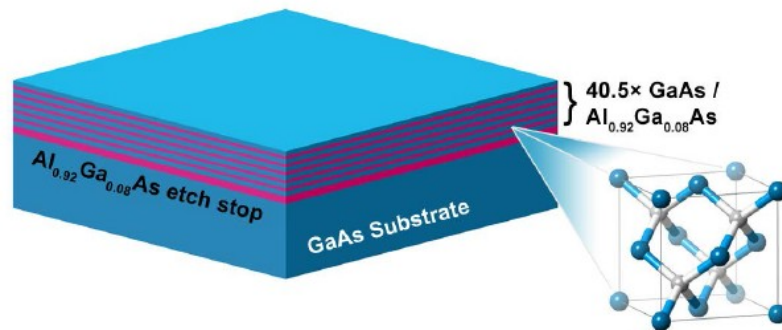
Coating thermal noise matters



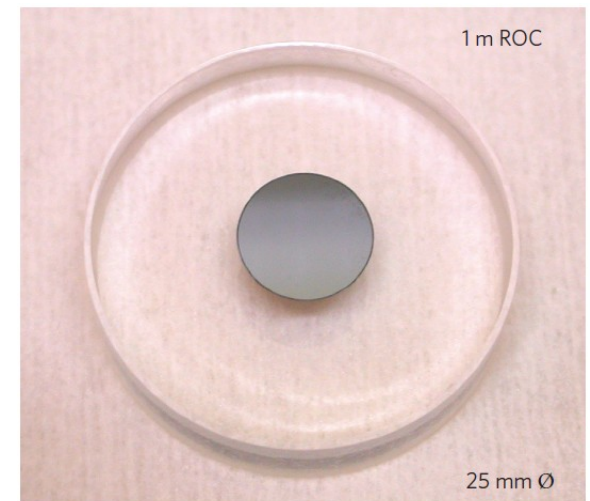
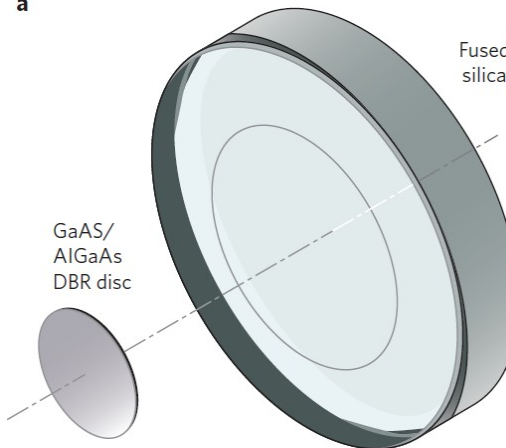
Here

Crystalline Coatings

- Conventional amorphous coating: high mechanical loss
- Low loss coating with high optical performance necessary



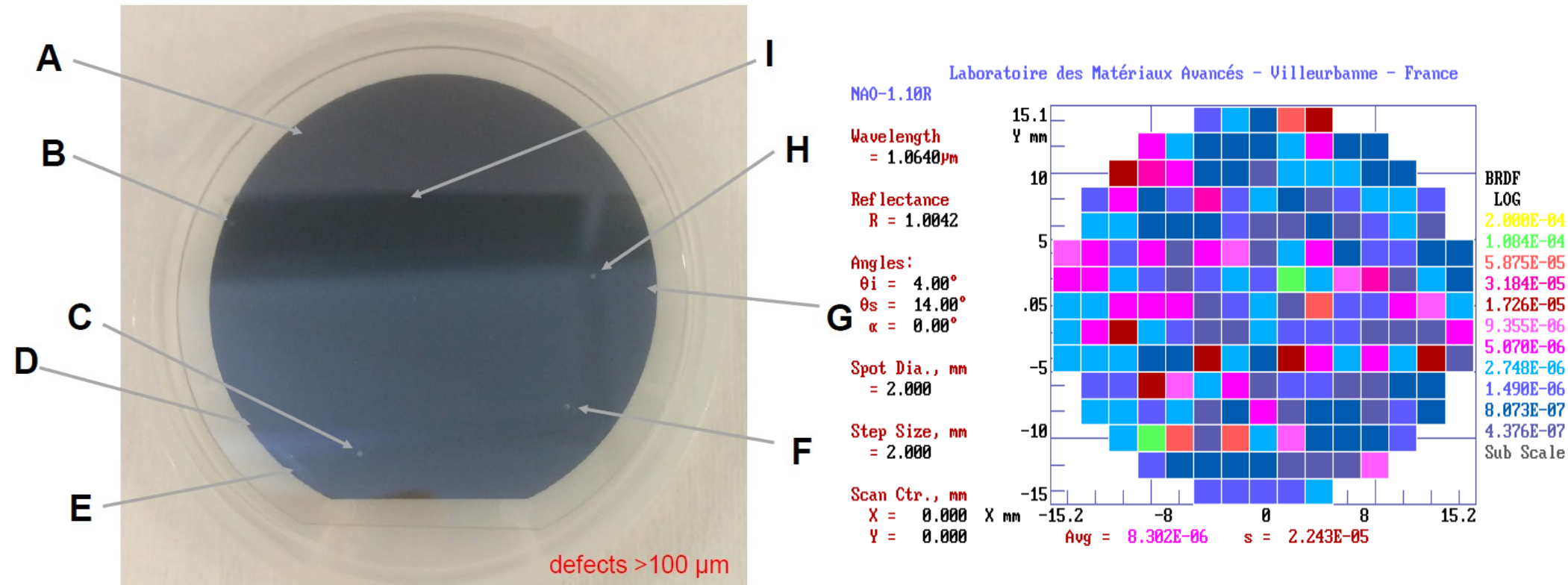
a



G.Cole et al. Nature Photonics 2013

2-inch GaAs/AlGaAs transferred to a sapphire substrate

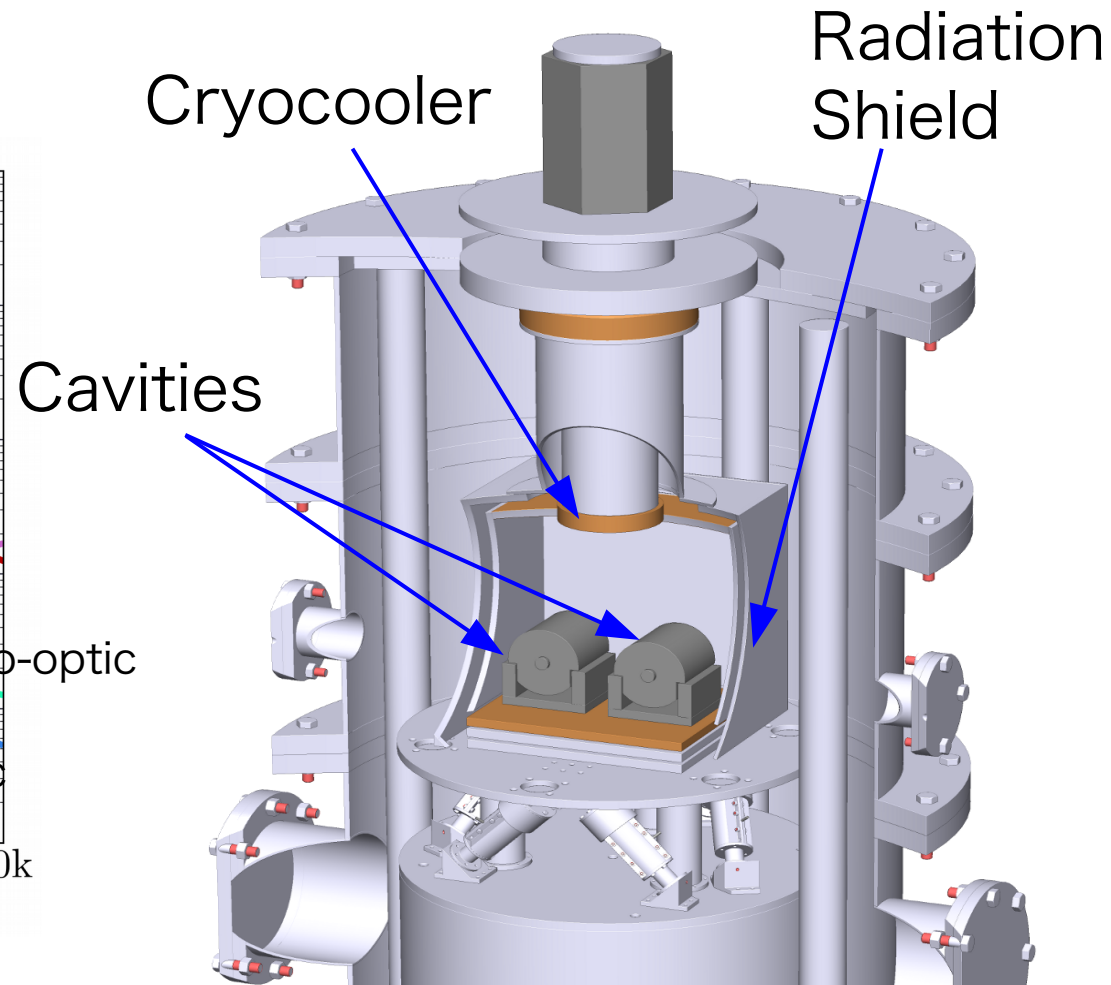
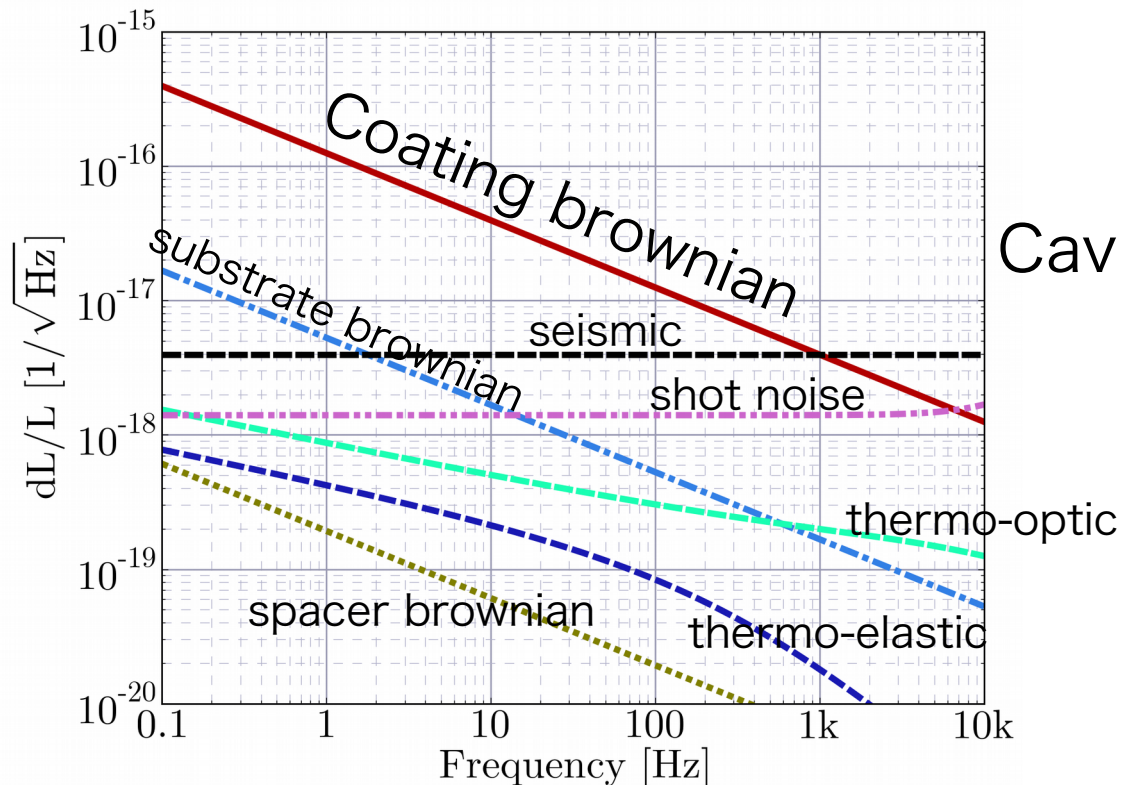
- Largest crystalline coating on Sapphire
- 6ppm scattering
- 9 large defects



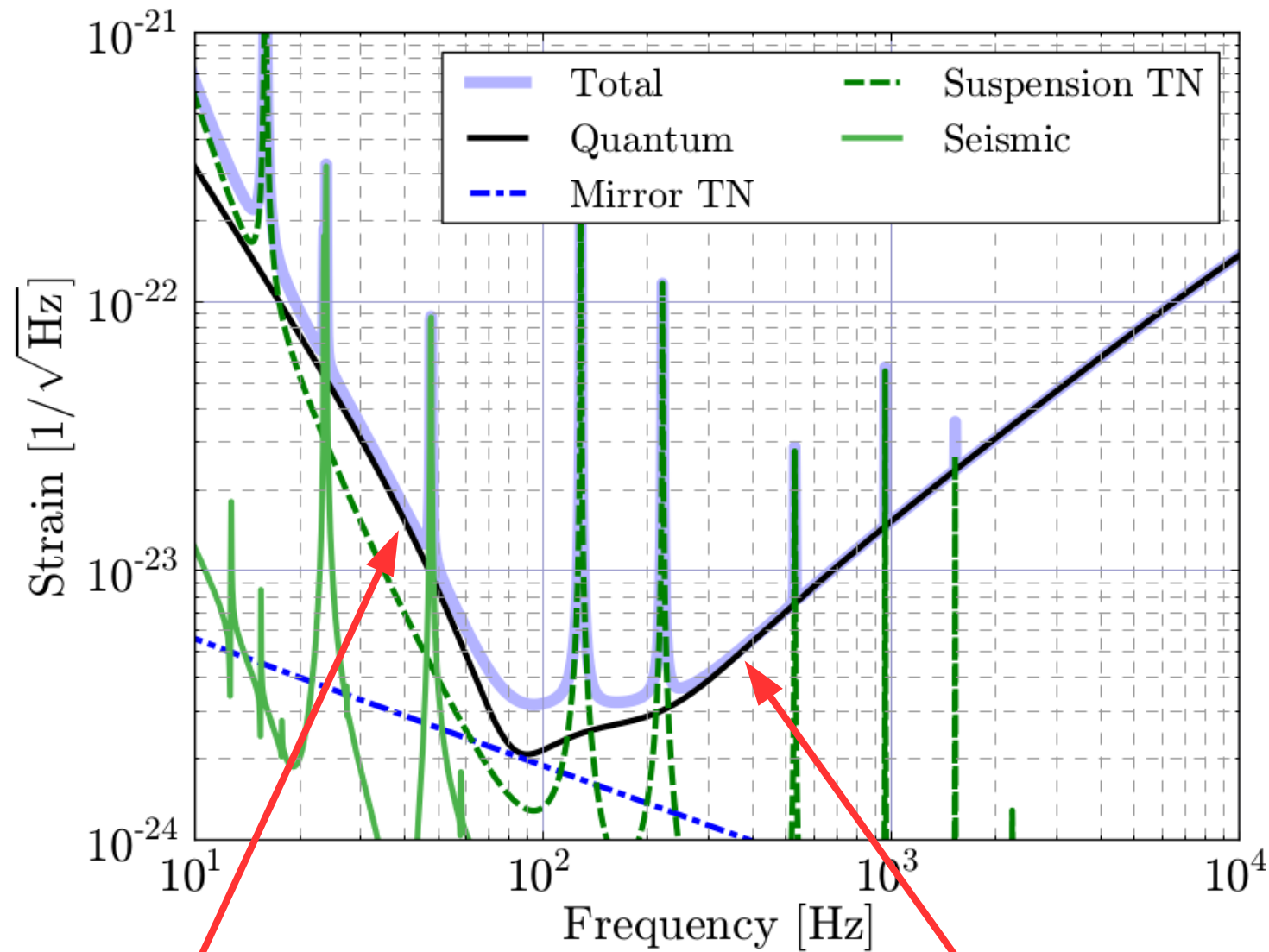
Cryogenic cavities for thermal noise measurement

- Mono-crystalline silicon cavities (1550nm)
- Cooled down to 5K
- Cryogenic operation successful with a large (20cm) cavity
- Working on a twin-cavity configuration

Noise estimates @ 20K



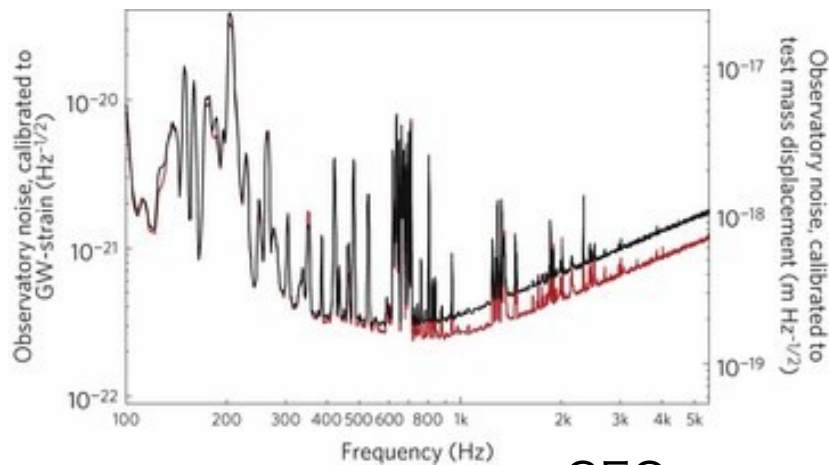
Squeezing more out of our detectors



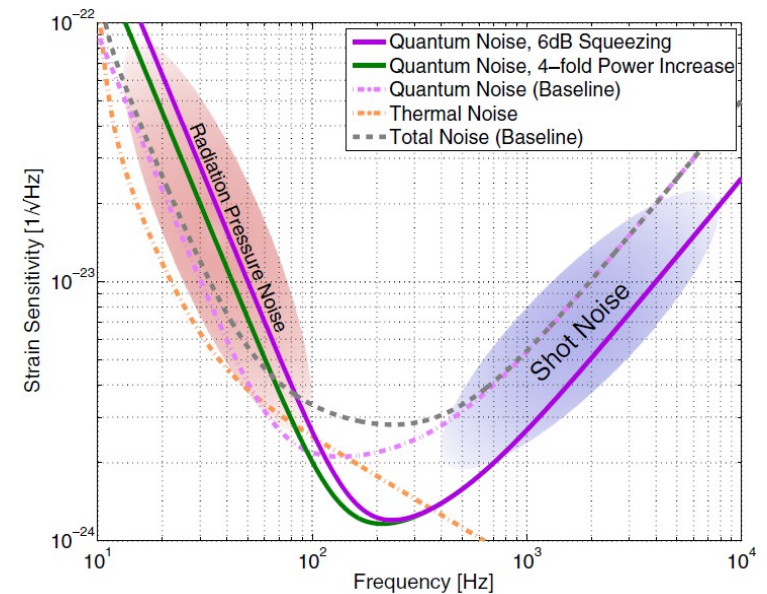
Here

There

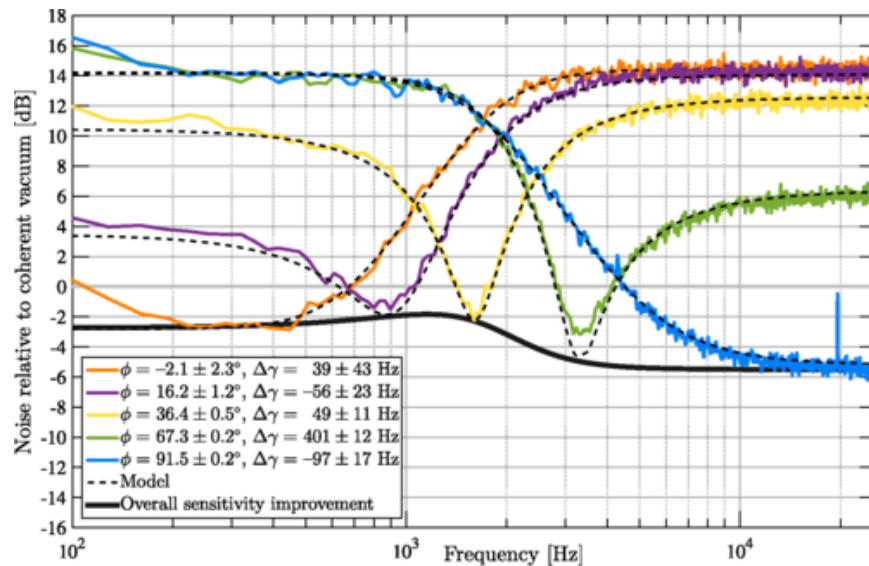
Squeezing: a promising path forward for tackling quantum noises



GEO

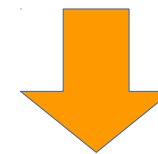


M. Evans et al. 2013



E. Oelker et al. 2016

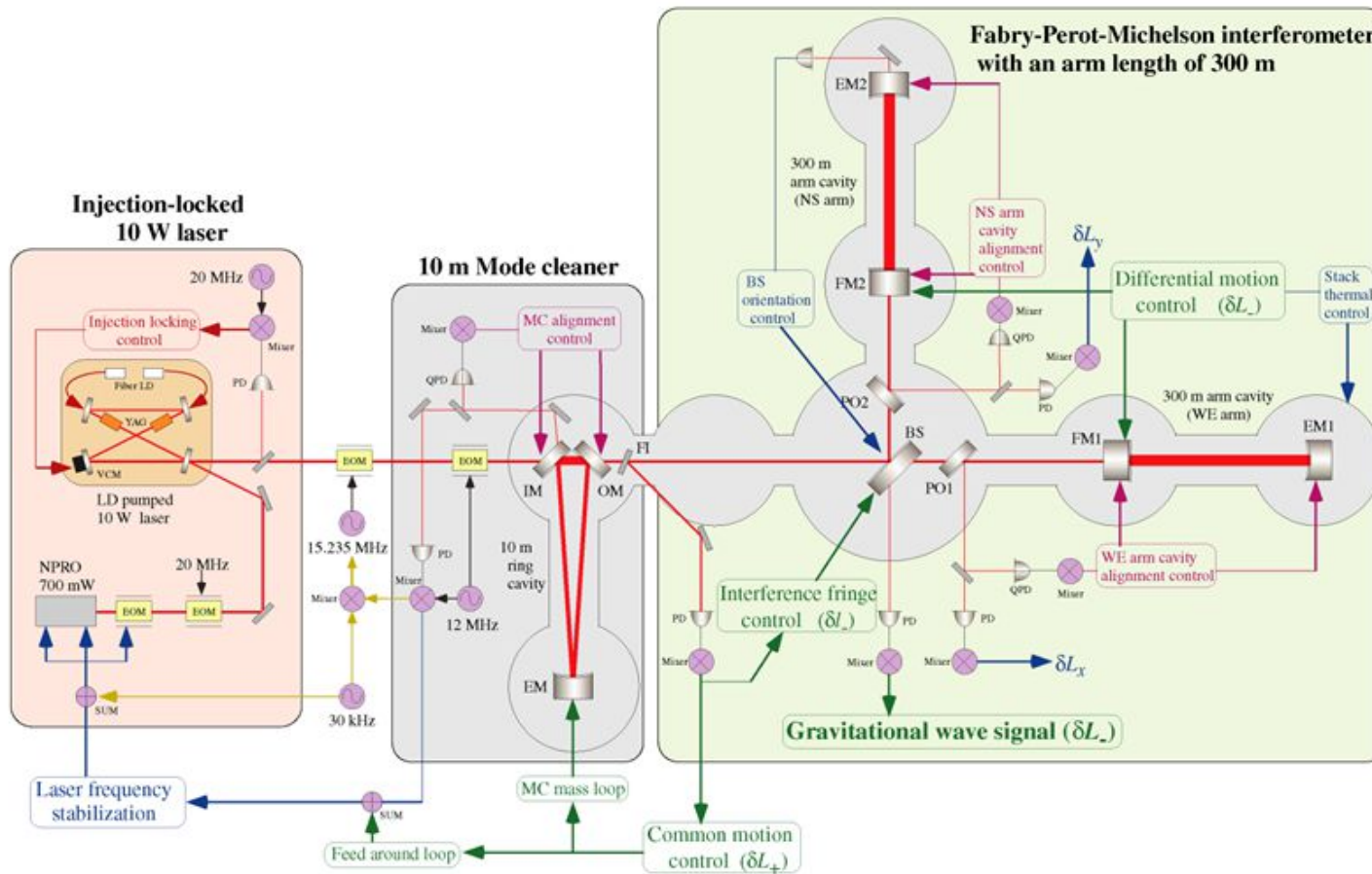
Frequency dependent
rotation of squeezing angle



Necessary for full band sensitivity
improvement

TAMA300

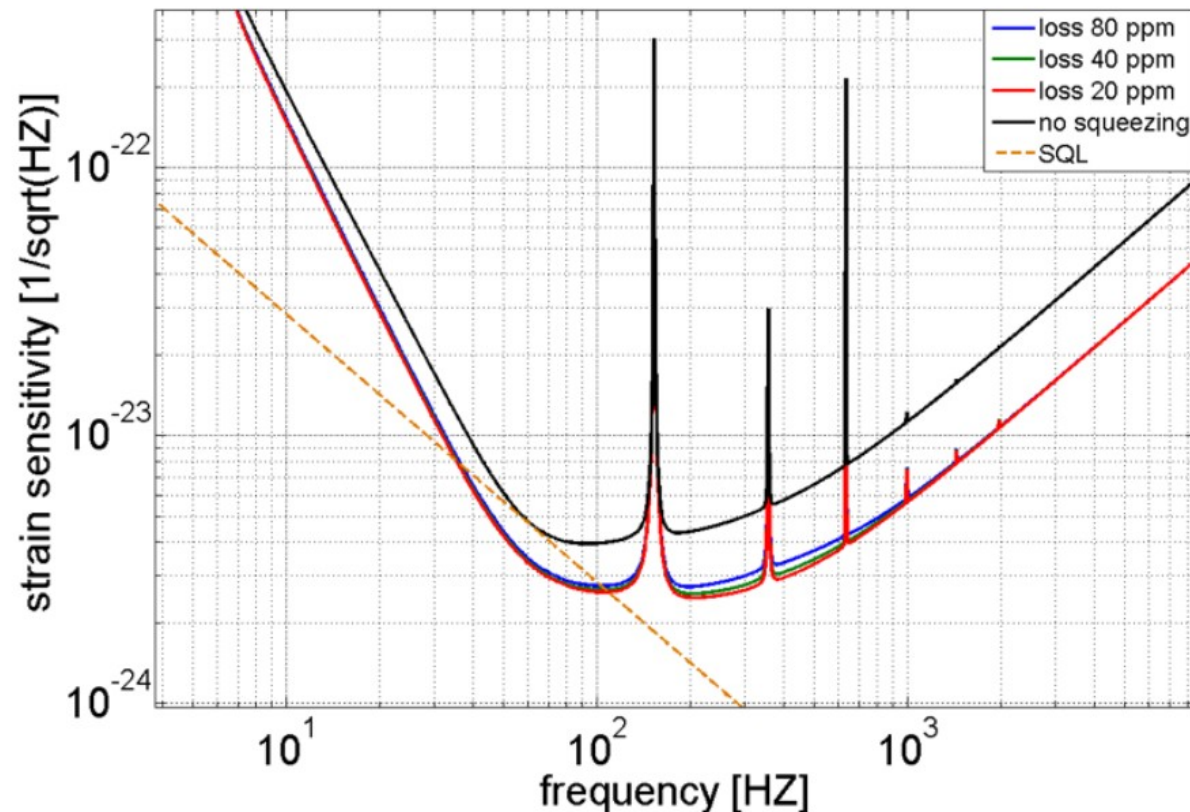
- 300m prototype interferometer
- NAOJ Mitaka Campus



300m Filter Cavity Experiment

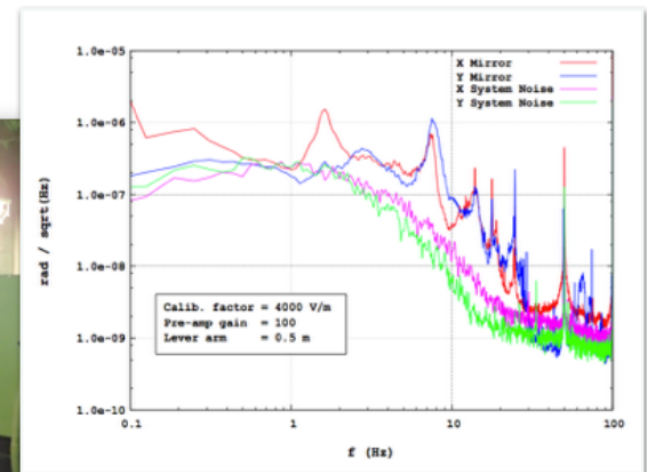
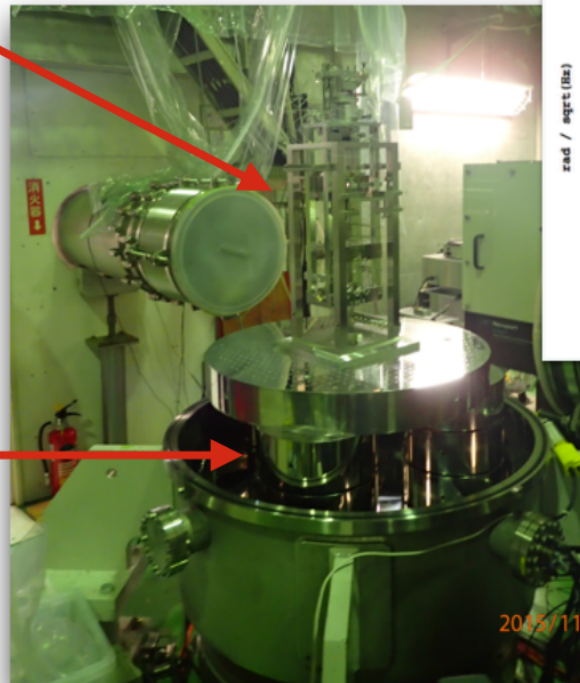
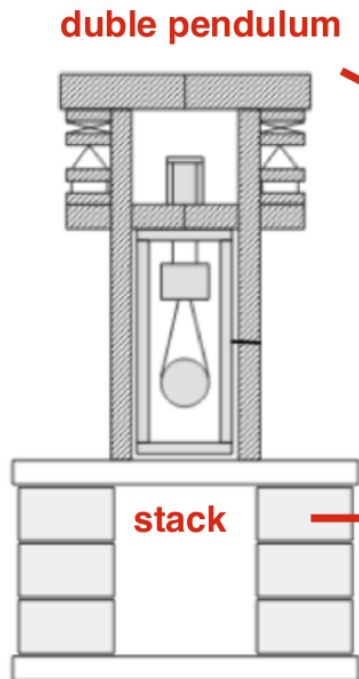
Daisuke Tatsumi, Eleonora Capocasa, Raffaele Flaminio, Matteo Barsuglia, Jérôme Degallaix
Laurent Pinard, Nicolas Straniero, Roman Schnabel, Kentaro Somiya, Yoichi Aso

- 300m long
- 10cm diameter mirrors
- Finesse ~ 4500
- Round trip loss $\sim 80\text{ppm}$



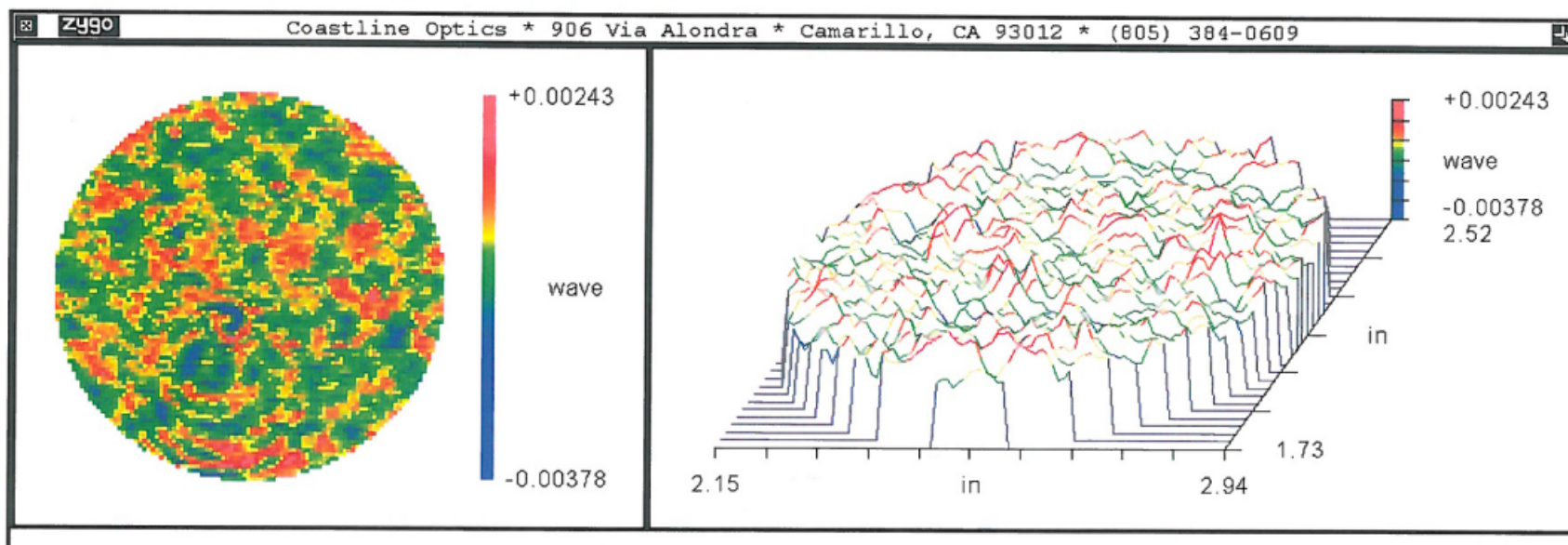
Current Status

Suspensions
being installed



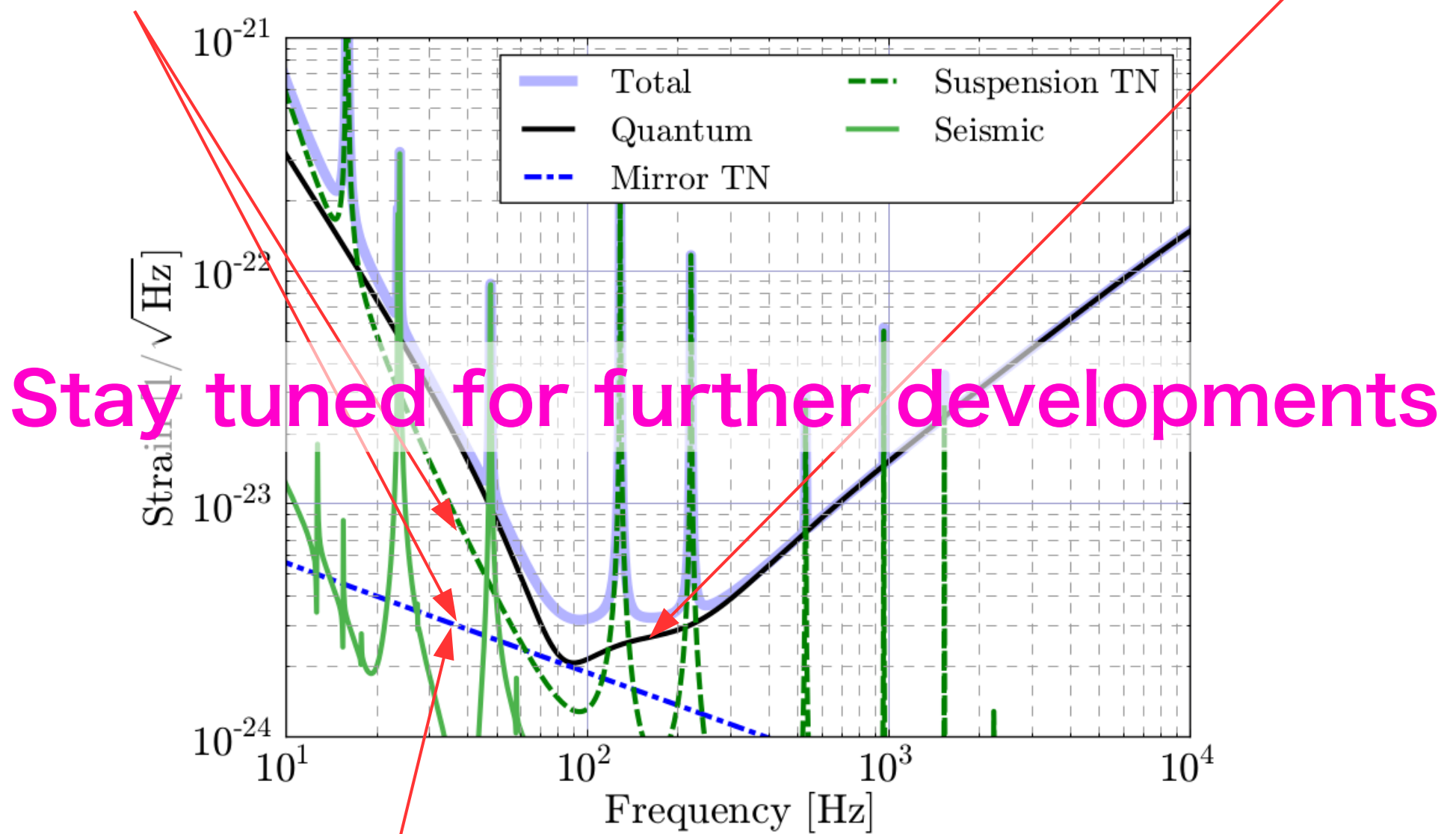
- First measurements of the mirror motion performed

Mirrors
Polished



Cryogenic Suspension
Mirror Absorption Measurements

300m Filter Cavity



Crystalline Coatings
Cryogenic Thermal Noise Measurements