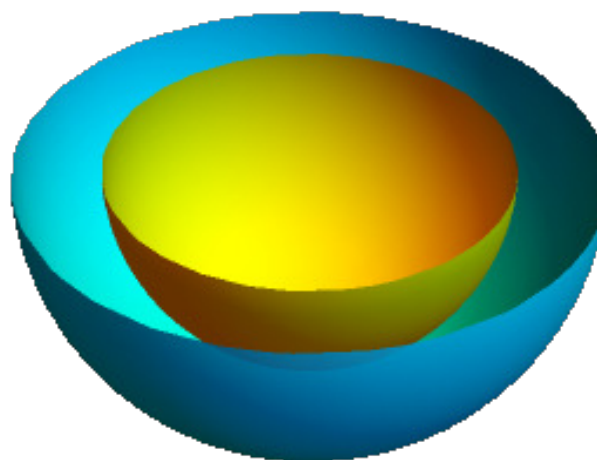


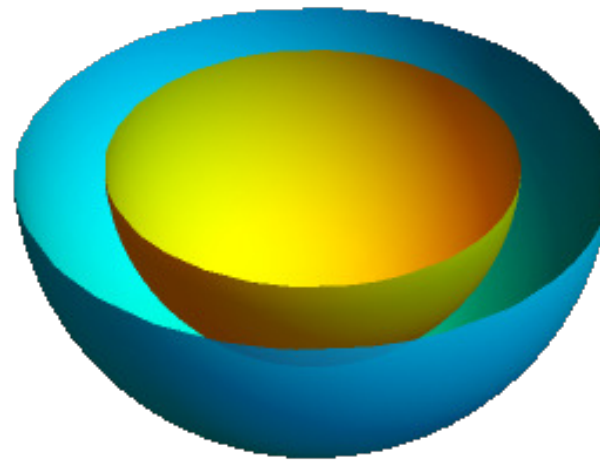
Dynamics of confined double-shells systems, critical behavior and chaos

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based on:

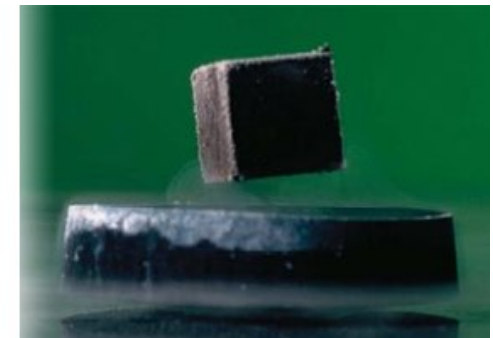
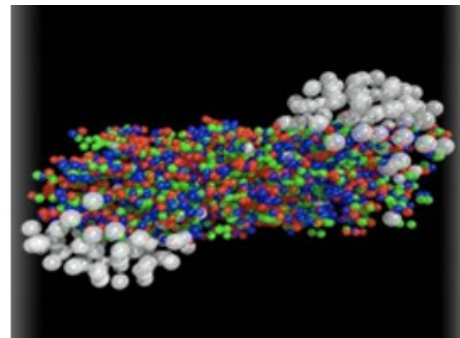
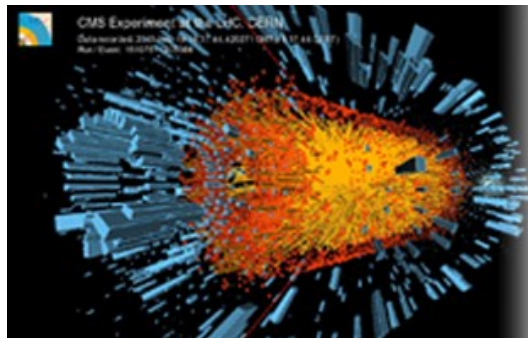
- ❖ 1601.07552 [gr-qc] w/ **Vitor Cardoso**
- ❖ 1602.03535 [hep-th] w/ **Richard Brito and Vitor Cardoso**

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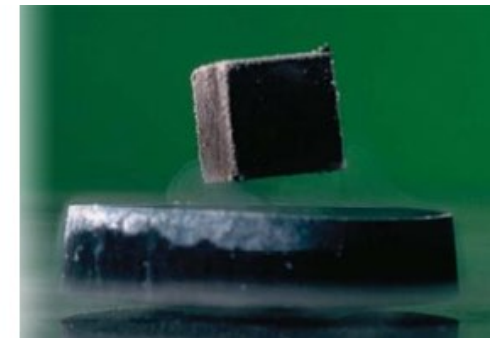
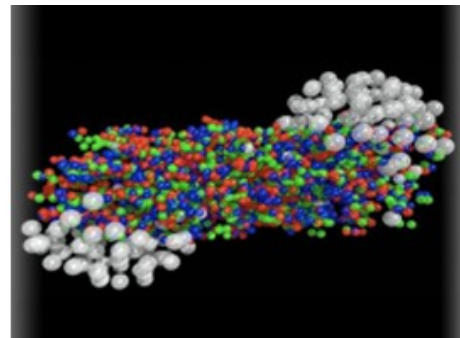
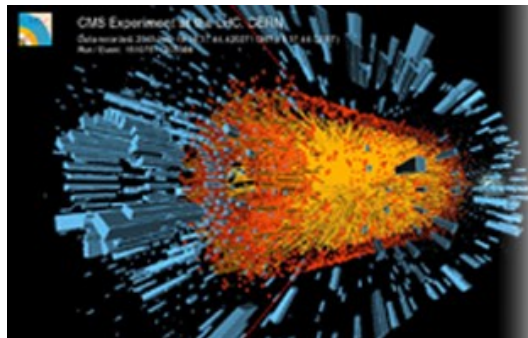
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- ✦ But the subject of **gravity in confined spaces** has intrinsic interest. Even in the context of **four-dimensional classical gravity**.

Introduction: Gravity in confined spaces

- ✦ More motivation: the **turbulent instability of AdS**.

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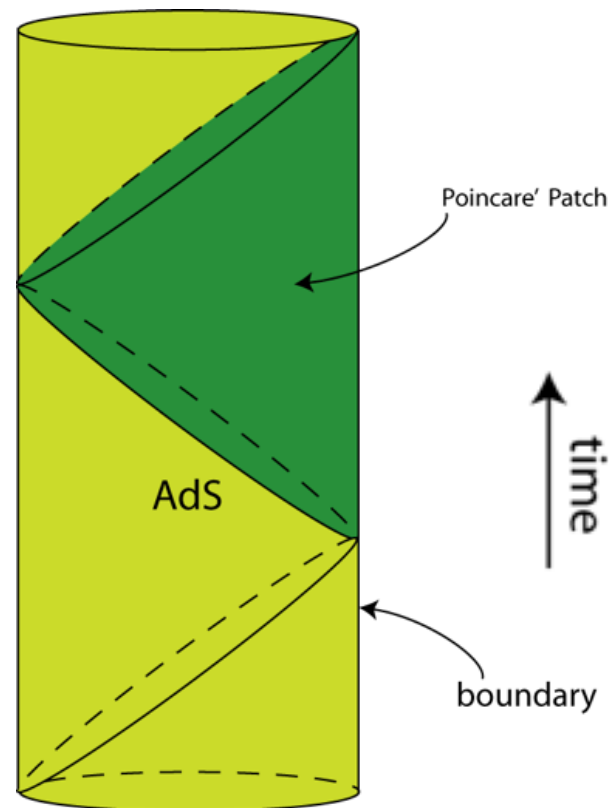
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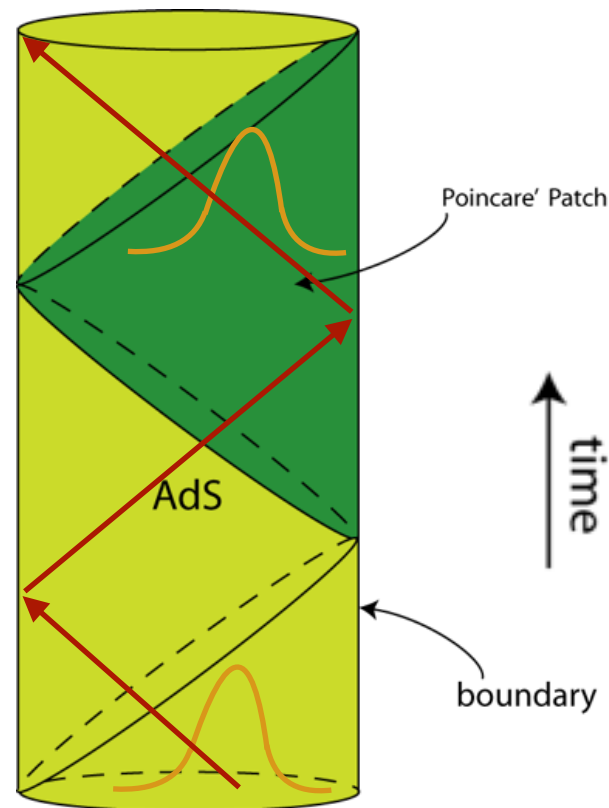
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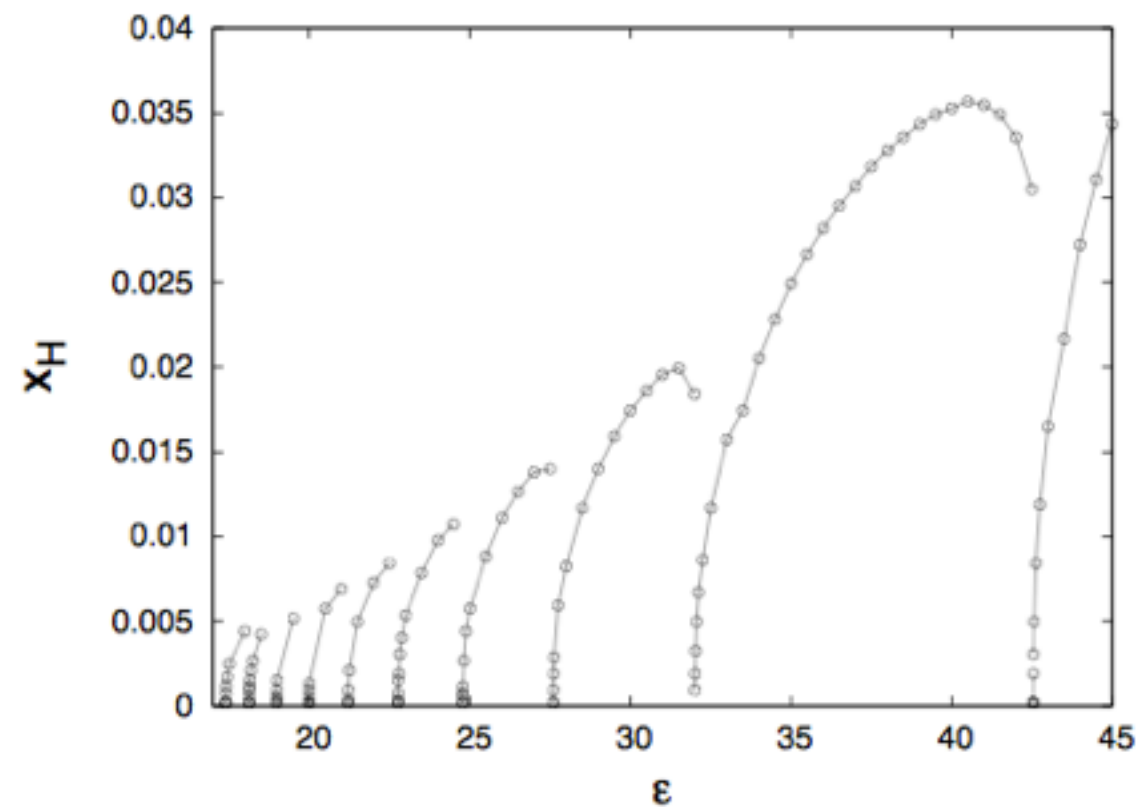
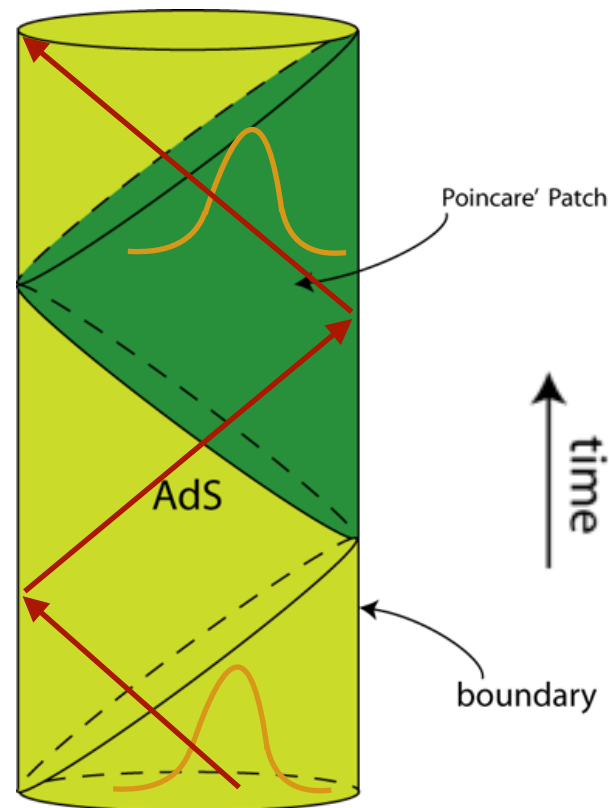
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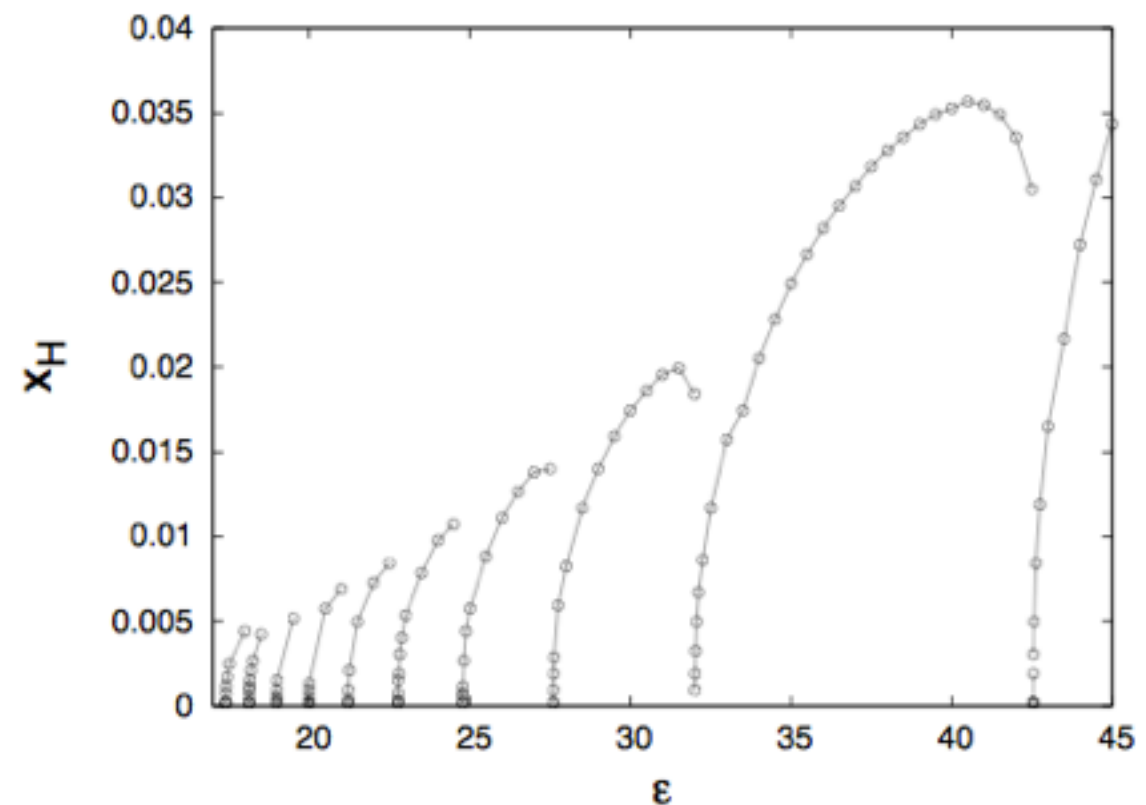
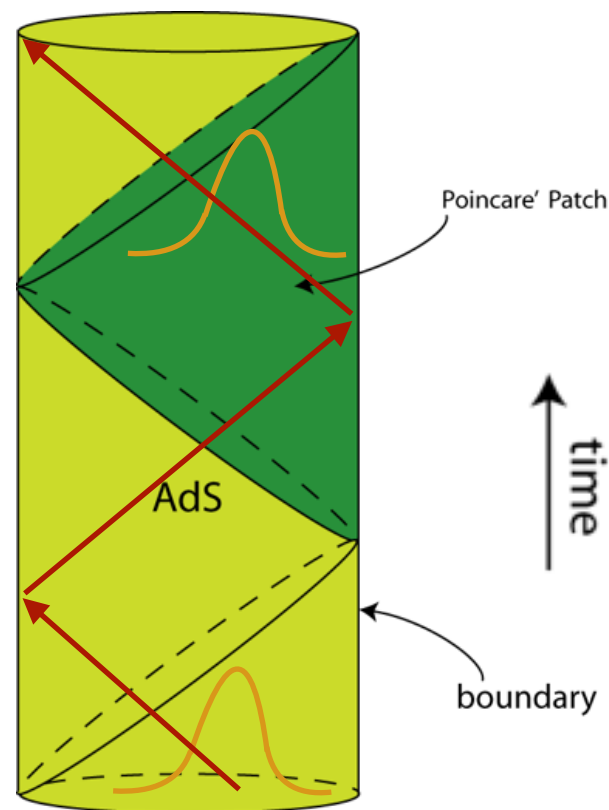
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- ♦ **Arbitrarily (?)** small perturbations of AdS typically (?) **collapse** after a sufficiently large number of **reflections**.

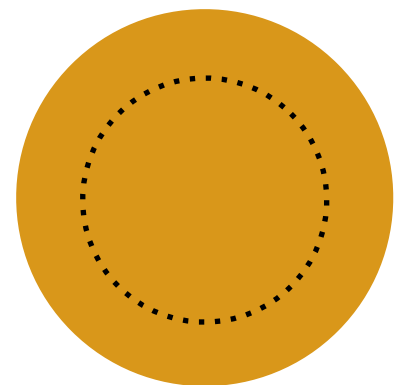
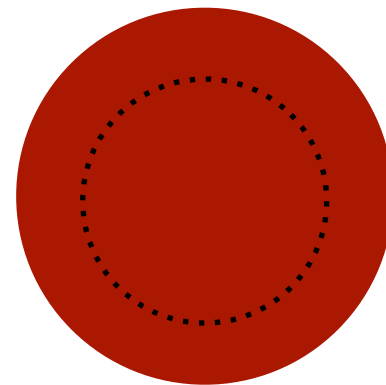
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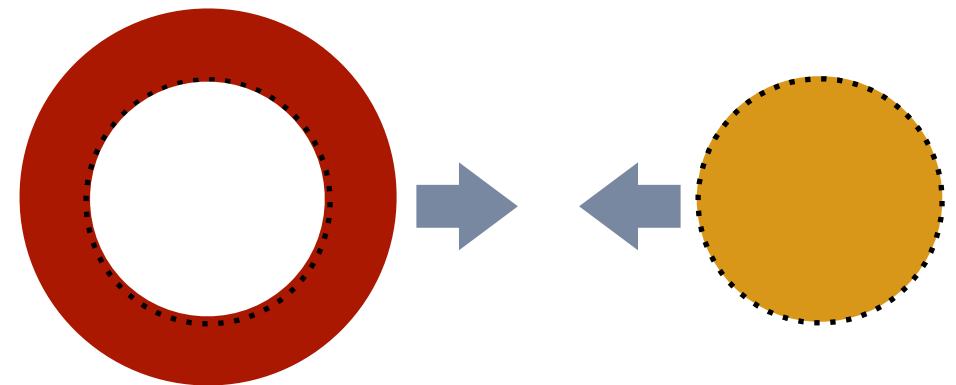
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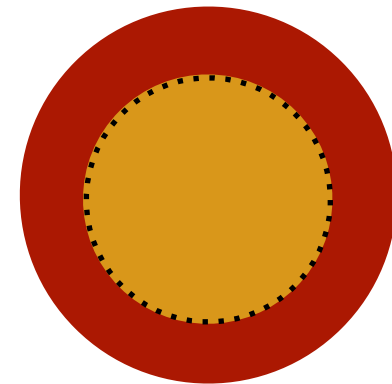
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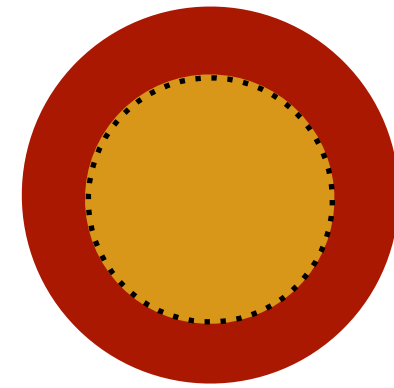
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- ✦ Note: the thin shell may be *dynamical* even if the two spacetimes matched are **static**.

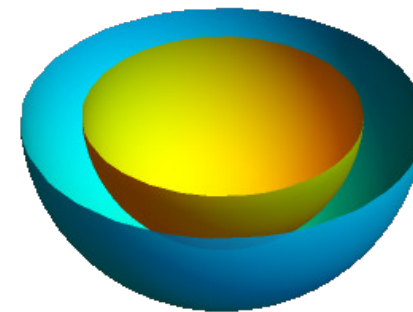


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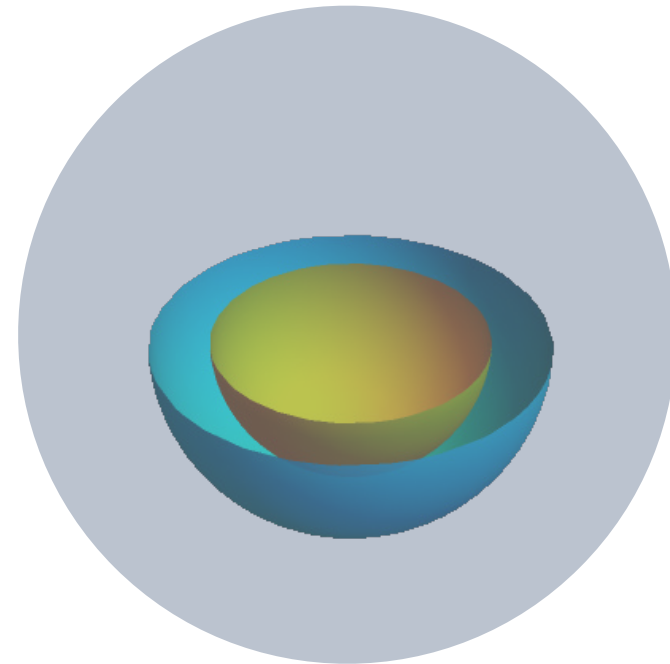
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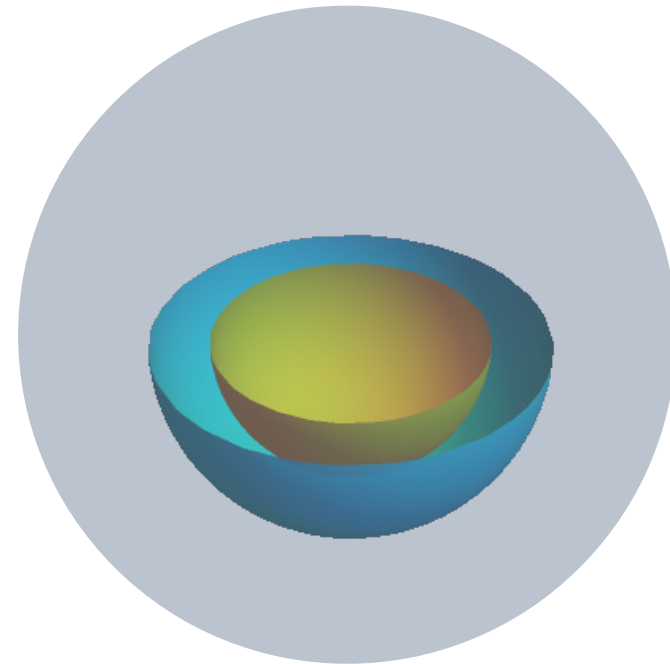
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This basic setting displays ***astonishingly rich dynamics*** and provides ***easily tractable time evolution***.

Introduction: Earlier literature

✦ Previous studies with relativistic multiple-shells systems:

— collision of two null shells

[Dray, 't Hooft (1985)] [Redmount (1985)]

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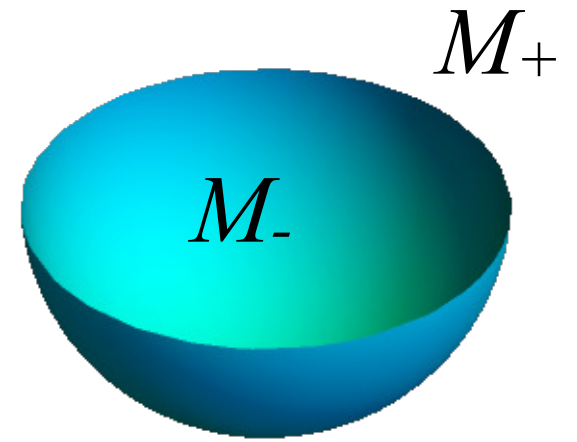
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Crucial difference: **absence of a confining mechanism**

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- ♦ For a **single** spherical shell: inner and outer geometries are both Schwarzschild-AdS.

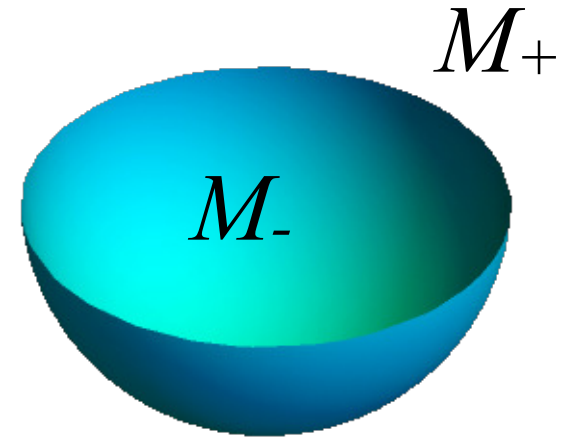
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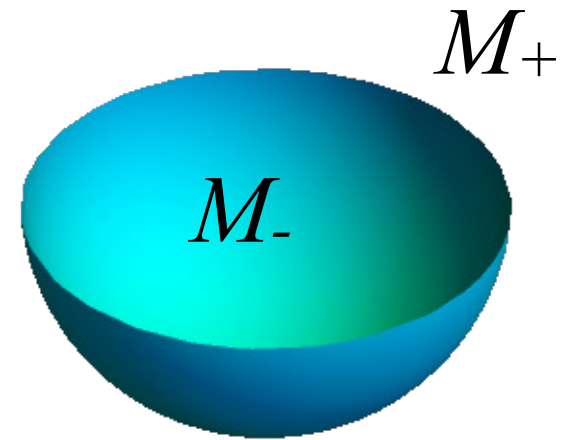
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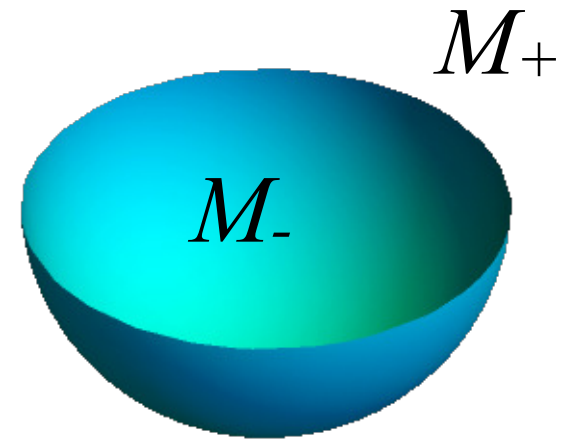
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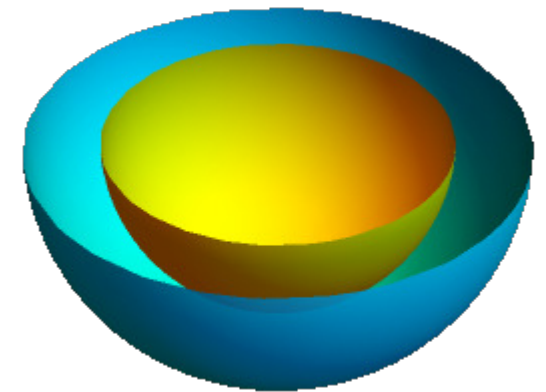
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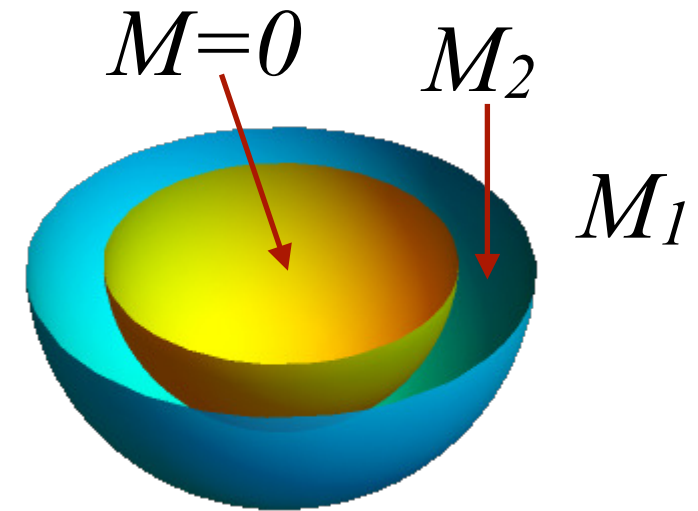
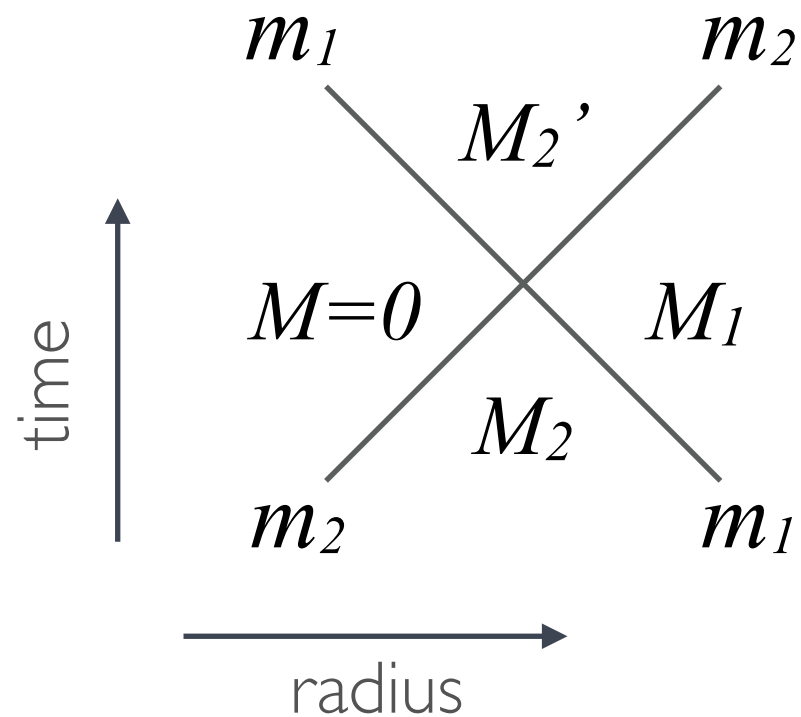
Double-shell systems: **Shell crossings**

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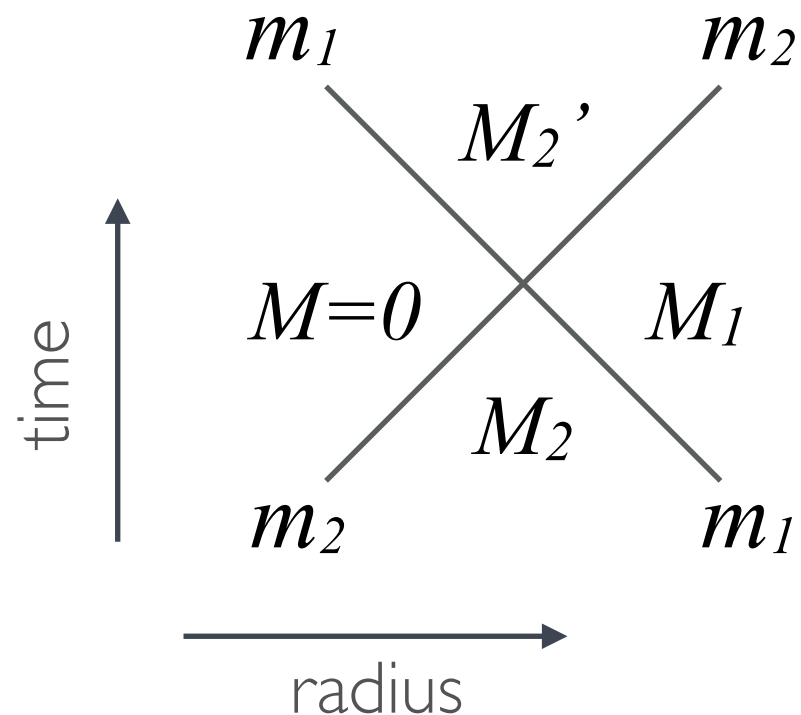
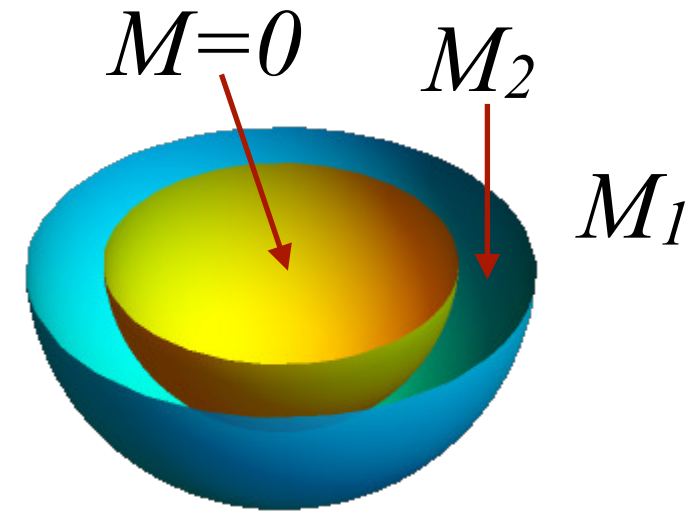
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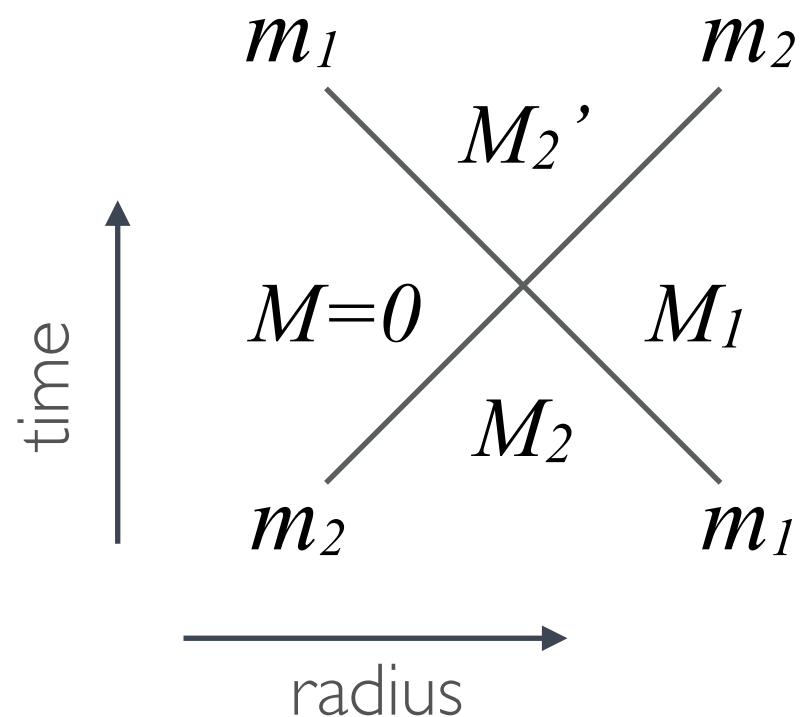
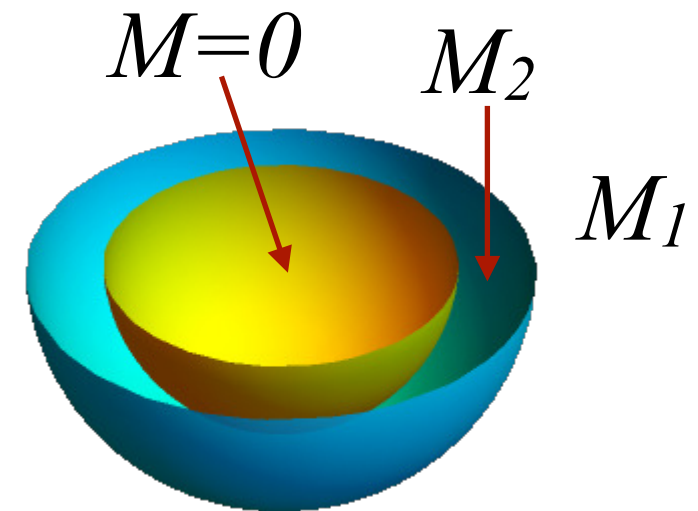
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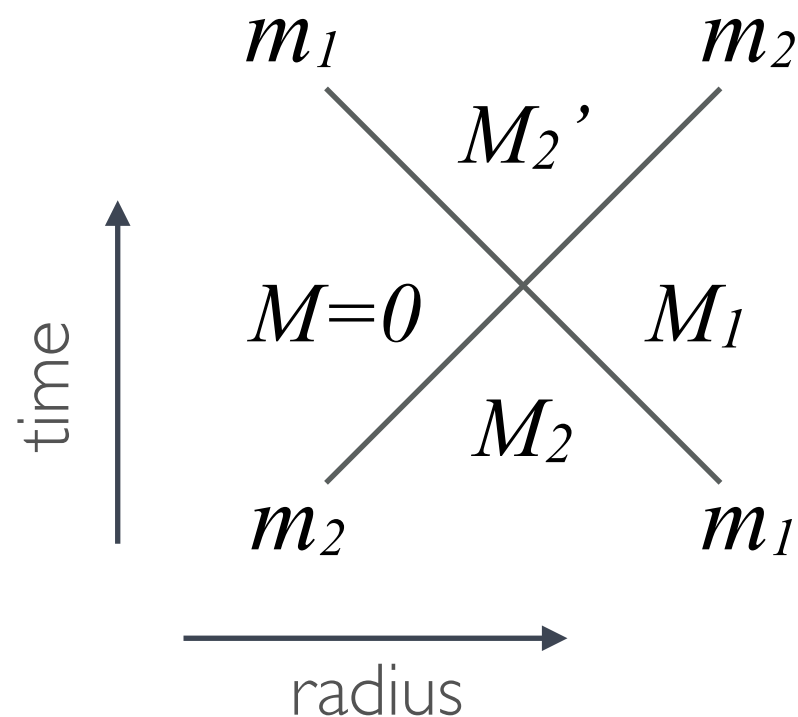
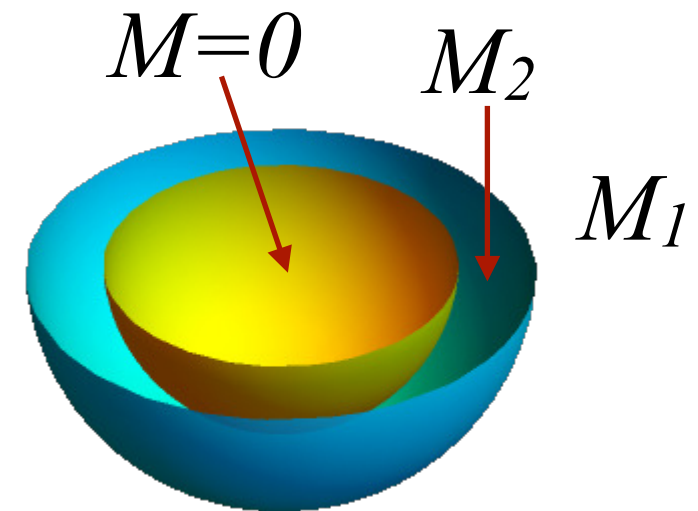
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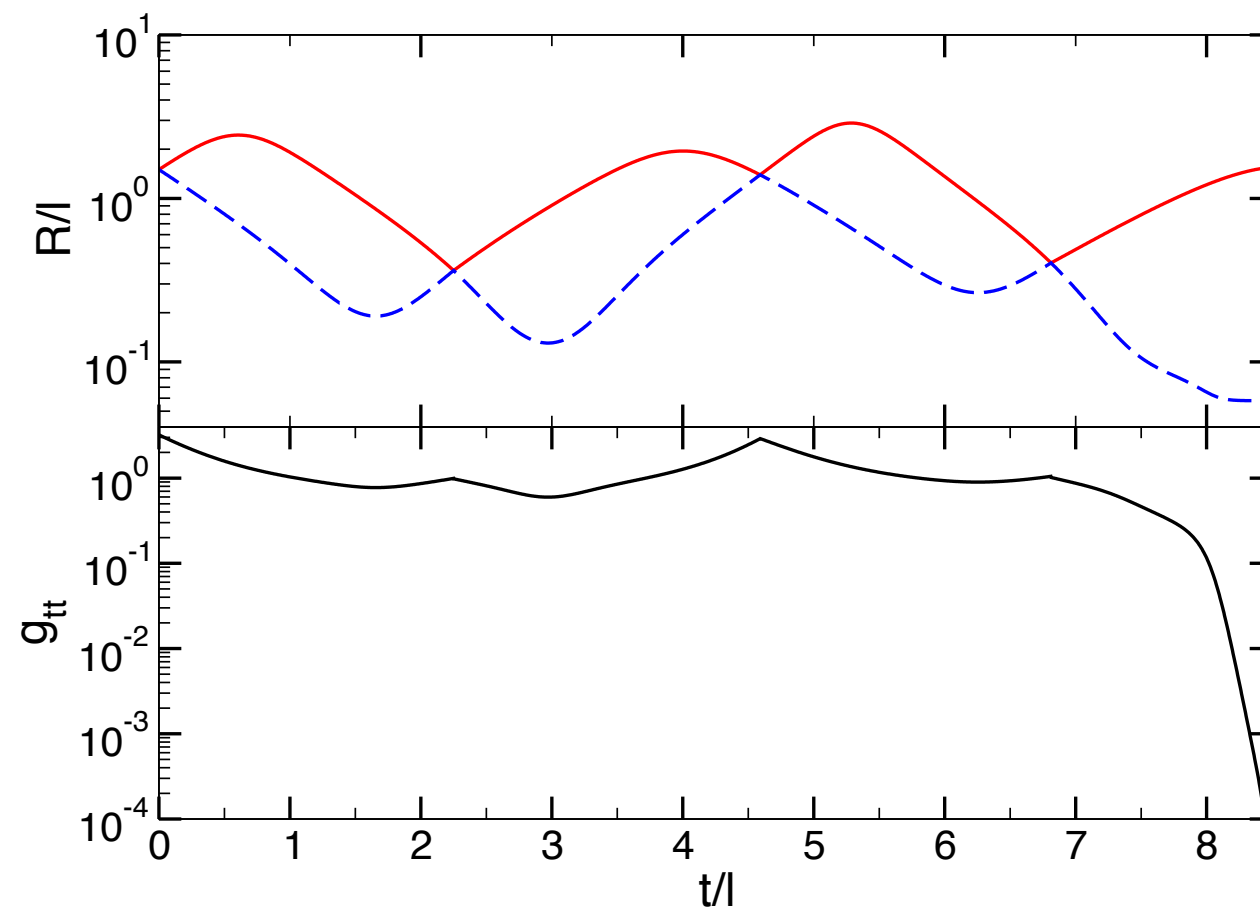
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Note: The outgoing shell always transfers energy to the ingoing shell.

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The confined double-shell system: Time evolution

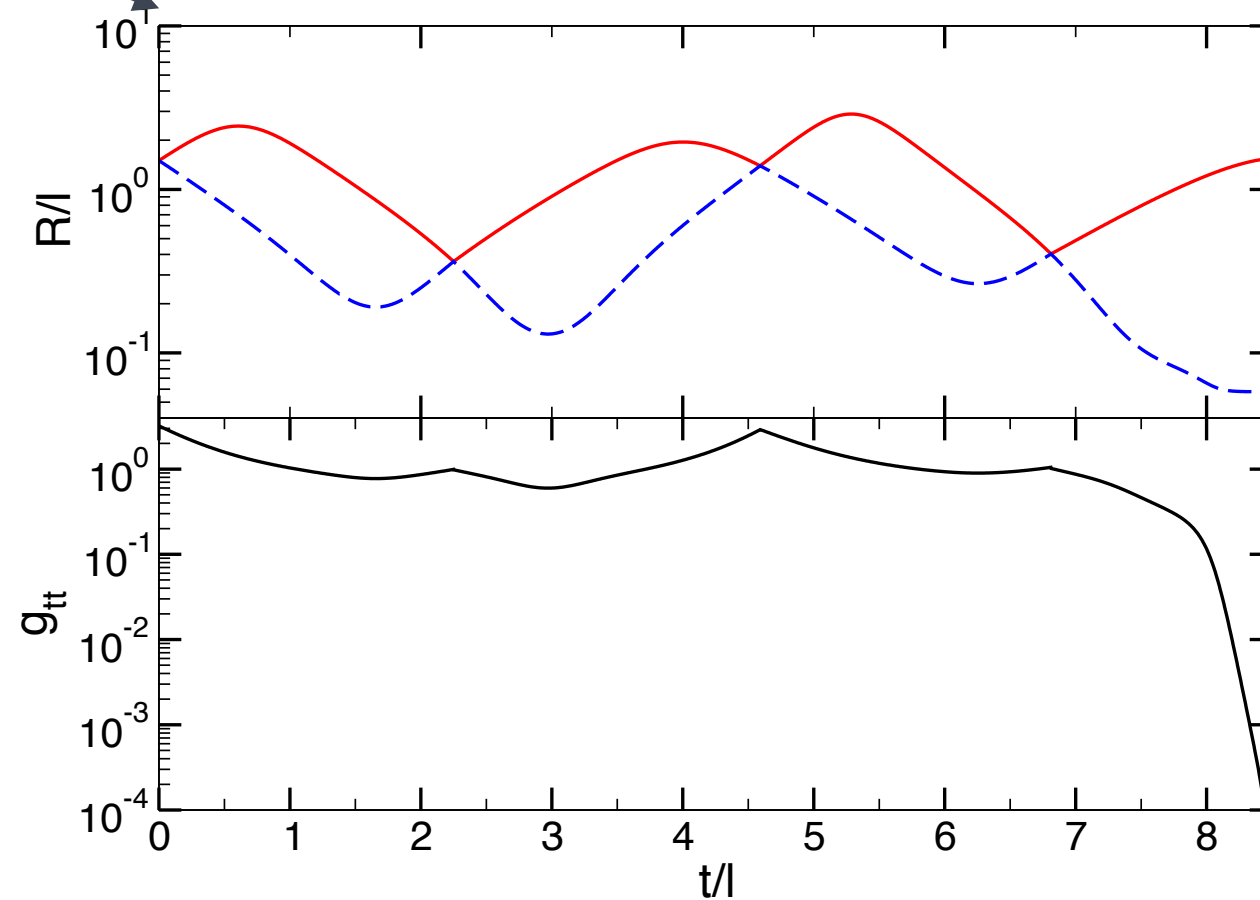
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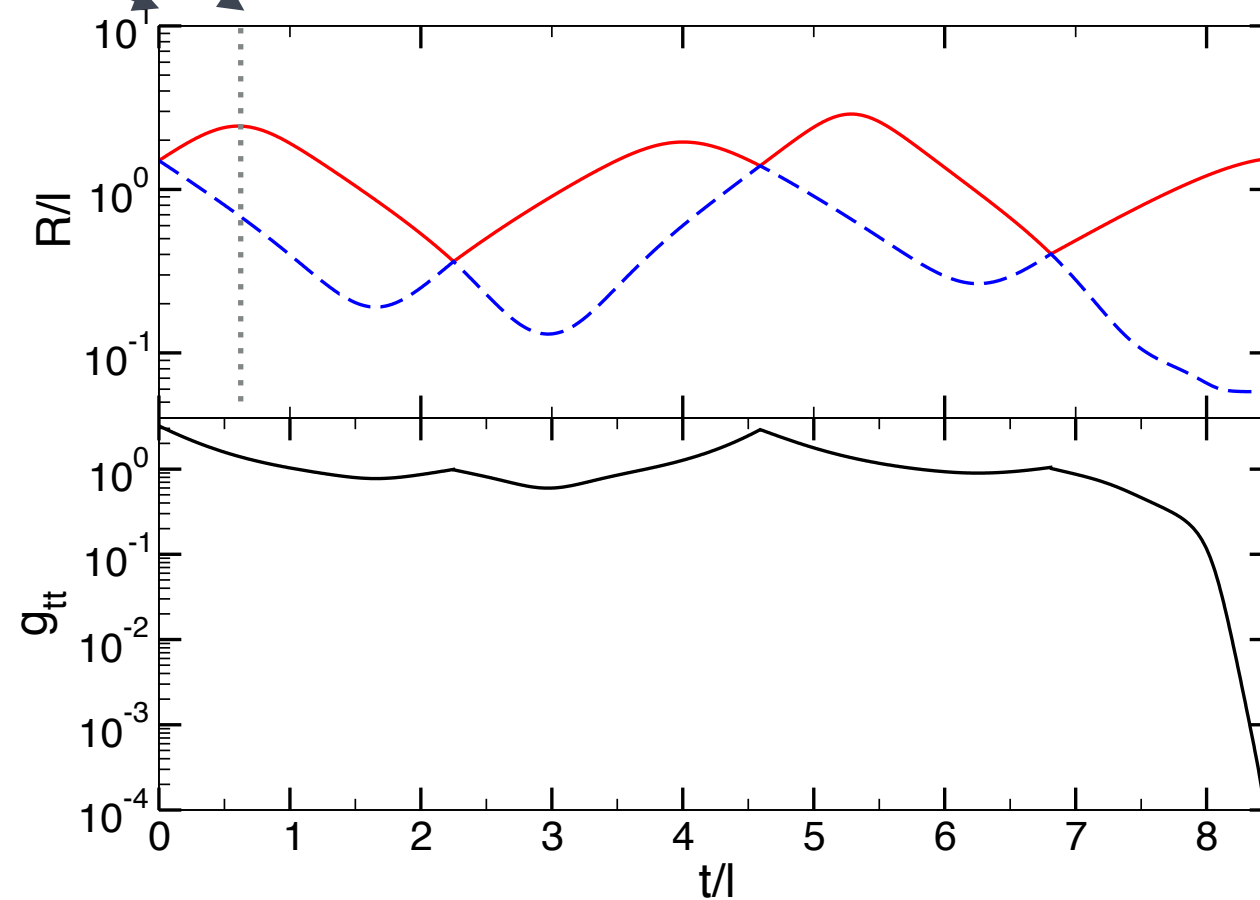
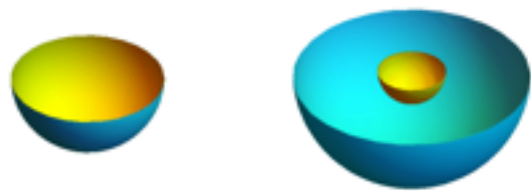
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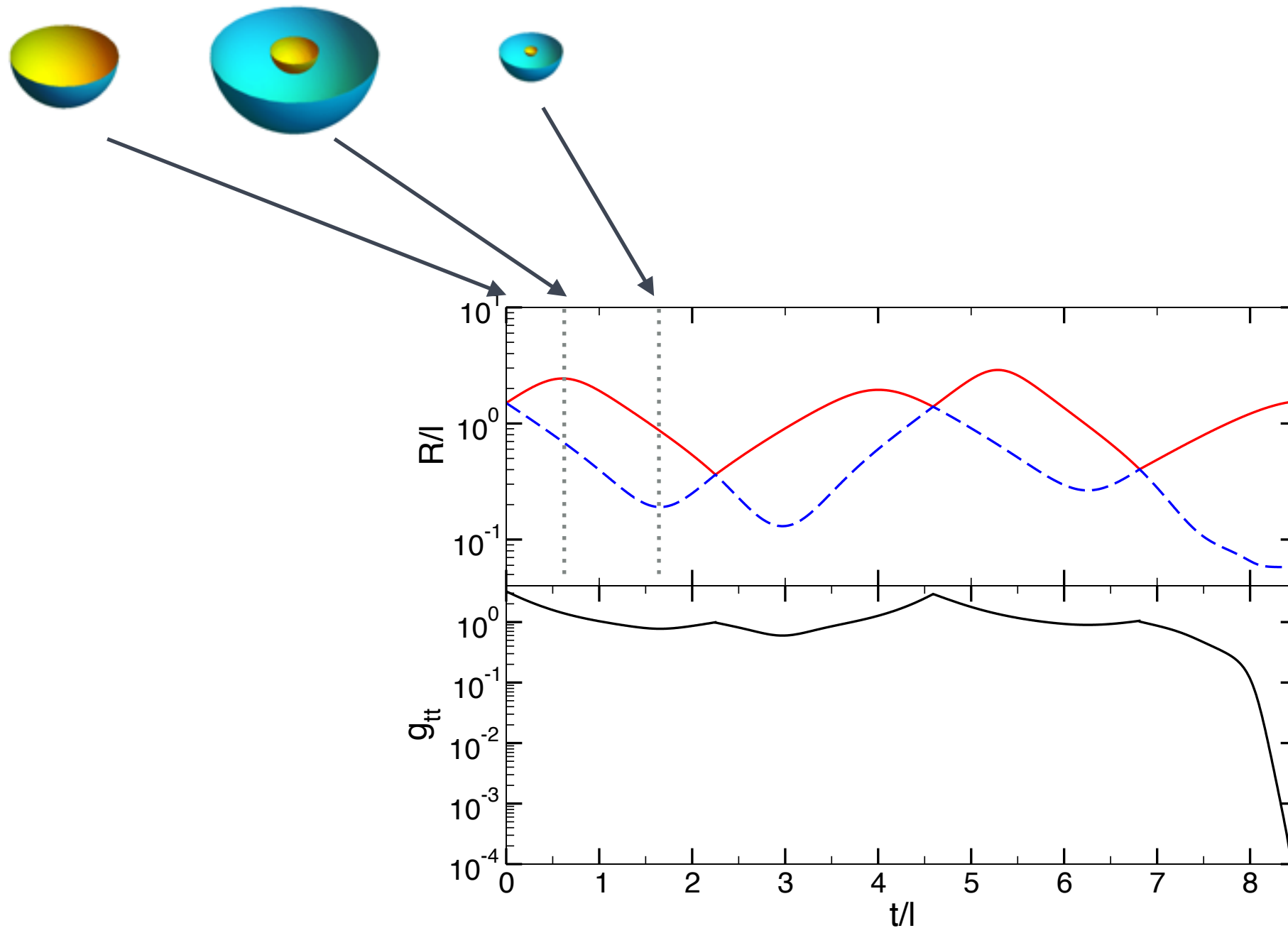
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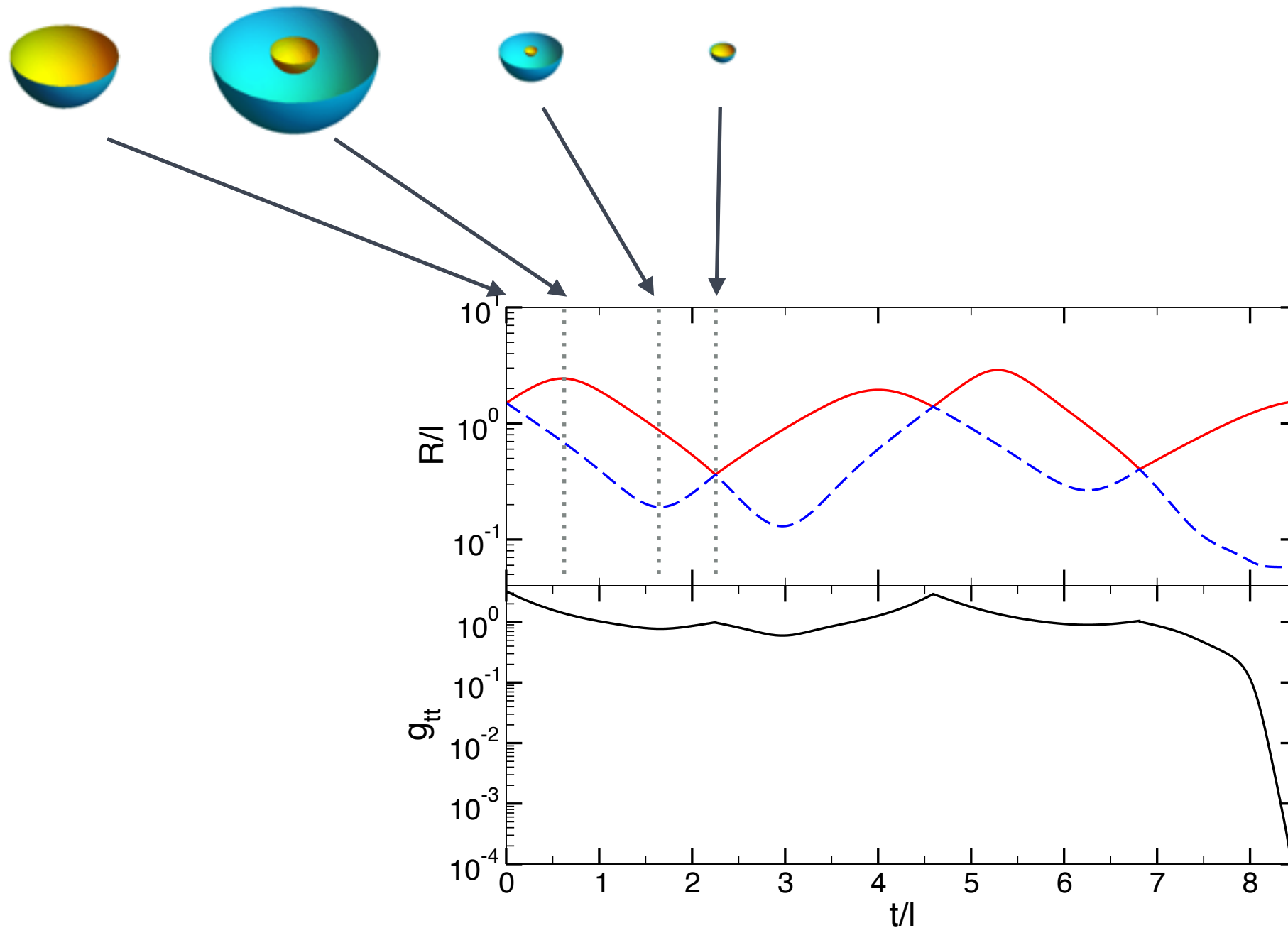
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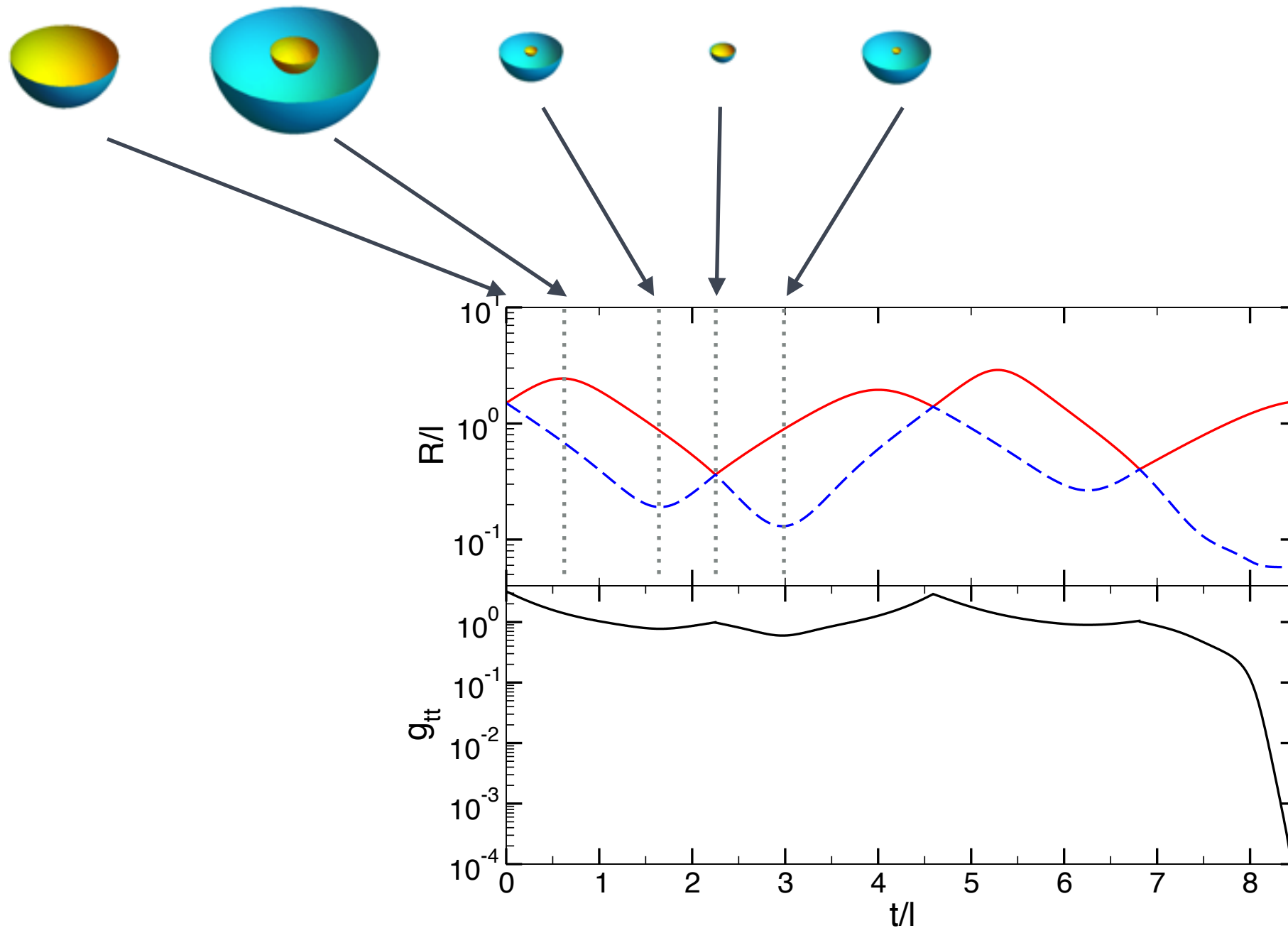
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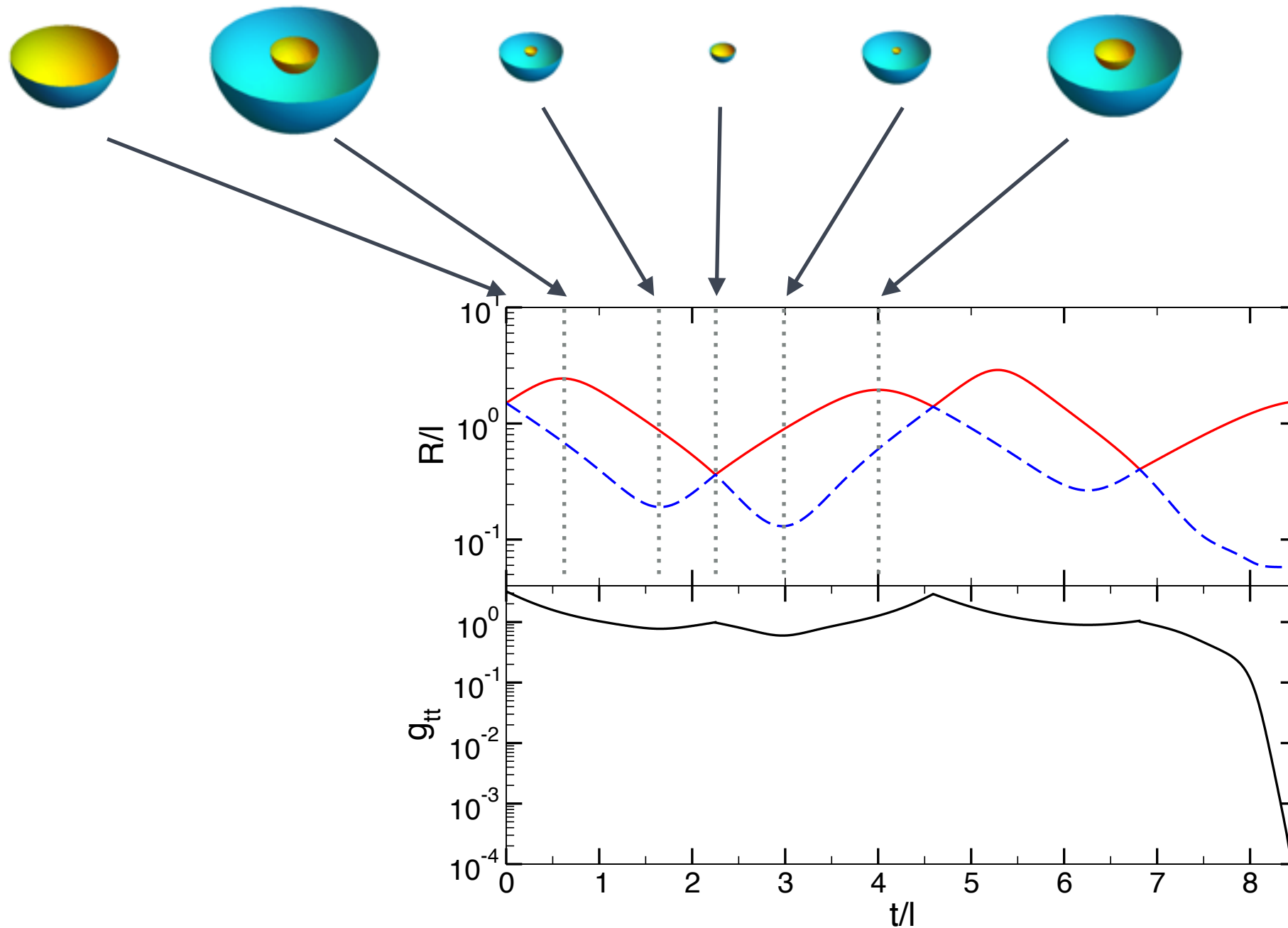
double-shell in AdS



$$\begin{aligned} M_1/l &= 0.05, \quad M_2/l = 0.025, \\ m_1/l &= m_2/l = 0.0136, \\ w_1 &= w_2 = 0.2, \\ R_i/l &= 1.5. \end{aligned}$$

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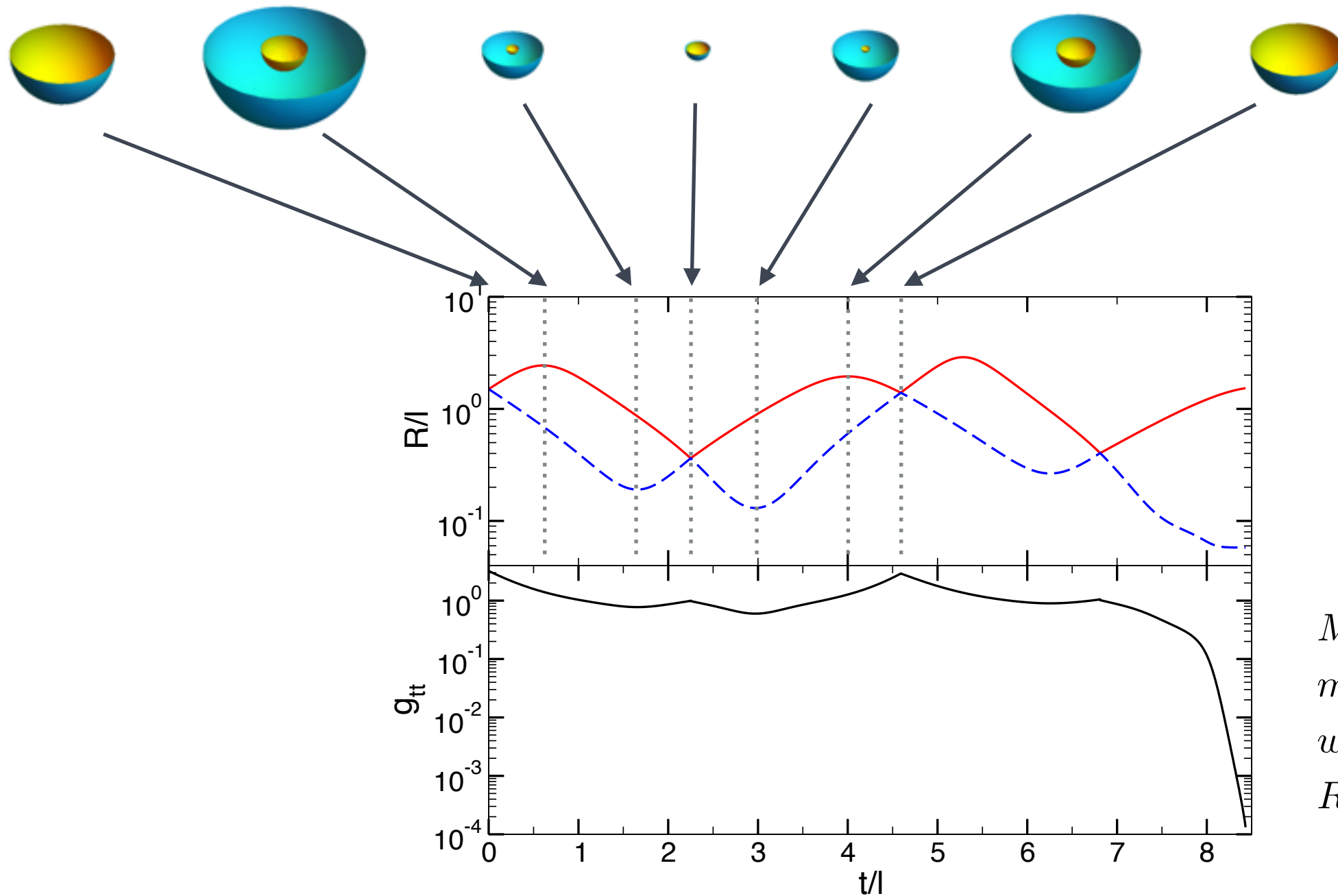
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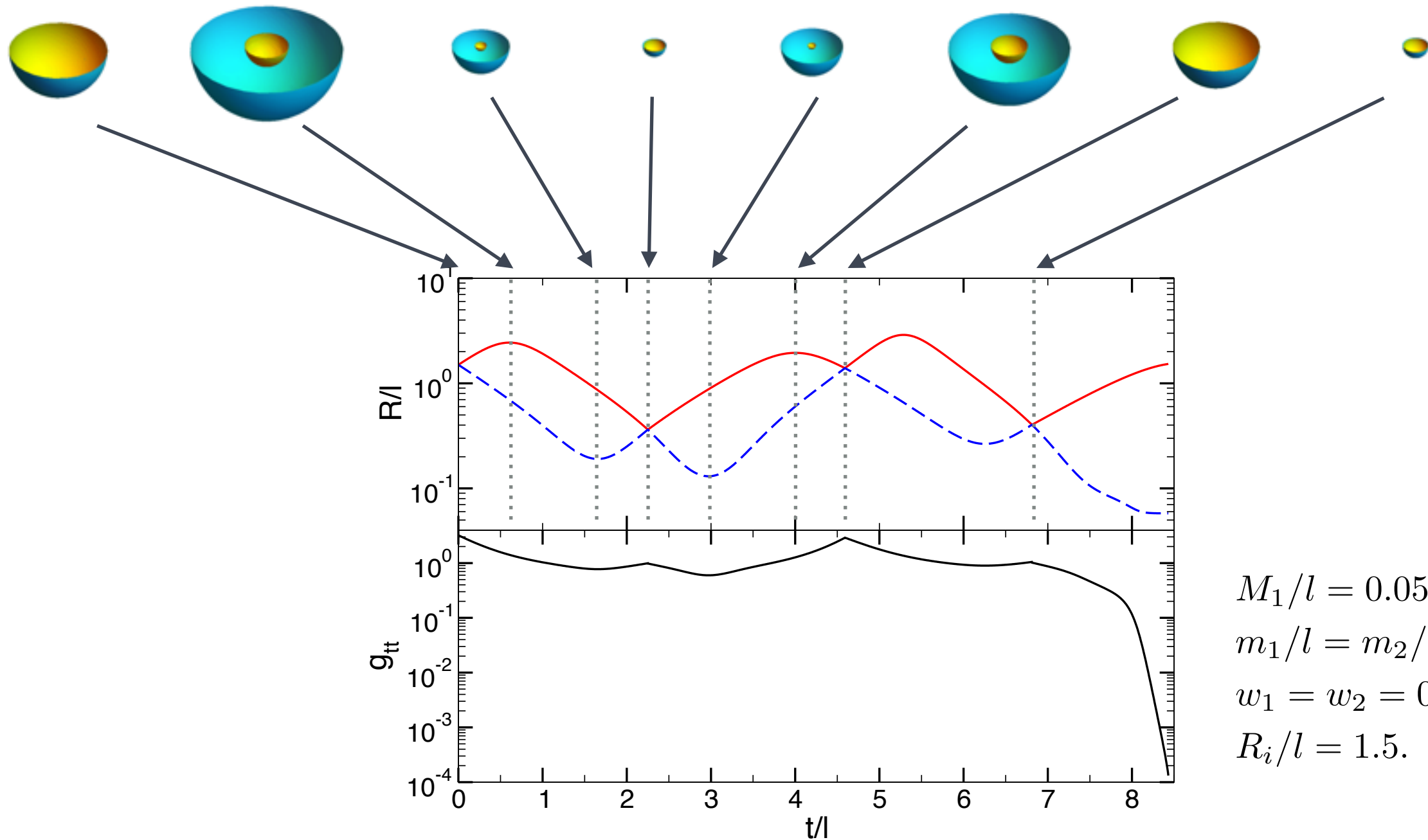
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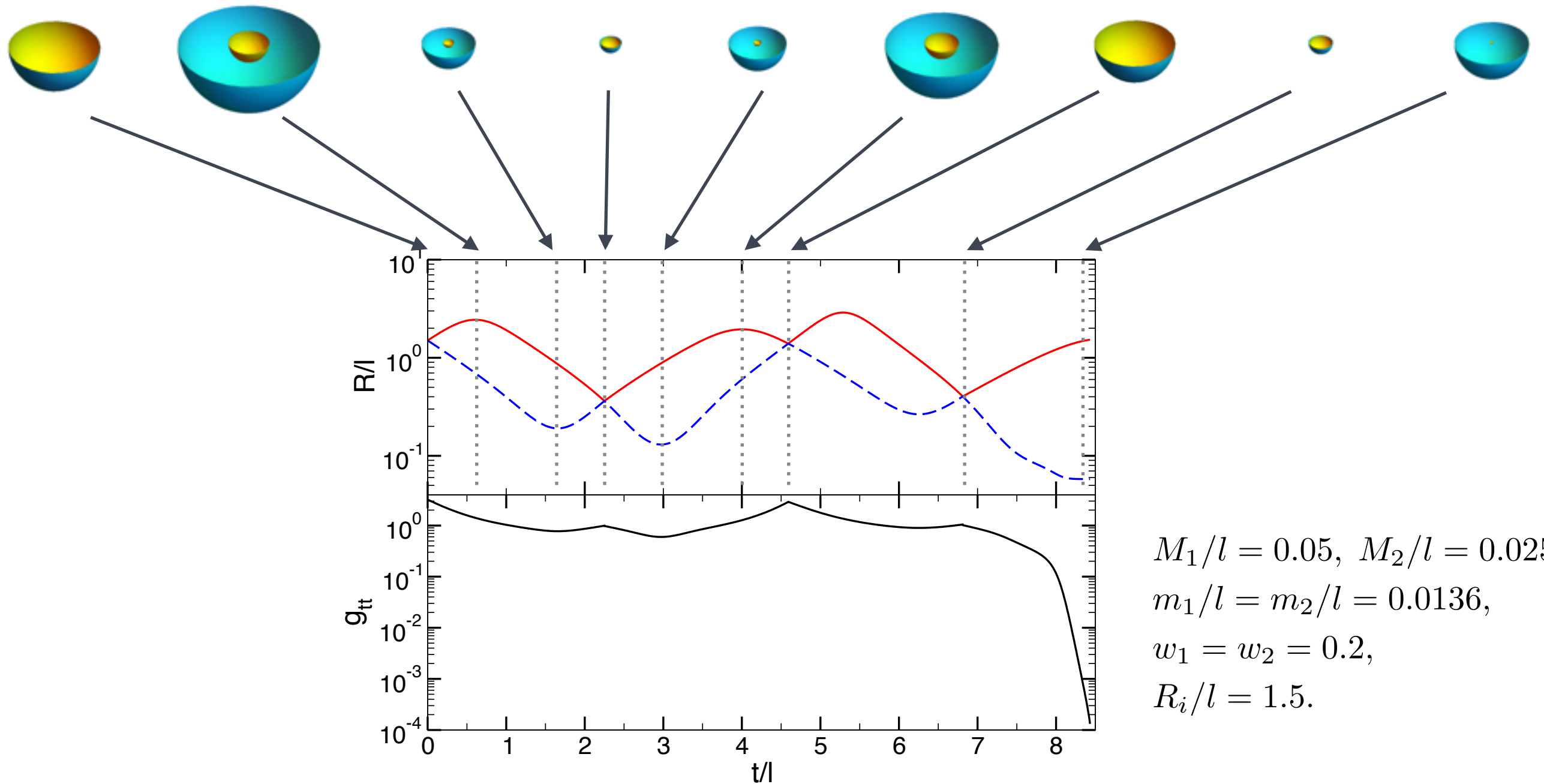
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Results: Critical phenomena

- ✦ Number of crossings before collapse **depends sensitively** on **initial conditions**.

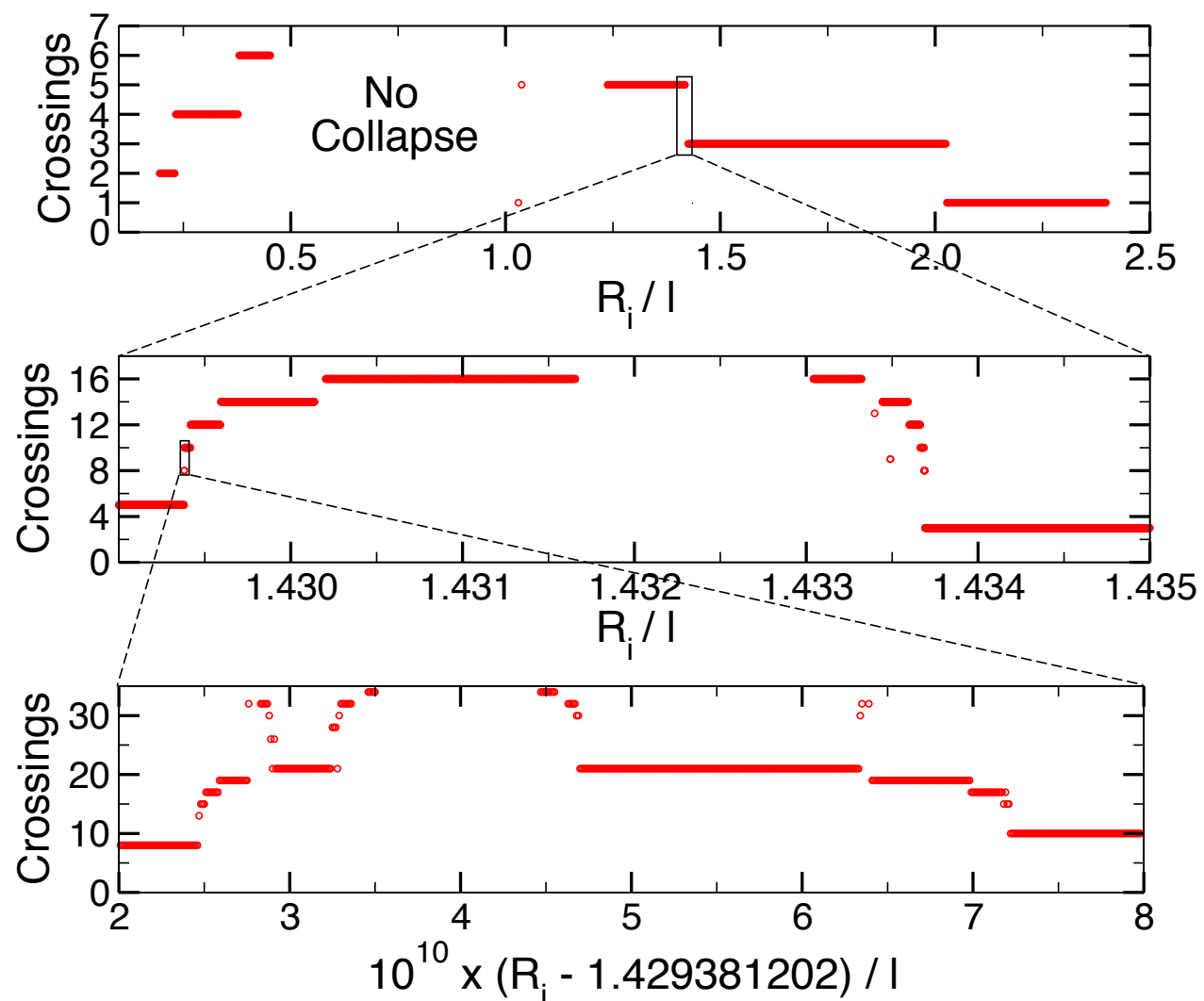
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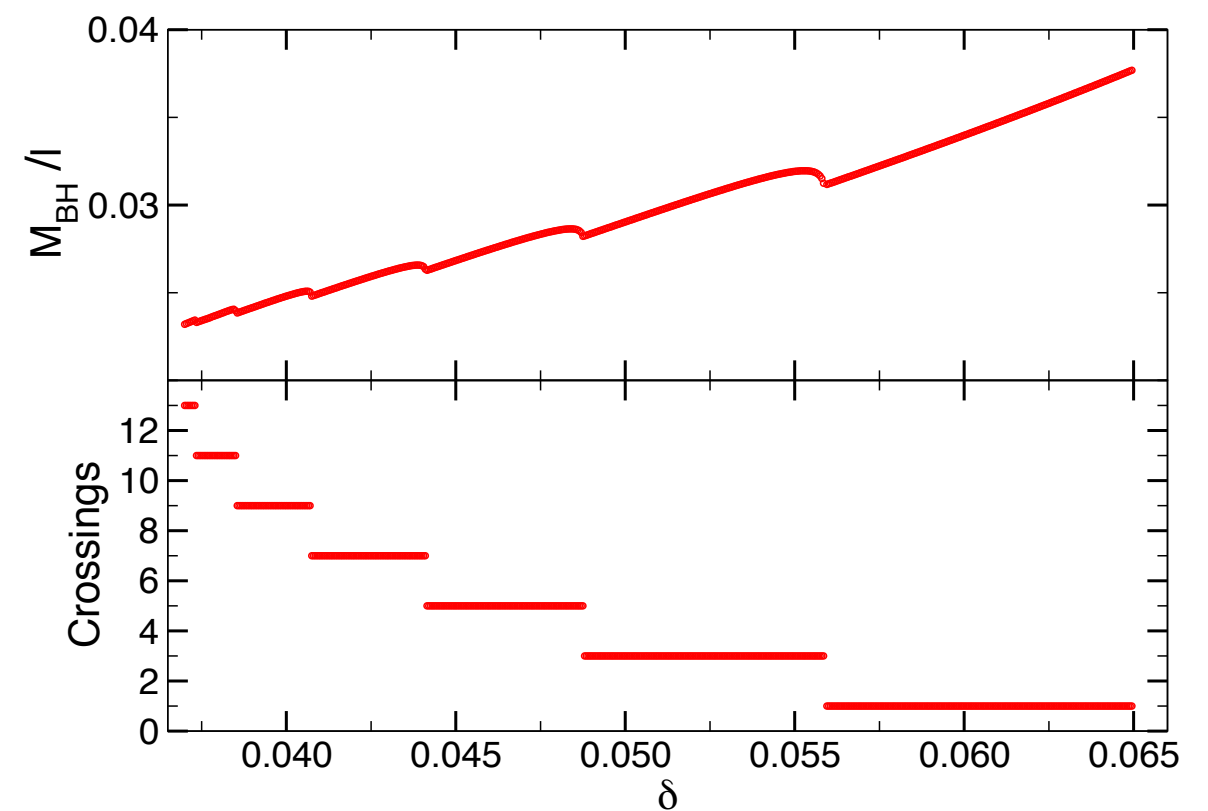
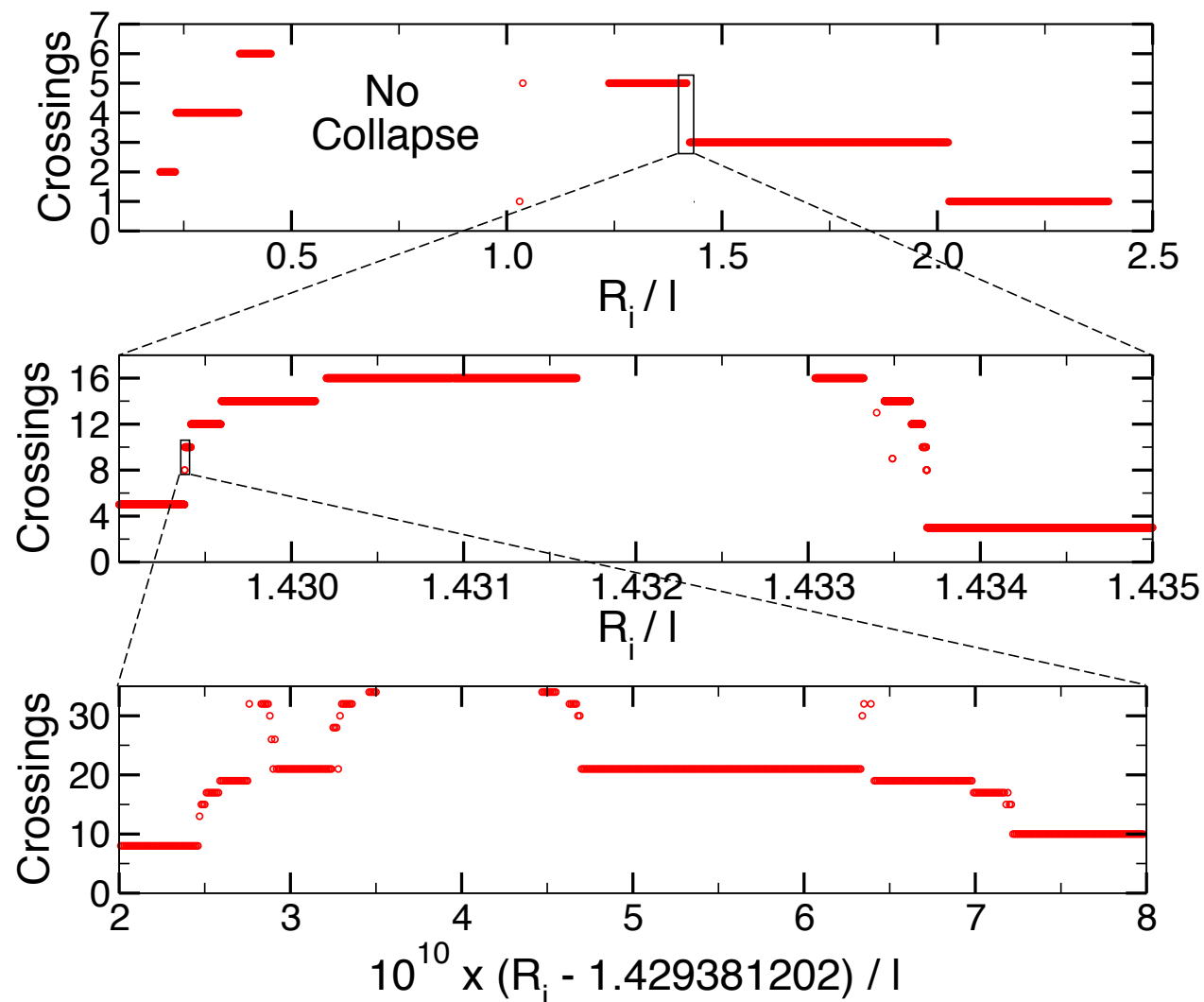
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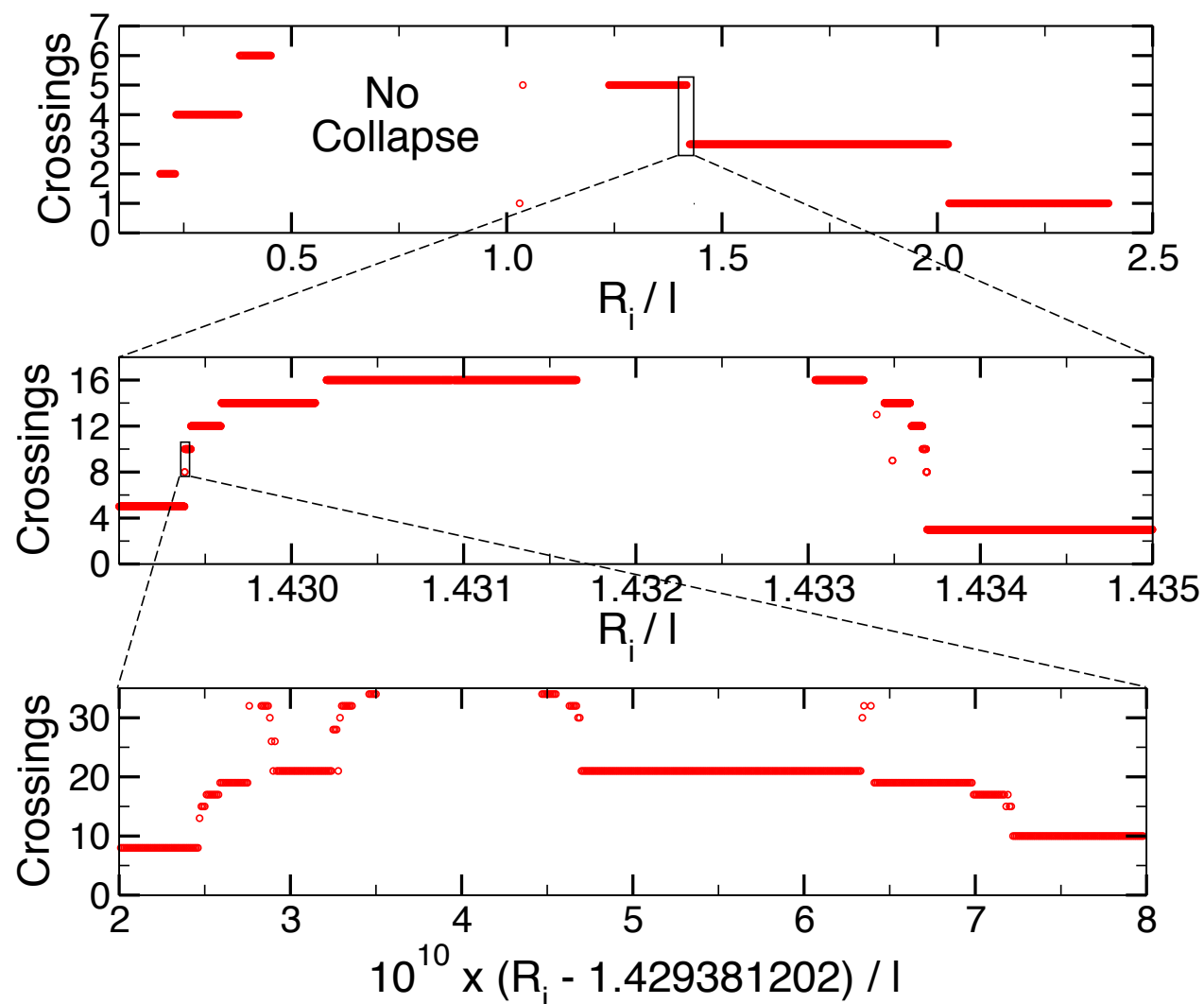


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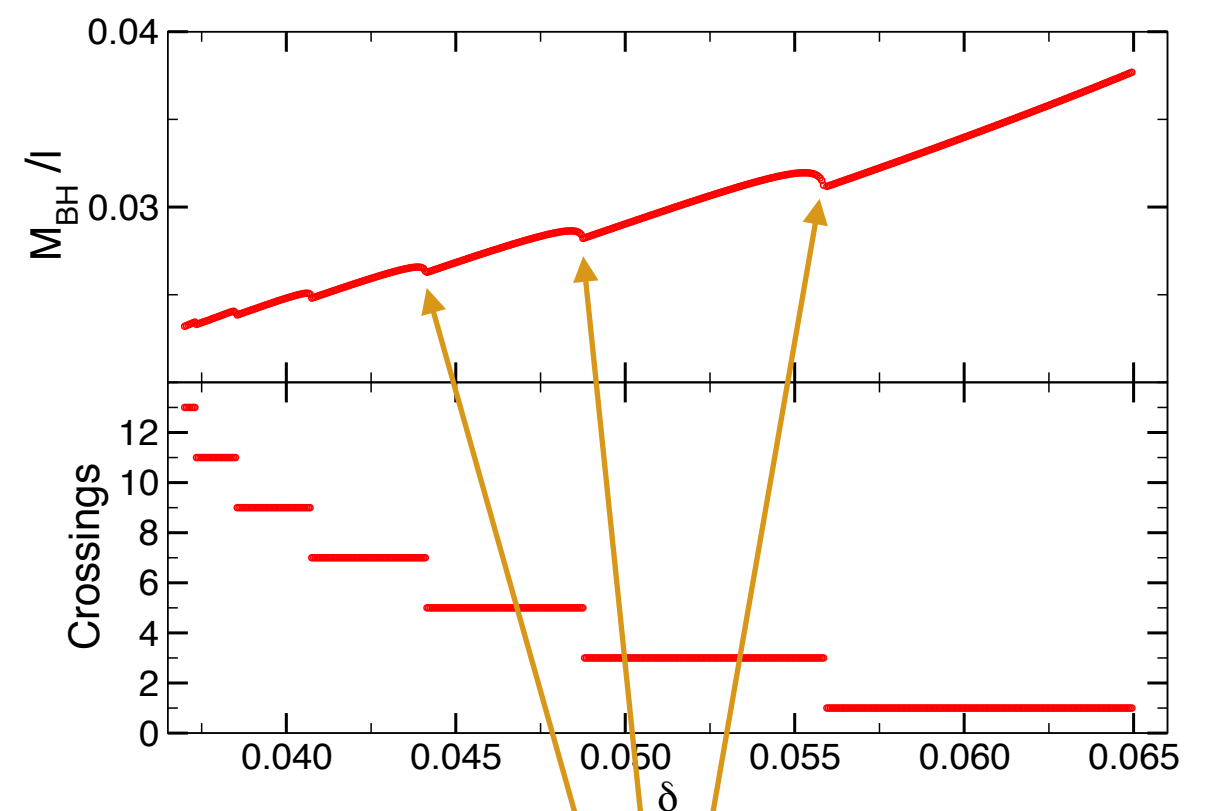
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$$M_{BH} - M_0 \propto |\delta - \delta_*|^\gamma$$

Results: Comparison between ‘cavity’ and ‘AdS’

- ✦ The critical exponent γ is sensitive to the matter content, i.e. the choice of w .

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- ✦ A similar trend has been observed recently with massless scalar fields.

[Santos-Oliván, Sopena (2015)]

[Cai, Yang (2016)]

Results: Chaotic behavior

- ♦ The double-shell system in Newtonian gravity (without CC) is chaotic.

How about in GR, in AdS?

[Miller, Youngkins (1997)]

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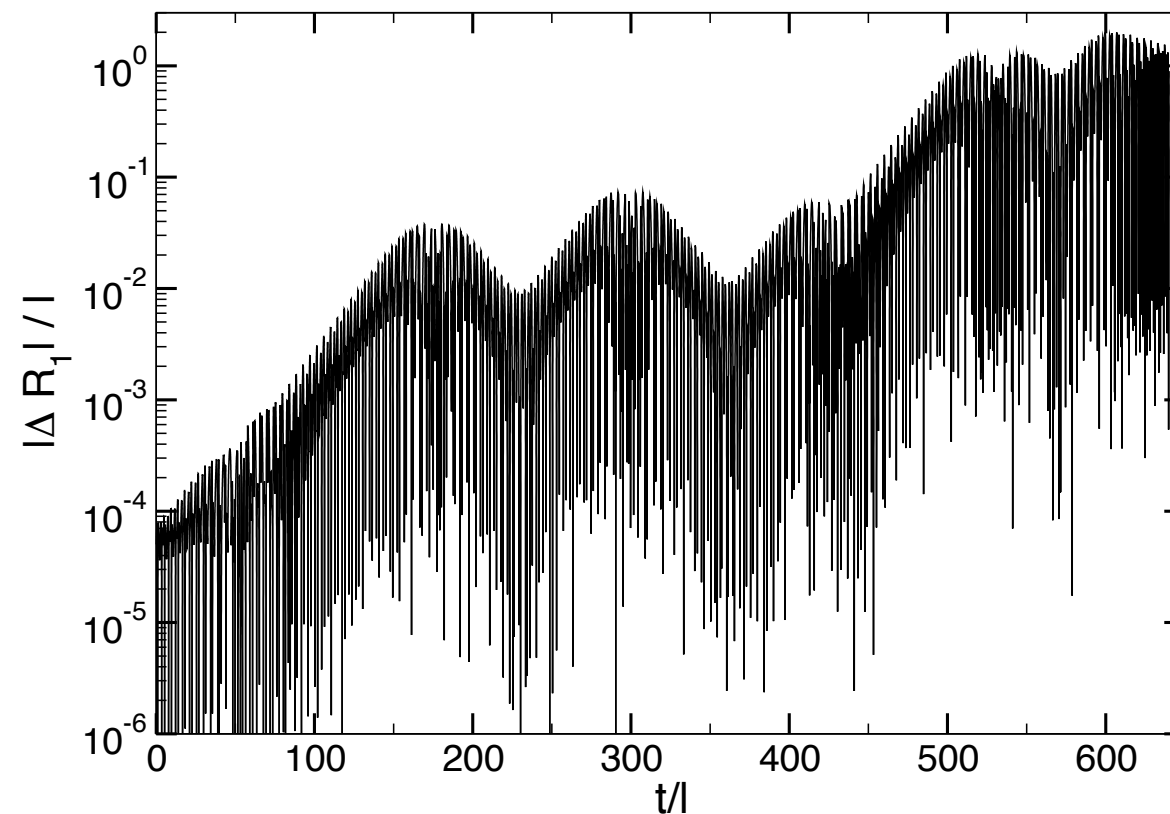
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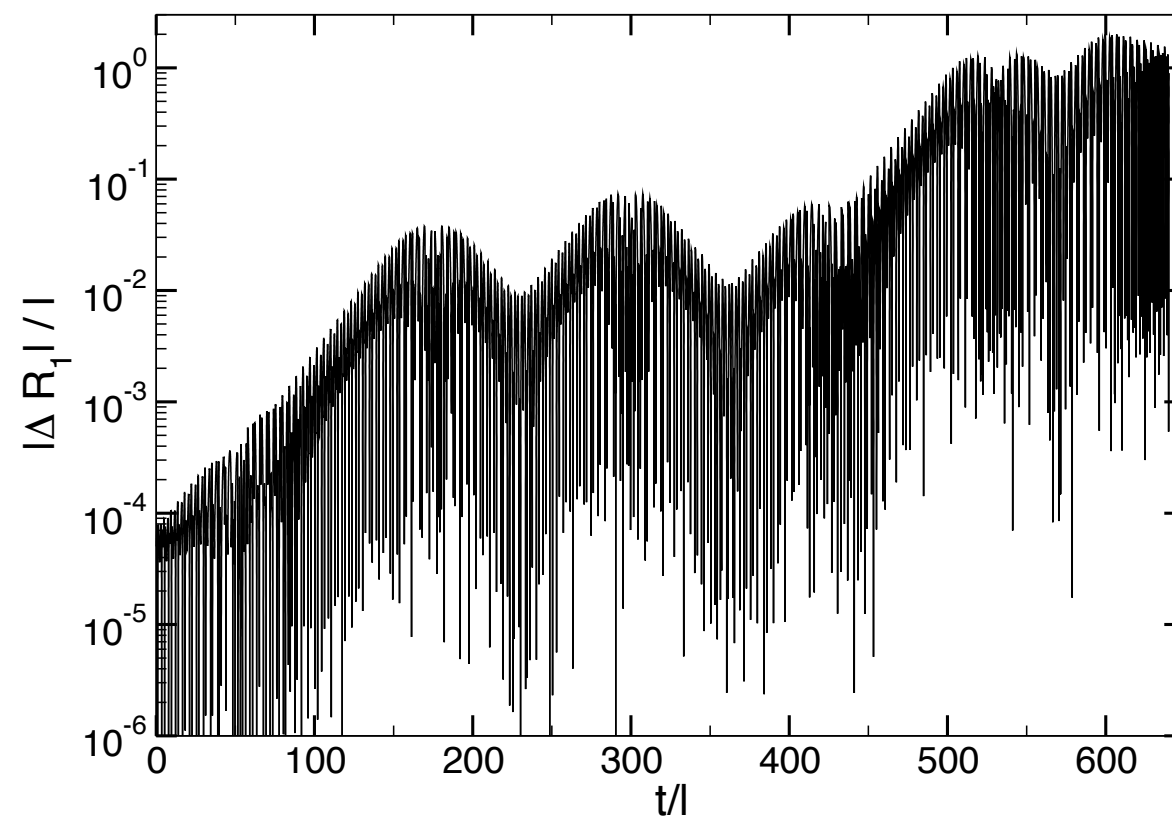
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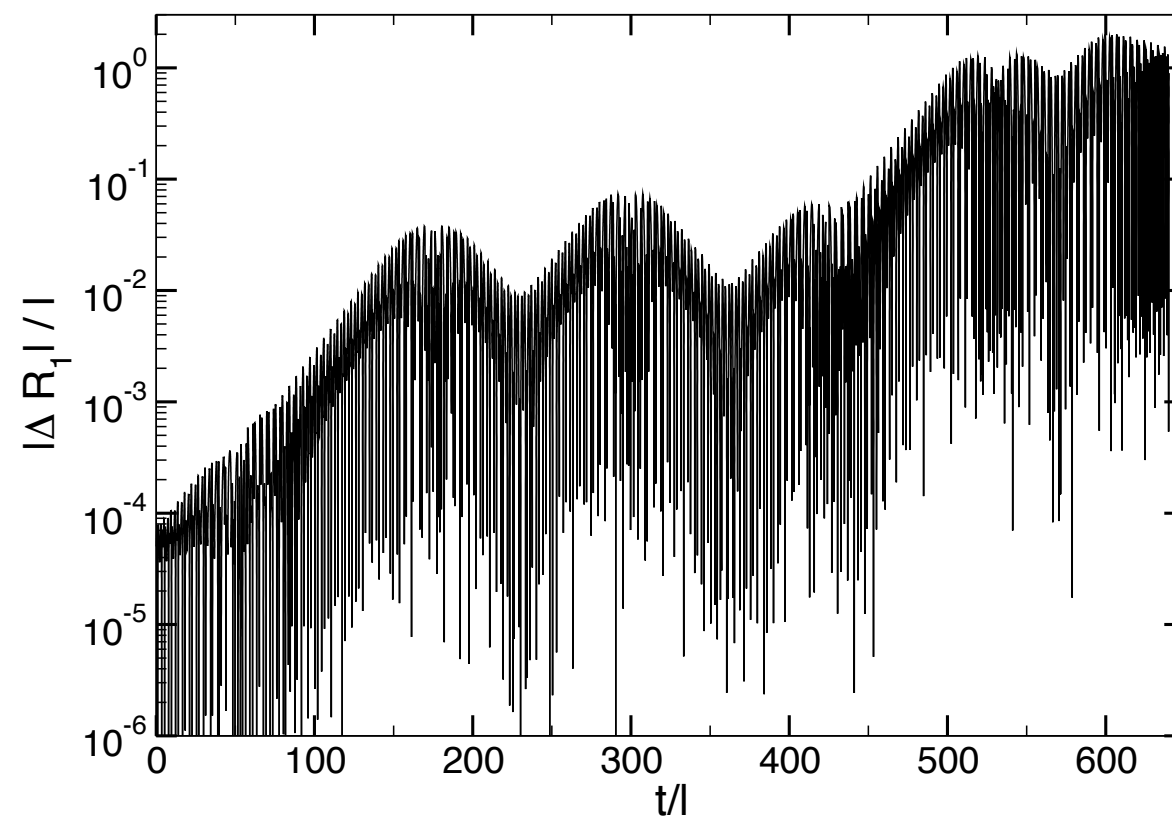
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- ♦ Observe **exponential growth** in the separation between two **initially nearby orbits**. Eventually growth **saturates** because system is **confined**.
- ♦ Similar behavior was observed for a massless scalar field in AdS.

[Oliveira, Pando Zayas, Terrero-Escalante (2012)]

[Farahi, Pando Zayas (2014)]

Conclusion:

- ✦ The **simplicity** + **richness** afforded by this setup provides an **ideal testbed** for explorations of *gravitational collapse in confining geometries* and its *holographic dual interpretations*.

The study of the dynamics of these systems only requires **solving two decoupled ODEs**.

- ✦ Depending on initial conditions one finds:
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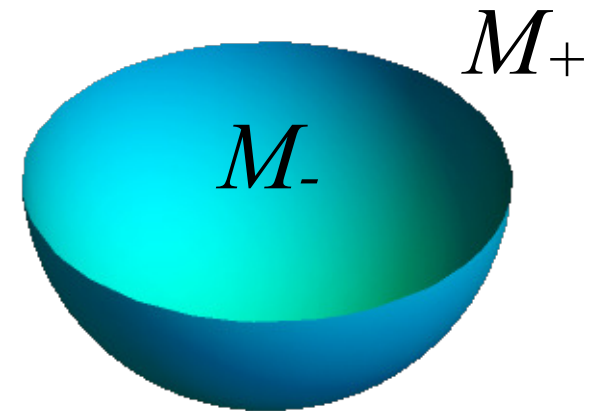
Thank you. 

Extra: Junction conditions

- Start with a **single** spherical shell.

- Geometries inside and outside are both Schwarzschild-AdS:

$$ds^2 = -f(r)dt^2 + f(r)^{-1}dr^2 + r^2d\Omega^2, \quad f(r) = \left(1 - \frac{2M}{r} + \frac{r^2}{l^2}\right)$$



- Apply **junction conditions** for the timelike hypersurface, $t = T(\tau)$, $r = R(\tau)$.

induced metric $\longrightarrow \mathfrak{g}_{ij}^{(+)} = \mathfrak{g}_{ij}^{(-)} \equiv \mathfrak{g}_{ij},$

extrinsic curvature $\longrightarrow (k_{ij}^{(+)} - k_{ij}^{(-)}) - \mathfrak{g}_{ij}(k^{(+)} - k^{(-)}) = -8\pi G \mathcal{S}_{ij}$

shell's stress-
-energy tensor

- The 2nd junction condition dictates the form of the shell stress-energy tensor.

It's a **perfect fluid**:

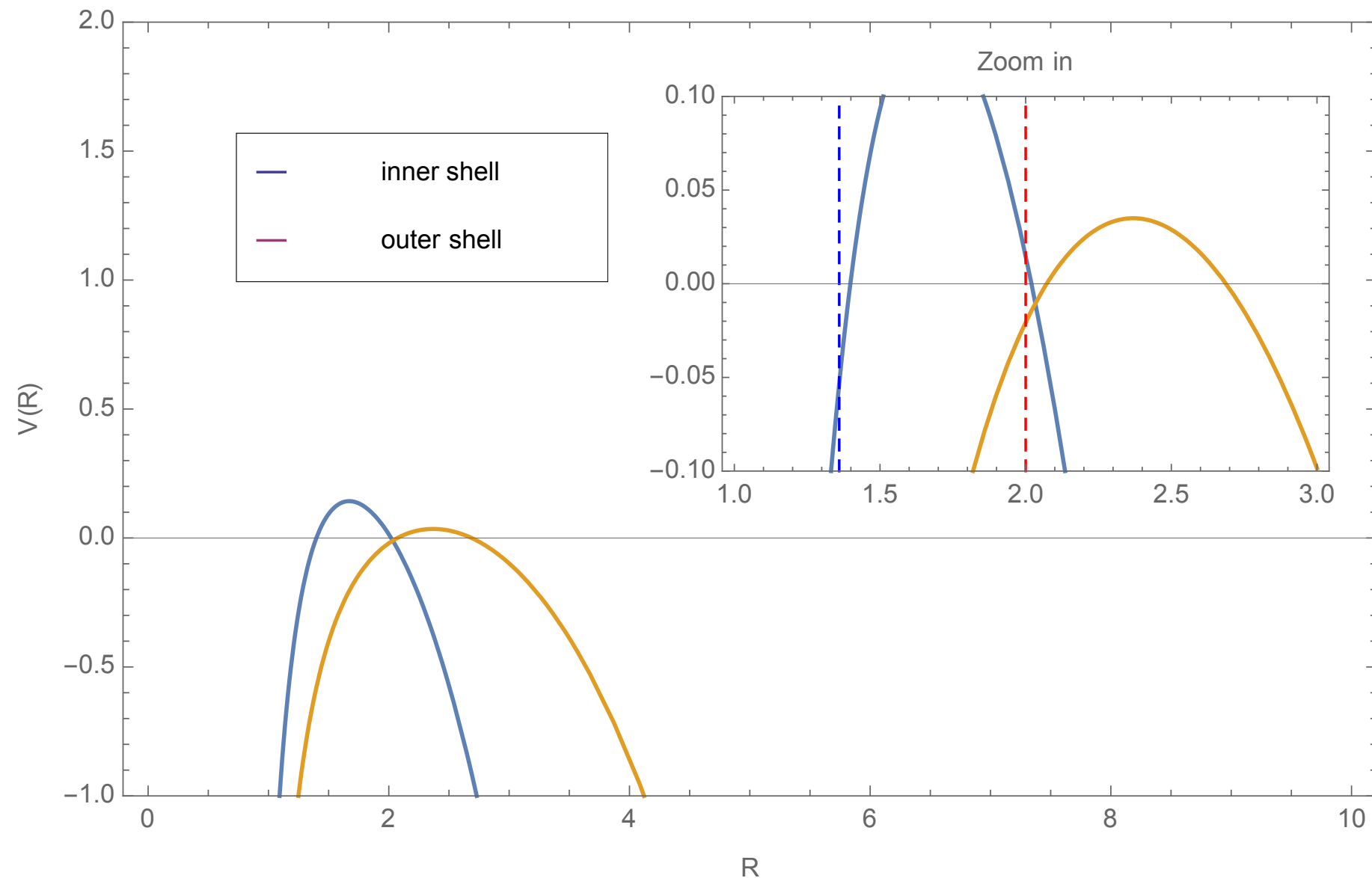
$$\mathcal{S}_{ij} = (\rho + P)u_i u_j + P \mathfrak{g}_{ij} \quad (u = \partial_\tau)$$

energy density

pressure

Extra: Non collapsing configurations

- ♦ **Self-sustained** non collapsing configurations of multiple shells can exist, even **without a confining mechanism**.



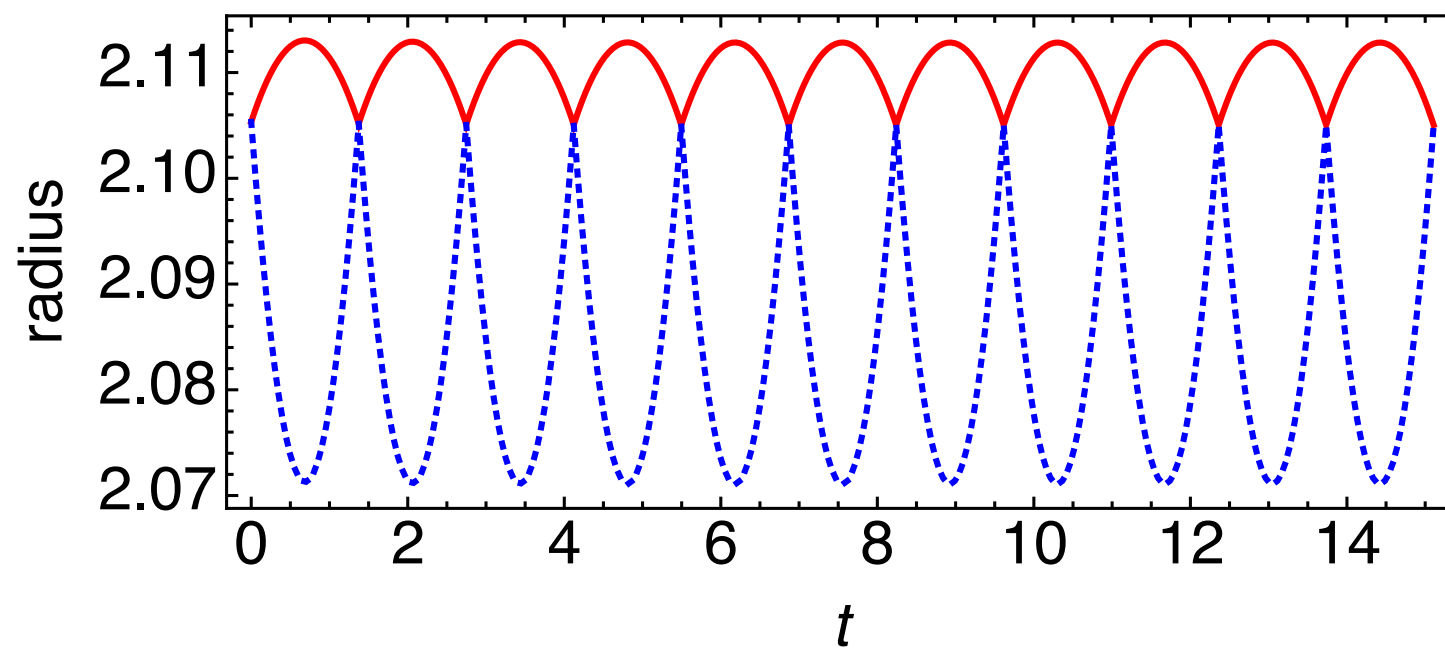
$$M_1 = 1$$

$$M_2 = 0.68$$

$$m_1 = m_2 = 2.9$$

$$w_1 = w_2 = 0.86$$

Extra: Non collapsing configurations



$$M_1 = 1$$

$$M_2 = 0.663$$

$$m_1 = m_2 = 2.9$$

$$w_1 = w_2 = 0.86$$

Extra: Ida-Nakao formula

- ✦ Post-collision 'ADM' mass:

$$M'_2 = -(\nu_1^2 + \nu_2^2) \frac{r_0}{4} + (f_2 - 1 - f_1) \frac{r_0}{4} - (\nu_1^2 - 1)(\nu_2^2 - f_1) \frac{r_0}{4f_2} + \frac{r_0}{2} \\ - \epsilon_1 \epsilon_2 \frac{r_0}{4f_2} \left[\nu_1^4 - 2(1 + f_2)\nu_1^2 + (1 - f_2)^2 \right]^{1/2} \left[\nu_2^4 - 2(f_1 + f_2)\nu_2^2 + (f_1 - f_2)^2 \right]^{1/2}$$

where

$$f_i = 1 - \frac{2M_i}{r_0},$$

$$\nu_i = \frac{m_i}{r_0^{1+2w}},$$

$$\epsilon_i = \pm 1.$$

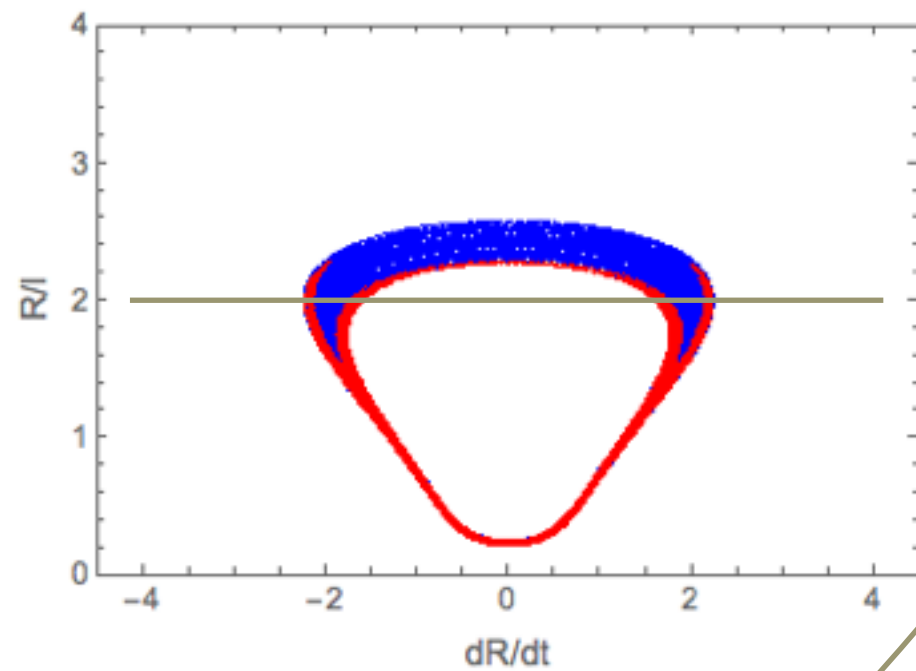
collision radius

expanding shell

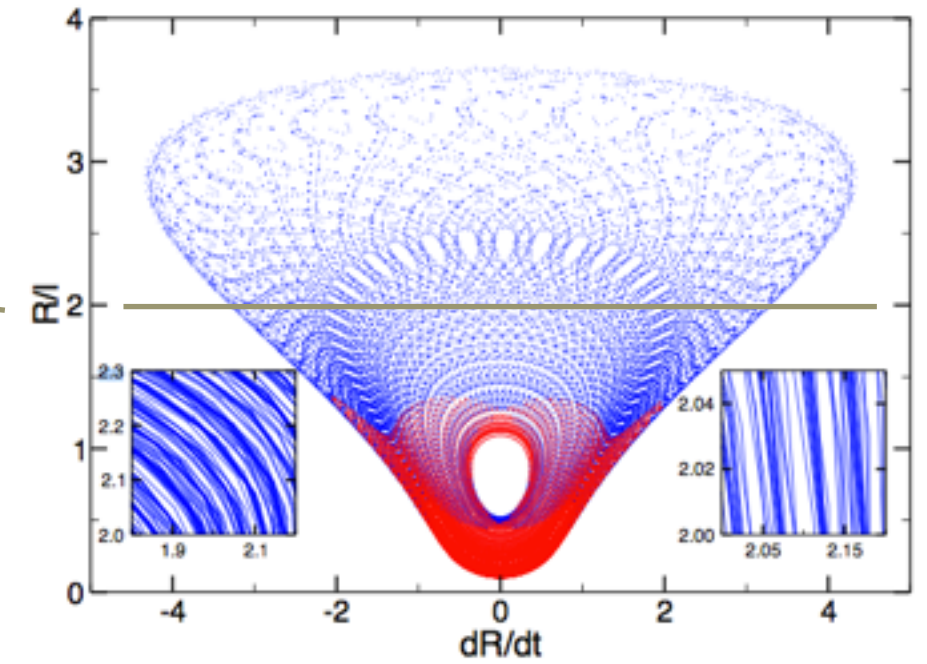
contracting shell

Extra: Chaotic phase space

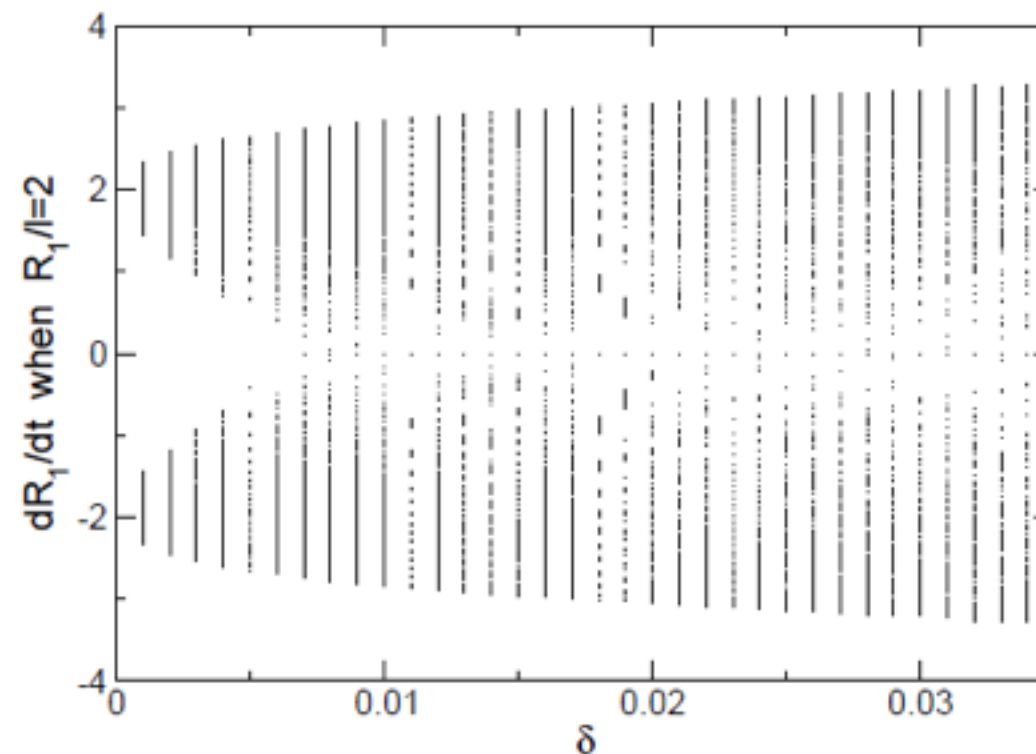
- ♦ Scan non-collapsing configurations: system is **quasi-periodic** in the low energy regime.



$\delta = 0.0005$



$\delta = 0.03$



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