

# The Hartle-Hawking Wave Function in Causal Set Quantum Gravity

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- CST 101
- Continuum Inspired Dynamics for CST : a review of results in 2d CST
- The Hartle-Hawking Wave Function in 2d CST
- The Large N limit.

L. Glaser and S. Surya, Class.Quant.Grav. (2016)

L. Glaser, D. O'Connor and S. Surya, In Preparation

- Spacetime continuum is replaced by a locally finite countable partially ordered set.

Order + Number  $\sim$  Spacetime geometry

- Continuum approximation: via a Poisson Process:

$$P_V(n) \equiv \frac{1}{n!} \exp^{-\rho V} (\rho V)^n, \quad \langle N \rangle = \rho V$$

- Local Lorentz invariance: there are no preferred directions

– L.Bombelli, J.Henson, R. Sorkin, Mod.Phys.Lett. 2009

- Non-locality: A causal set need not be a fixed valency graph.

- ## Order + Number $\sim$ Spacetime geometry

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# A Continuum Inspired Dynamics For CST

- **First principles:** Quantum sequential growth using the *quantum measure formulation*.
- **Continuum Inspired Dynamics:**

$$Z_{\Omega} = \sum_{c \in \Omega} \exp \frac{i}{\hbar} S(c)$$

- $S(C)$  is the *Benincasa-Dowker action* which is the analog of the Einstein-Hilbert action in CST.  
–D. Benincasa and F. Dowker, Phys.Rev.Lett. (2010)  
–F. Dowker and L. Glaser, CQG (2013)
- $\Omega$  is a sample space of causal sets ( e.g.: the set of all past-finite causal sets)

$$Z_{\Omega} = \sum_{c \in \Omega} \exp \frac{i\beta}{\hbar} S(c) \longrightarrow Z_{\Omega} = \sum_{c \in \Omega} \exp -\frac{\beta}{\hbar} S(c)$$

- Space of Configurations  $\Omega$  is unchanged: **There are no Euclidean causal sets!**
- Analytic continuation of parameter:  $i\beta \rightarrow -\beta$
- MCMC for  $\Omega_N$ :
- Covariant/label invariant observables  $\langle O \rangle$

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- MCMC for  $\Omega_N$ :
  - $\beta = 0$ :
    - In the  $N \rightarrow \infty$  limit the *Kleitman-Rothschild* posets dominate.
    - The onset of the asymptotic regime occurs for  $N > 80$
  - Studies have begun on  $\beta \neq 0$  – a challenge!
- Covariant/label invariant observables  $\langle O \rangle$

– J. Henson, D. Rideout, R. Sorkin and S.Surya, 2016

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  - Ordering Fraction:  $R / \binom{N}{2}$
  - Myrheim-Myer Dimension
  - Homology
  - Height
  - Abundance of Causal Diamonds of a given volume



$$Z_{2d} = \sum_{c \in \Omega_{2d}(N)} \exp^{-\frac{1}{\hbar} \beta S_{2d}}$$

–G. Brightwell, J. Henson, S.Surya, Class.Quant.Grav. 25, 2008

–S. Surya, Class.Quant.Grav. 29, 2012

–L. Glaser, D. O'Connor and S. Surya, in preparation

- $\Omega_{2d}(N) \subset \Omega(N)$ :  $N$ -element **2d-orders**:  $U \cap V$ , where  $U$  and  $V$  are total orders:

- 2d Benincasa-Dowker Action:  $\frac{1}{\hbar} S(\epsilon) = 4\epsilon \left( N - 2\epsilon \sum_{n=0}^{N-2} N_n f(n, \epsilon) \right)$

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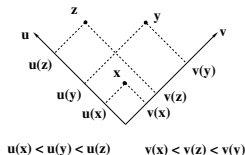
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$$x \prec y \iff u(x) < u(y) \text{ and } v(x) < v(y)$$



- Includes all continuum-like causal sets  $\sim$  topologically trivial causal diamond in 2d.
- Includes posets with no continuum approximation.
- 2d Benincasa-Dowker Action:  $\frac{1}{\hbar} S(\epsilon) = 4\epsilon \left( N - 2\epsilon \sum_{n=0}^{N-2} N_n f(n, \epsilon) \right)$

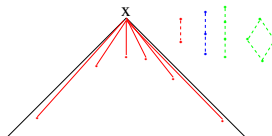
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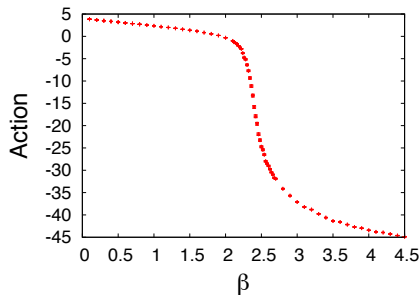
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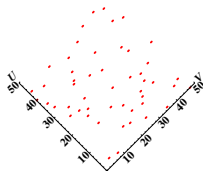
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- 2d Benincasa-Dowker Action:  $\frac{1}{\hbar} S(\epsilon) = 4\epsilon \left( N - 2\epsilon \sum_{n=0}^{N-2} N_n f(n, \epsilon) \right)$ 
  - $N_n$ : # of  $n$ -element order intervals



- Mesoscale:  $\epsilon = \left( \frac{p}{k} \right)^2 \in (0, 1]$
- $f(n, \epsilon) = (1 - \epsilon)^n - 2\epsilon n(1 - \epsilon)^{n-1} + \frac{1}{2}\epsilon^2 n(n-1)(1 - \epsilon)^{n-2}$
- $\epsilon = 1$ :  $\frac{1}{\hbar} S^{(2)}(C) = N - 2N_0 + 4N_1 - 2N_2$

$\langle S \rangle$  vs  $\beta$  for  $N = 50, \epsilon = 0.12$

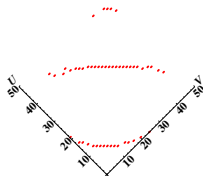




- The typical causal set is a 2d “random order”:  $U$  and  $V$  chosen randomly and independently.
- 2d random orders dominate when  $\beta = 0$ . —Peter Winkler, Order 1, 317, (1985), El-Zahar and N.W. Sauer, Order 5, 239, (1988)
- 2d random order  $\sim$  Minkowski spacetime  $^2M$ . —G. Brightwell, J. Henson, S.Surya, Class.Quant.Grav. 25, 2008

Continuum Phase  $\sim$  2d causal diamond

# Crystalline Phase

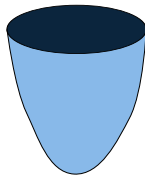


- The typical causal set is layered, with a large number of links, but much fewer small intervals.
- It is distinctly non-manifold like.

Crystalline Phase  $\not\sim$  continuum

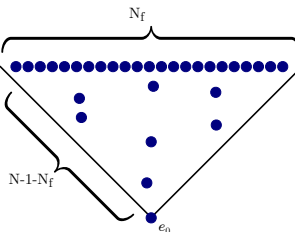
# The Hartle-Hawking Prescription in CST

- Continuum Proposal:  $\Psi_0(h_{ab}, \Sigma) = A \sum_M \int dg \exp^{-\frac{1}{\hbar} I_E(g)}, \quad \partial M = \Sigma, g|_{\Sigma} = h$ 
  - Path integral over Riemannian geometries on  $M$ .
  - $M$  is compact with a “final” boundary geometry  $(\Sigma, h_{ab})$ .
  - Initial spatial “zero” geometry, *“a single point, which captures the idea of a universe emerging from nothing.”*



# The Hartle-Hawking Prescription in CST

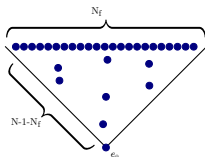
- Continuum Proposal:  $\Psi_0(h_{ab}, \Sigma) = A \sum_M \int dg \exp^{-\frac{1}{\hbar} I_E(g)}, \quad \partial M = \Sigma, g|_{\Sigma} = h$
- CST Proposal:  $\Psi_0^{(N)}(\mathcal{N}_f, \beta) \equiv A \sum_{c \in \Omega_N} \exp^{-\frac{1}{\hbar} \beta S(c)}, \quad |\text{Max}(C)| = \mathcal{N}_f$ 
  - The sum is over “discrete Lorentzian” geometries or causal sets,  $c \in \Omega_N$  of finite cardinality.
  - Initial spatial geometry is a single element to the past of all other elements in  $c$ . This is “a single point” from which the universe emerges.
  - The final geometry is a *maximal antichain*  $\mathcal{A}_f$  whose only characteristic feature is cardinality which is therefore fixed  $|\mathcal{A}_f| = \mathcal{N}_f$ .





# The 2d Hartle-Hawking Wave Function: Analytic Results

$$\psi_0^{(N)}(\mathcal{N}_f, \beta) \equiv A \sum_{c \in \Omega_{2d}} \exp^{-\frac{1}{\hbar} \beta S_{2d}(c)}, \quad \mathcal{N}_f = N - p$$



- $\mathbf{p} = \mathbf{1}$  :  $\psi_0(N-1) = A \exp^{-\beta R}$ ,  $R = 2\epsilon N(1-2\epsilon) + 4\epsilon^2$

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- $\mathbf{p} = \mathbf{2}$  :  $\psi_0(N-2) = \frac{A \exp^{-\beta(R-Q)}}{(1 - \exp^{\beta Q})^2} \left( N - 2 - (N-1) \exp^{\beta Q} + \exp^{\beta Q(N-1)} \right), Q = 4\epsilon^2(1 - 3\epsilon).$

# The 2d Hartle-Hawking Wave Function: Analytic Results

$$\psi_0^{(N)}(\mathcal{N}_f, \beta) \equiv A \sum_{c \in \Omega_{2d}} \exp^{-\frac{1}{\hbar} \beta S_{2d}(c)}, \quad \mathcal{N}_f = N - p$$

- **p = 3 :**

- antichain

$$\psi_0^{(a,i)}(N-3) = A \exp^{-\beta R} \sum_{\ell_1=1}^{N-3} \sum_{\ell_2=1}^{N-3} \sum_{m=m_0}^{m_f} (N-2-\ell_1-\ell_2+m) \exp^{\beta P m} \exp^{\beta Q(\ell_1+\ell_2)}$$

- chain

$$\psi_0^{(a,ii)}(N-3) = A \exp^{-\beta R} \sum_{\ell_1=1}^{(N-3-1)} \sum_{\ell_2=1}^{(N-3-\ell_1)} \sum_{\tilde{m}=0}^{(N-3-\ell_1-\ell_2)} (N-2-\ell_1-\ell_2-\tilde{m}) \exp^{\beta Q(\ell_1+\ell_2)}$$

$$P = 24\epsilon^4, \quad m_0 = \max(1, \ell_1 + \ell_2 - N + 3), \quad m_f = \min(\ell_1, \ell_2)$$

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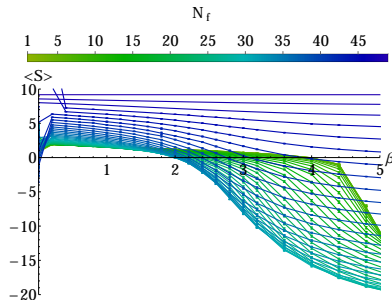
- $p > 3$  : Analytically challenging/impossible!

$$\Psi_0(\mathcal{N}_f) = A\mathcal{Z}_\beta(\mathcal{N}_f) = A\mathcal{Z}_0(\mathcal{N}_f) \exp\left(-\int_0^\beta d\beta' \langle S_{\beta'}(\mathcal{N}_f) \rangle\right)$$

- Calculation of  $\langle S_\beta(\mathcal{N}_f) \rangle$  using MCMC methods.
- Numerical Integration:  $\int_0^\beta d\beta' \langle S_{\beta'}(\mathcal{N}_f) \rangle$
- Estimation of  $\mathcal{Z}_0(\mathcal{N}_f)$ .
- Normalise to get  $A$ .
- Calculations performed for  $N = 50, \epsilon = 0.12, 0.5, 1$ .

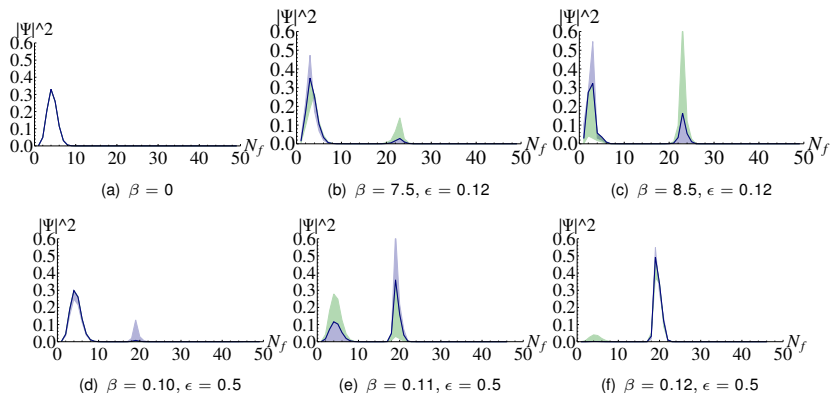
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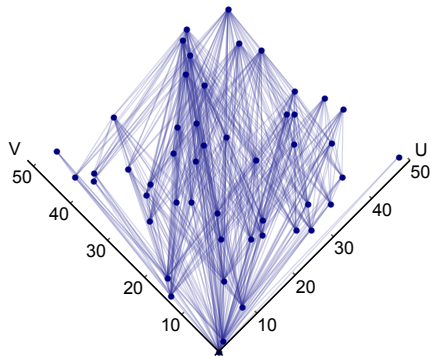
- Phase transition for smaller  $\mathcal{N}_f$ .
- As  $\mathcal{N}_f$  increases to  $N - 1$ , the phase transition is wiped out
- Minimum value of  $\beta_c$  at  $\mathcal{N}_f \sim 30$  :  $\beta_c(\mathcal{N}_f)$  is not a monotonic function.
- Numerical Integration:  $\int_0^\beta d\beta' \langle S_{\beta'}(\mathcal{N}_f) \rangle$

# The Hartle-Hawking Wave Function.



# The Two Peaks

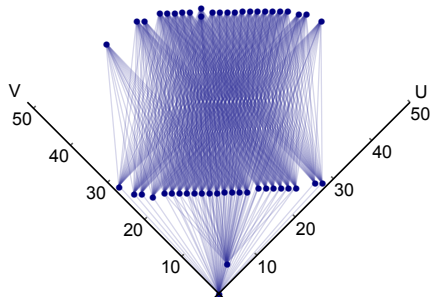
- The Peak at  $\mathcal{N}_f \sim 4$





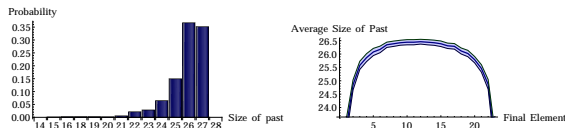
# The Two Peaks

- The Peak at  $\mathcal{N}_f \sim 23$



# Features of the second peak geometry

- Rapid expansion from a single element to a large spatial slice:  $\mathcal{N}_f/\text{height} \sim 6$ .
- Homogeneity determined from causal past.

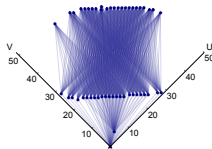


- Causal pasts of final elements maximally overlap
- Non-manifold like.

Initial Conditions for the Universe from Quantum Gravity?

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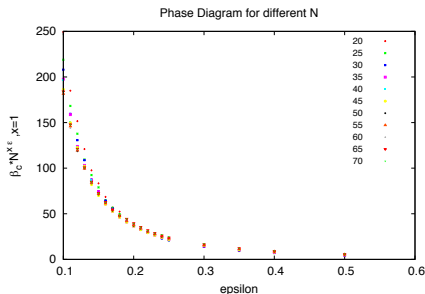
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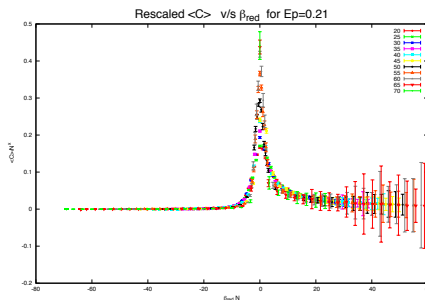
- Nature of phase transition: First order for  $\epsilon = 0.02 - -0.5$ .



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Thermodynamic Limit Exists

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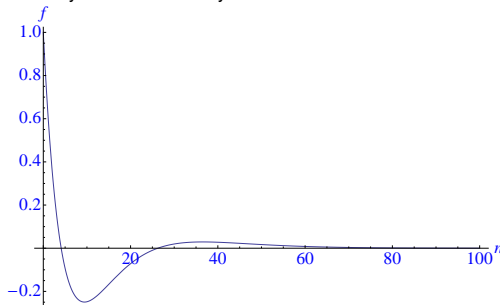
# Summary and Open Questions and Speculations

- A concrete *illustration* of how physically interesting initial conditions can arise from a theory of quantum gravity.
- Importance of non-continuum structures.
- 2d is NOT 4d, but there are universal dimension-independent features:
  - $\beta$  parameter: In  $d > 2$  rescales the Planck volume.
  - Large  $\beta$  limit is dominated by the Action  $\Rightarrow$  Crystalline Phase will dominate.
  - Speculations:

*Simulations on RRI HPC cluster. Supported in part by FQXi via Theiss Research*

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  - Speculations:  $N \sim$  Age of Universe.  $\epsilon$  could flow to smaller values for which spacetime emerges.

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