

# ELECTRODYNAMIC EFFECTS OF INFLATIONARY GRAVITONS

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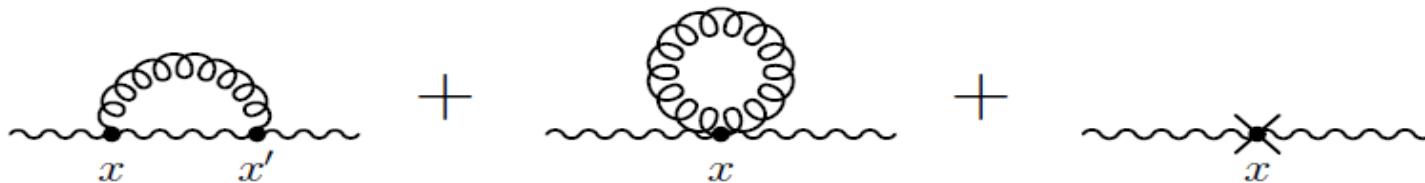
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# What We Did

- $\mathcal{L} = \frac{(R-2\Lambda)\sqrt{-g}}{16\pi G} - \frac{1}{4}F_{\mu\nu}F_{\rho\sigma}g^{\mu\rho}g^{\nu\rho}\sqrt{-g} + \text{c-terms}$
- 1. Compute  $i[\mu\Pi^\nu](x; x')$  from gravitons at 1 loop in dim. reg.



- 2. Renormalize with BPHZ counter terms (cf. LEFT)

$$\begin{aligned}\Delta\mathcal{L} = & C_4 D_\alpha F_{\mu\nu} D_\beta F_{\rho\sigma} g^{\alpha\beta} g^{\mu\rho} g^{\nu\sigma} \sqrt{-g} + \bar{C} H^2 F^{\mu\nu} F_{\rho\sigma} g^{\mu\rho} g^{\nu\sigma} \sqrt{-g} \\ & + \Delta C H^2 F_{ij} F_{kl} g^{ik} g^{jl} \sqrt{-g}\end{aligned}$$

- 3. Quantum-correct Maxwell's equations

$$\partial_\nu [\sqrt{-g} g^{\nu\rho} g^{\mu\sigma} F_{\rho\sigma}(x)] + \int d^4 x' [\mu\Pi^\nu](x; x') A_\nu(x') = J^\mu(x)$$

- $J^\mu = 0 \rightarrow$  dynamical photons
- $J^\mu \neq 0 \rightarrow$  electromagnetic forces

# Why there should be an effect

- Neither photons nor gravitons are charged
  - So how can they polarize the vacuum?
- But they DO carry momentum
  - Dynamical photons have  $\vec{p} = \hbar c \vec{k}$
  - Virtual photons carry  $\vec{F} = \frac{\Delta \vec{p}}{\Delta t}$
- Virtual gravitons can add or subtract to this
  - de Sitter gravitons  $(\vec{k}, \lambda)$  have  $N(t, k) = \left[ \frac{Ha(t)}{2ck} \right]^2$
  - Gravitons with  $k_{phys} \sim H \gg k_{photon} = \frac{k}{a(t)}$

# What happens on a flat background

arXiv:1202.5800 by Leonard & Woodard

- $\kappa^2 \equiv 16\pi G$        $\Delta x^2 \equiv \|\vec{x} - \vec{x}'\|^2 - (|t - t'| - i\epsilon)^2$
- $i[\mu \Pi^\nu](x; x') = (\eta^{\mu\nu} \partial' \cdot \partial - \partial'^\mu \partial^\nu) \left[ \frac{\# \kappa^2}{\Delta x^{2D-2}} \right]$
- $\Delta \mathcal{L} = C_4 D_\alpha F_{\mu\nu} D_\beta F_{\rho\sigma} g^{\alpha\beta} g^{\mu\rho} g^{\nu\sigma} \sqrt{-g}$
- $i[\mu \Pi_{ren}^\nu](x; x') = (\eta^{\mu\nu} \partial' \cdot \partial - \partial'^\mu \partial^\nu) \partial^4 \left[ \frac{\kappa^2}{192\pi^2} \frac{\ln(\mu^2 \Delta x^2)}{\Delta x^2} \right]$
- No change in dynamical photons
- Coulomb:  $\Phi(r) = \frac{q}{4\pi r} \left[ 1 + \frac{2G}{3\pi r^2} + O(G^2) \right]$ 
  - Radkowski (1970)
  - 0<sup>th</sup> order field distorts virtual gravitons nearby & these add momentum to virtual photons carrying force

# What happens on de Sitter

arXiv:1304.7265 by Leonard & Woodard ; arXiv:1308.3453 by Glavan, Miao, Prokopec & Woodard

- de Sitter breaking  $\rightarrow$  2 structure functions
- $i[\mu \Pi^\nu] = (\eta^{\mu\nu}\eta^{\rho\sigma} - \eta^{\mu\sigma}\eta^{\nu\rho})\partial_\rho\partial'_\sigma F(x; x') + (\bar{\eta}^{\mu\nu}\bar{\eta}^{\rho\sigma} - \bar{\eta}^{\mu\sigma}\bar{\eta}^{\nu\rho})\partial_\rho\partial'_\sigma G(x; x')$
- Dynamical photons: E grows and B falls
  - $F_{(1)}^{0i} \rightarrow F_{(0)}^{0i} \times \frac{\kappa^2 H^2 \ln(a)}{8\pi^2}$  ,  $F_{(1)}^{ij} \rightarrow F_{(0)}^{ij} \times \frac{\kappa^2 H^2}{8\pi^2} \frac{ik \ln(a)}{Ha}$
- Co-moving charge & magnetic dipole
  - $J^\mu(\eta, \vec{x}) = q\delta_0^\mu\delta^3(\vec{x}) \rightarrow \Phi = \frac{q}{4\pi x} \left\{ 1 + \frac{\kappa^2 H^2}{8\pi^2} \left[ \frac{1}{3a^2 H^2 x^2} + \ln(aHx) \right] + \dots \right\}$
  - $J^\mu(\eta, \vec{x}) = \epsilon^{0\mu jk}m_j\partial_k\delta^3(\vec{x}) \rightarrow \vec{A} = \vec{m} \times \vec{\nabla} \left( \frac{1}{4\pi x} \left\{ 1 + \frac{\kappa^2 H^2}{8\pi^2} \left[ \frac{1}{3a^2 H^2 x^2} - \frac{2}{3} \ln(Hx) \right] + \dots \right\} \right)$
- Static charge & magnetic dipole
  - Point Charge  $\rightarrow \tilde{\Phi} = \frac{q}{4\pi r} \left\{ 1 + \frac{\kappa^2 H^2}{8\pi^2} \left[ \frac{1}{3H^2 r^2} + \ln(Hr) \right] + \dots \right\}$
  - Point magnetic dipole  $\rightarrow$  show secular growth

# The Gauge Issue: Photon & gravitons require fixing!

- EM gauge irrelevant because coupling to  $F_{\mu\nu}$
- But flat  $\partial^\mu h_{\mu\nu} = (\frac{b}{2})\partial_\nu h$  in flat space gives

$$i[\mu\Pi^\nu](x; x') = \frac{\kappa^2}{384\pi^4} \left(\frac{2b-1}{b-2}\right)^2 [\eta^{\mu\nu} \partial' \cdot \partial - \partial'^\mu \partial^\nu] \partial^2 \left[ \frac{\ln(\mu^2 \Delta x^2)}{\Delta x^2} \right]$$

- Spin 2 part vanishes at  $D = 4$
- The result of Bjerrum-Bohr (hep-th/0206236)
  - Effect in gauge independent S-matrix

# Our Conjecture:

arXiv:1204.1784 by Miao & Woodard

- Gauge fixed GF's **are not complete nonsense**
  - We construct the flat space S-matrix from them!
  - Need to separate physical information from unphysical
- Gauge invariant GF's **cannot solve everything**
  - Every gauge fixed GF represents some invariant
$$A_0(t, x) = 0 = A_1(0, x) \rightarrow A_1(t, x) = \int_0^t ds F_{01}(s, x)$$
- $i[\mu \Pi^\nu]$  on de Sitter must inherit the gauge dependence of  $i[\mu \Pi^\nu]$  on flat background
- **Maybe leading secular effects gauge independent**
  - NB these effects aren't present in flat space
  - Cf. poles terms of gauge fixed GF's in flat space QFT
- How to check?  $\rightarrow$  re-compute in other gauges
  - **EM gauge doesn't matter**
  - General covariant GR gauge:  $D^\mu h_{\mu\nu} = \frac{b}{2} \partial_\nu h_\mu^\mu$
  - arXiv:1205.4468 by Mora, Tsamis & Woodard

# Conclusions

- Dynamical photons:
- Spin 2 shows secular growth & no gauge dependence
- Spin 0 doesn't have leading order contribution
  - Consistent with our conjecture
- Intuition says the effect is real
  - Inflation really does produce huge ensemble of IR
  - Should scatter photons more the further they go
  - This would not even be doubted in flat space
- But proving it IS an important step in cosmo QFT