



MAX-PLANCK-GESELLSCHAFT



Fourier-domain modulation and delay of  
gravitational wave signals: application to the  
response of LISA-type detectors and to  
precessing binaries

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# Motivation

## (e)LISA prospective parameter estimation

- Bayesian sampling for parameter extraction is expensive: several  $10^6$  samples - limitation so far to inspiral and/or Fisher matrix
- Impact of merger/ringdown, higher modes, spins, precession, instrument design ?
- Available fast FD IMR waveform models: aligned spins (SEOBNRv2 ROM, PhenomD), precession (PhenomP, NR surrogate), compact FD amplitude/phase representation
- Fourier-domain response of the instrument for IMR waveforms ?

## Modeling waveforms from precessing binaries

- Precessing waveform as a frame rotation of non-precessing waveform
- Fourier-domain transposition: beyond the inspiral, extension to merger-ringdown ?

# Delayed and modulated signals

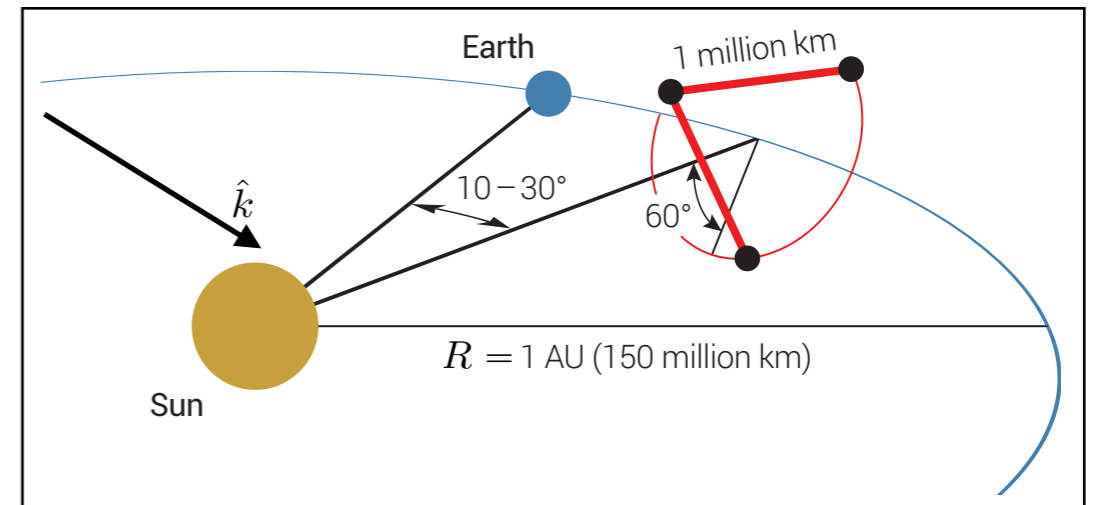
## (e)LISA response

Frequency observables:

$$y = \Delta\nu/\nu$$

$$y_{slr} = \Phi_l(t_s - \hat{k} \cdot p_s) - \Phi_l(t_r - \hat{k} \cdot p_r)$$

$$\Phi_l = \frac{1}{2} \frac{1}{1 - \hat{k} \cdot n_l} n_l \cdot h^{\text{TT}} \cdot n_l$$



$$\text{FT}[F(t)h(t + d(t))] \leftrightarrow \tilde{h}(f), F(t), d(t)$$

$F, d$  Instrument timescale (1yr)  $\longleftrightarrow$   $h$  Waveform timescale

## Frame rotation for precessing waveforms

$$h_{\ell m}^{\text{I}} = \sum_{m'} D_{m' m}^{\ell*}(\alpha, \beta, \gamma) h_{\ell m'}^{\text{P}}$$

- Inertial-frame  $h^{\text{I}}$  obtained as a rotation of a precessing-frame  $h^{\text{P}}$
- Approximate  $h^{\text{P}}$  as a non-precessing waveform (SpinTaylorF2, PhenomP, SEOBNR inspiral)

$$\text{FT}[F(t)h(t)] \leftrightarrow \tilde{h}(f), F(t)$$

$F$  Precessional timescale  $\longleftrightarrow$   $h$  Waveform timescale

# Delays and modulations in Fourier domain

A general view

$$\tilde{h}(f) = A(f)e^{-i\Psi(f)}$$

$$s(t) = F(t)h(t + d(t))$$

$$\tilde{s}(f) = \int df' \tilde{h}(f - f')\tilde{G}(f - f', f')$$

$$\tilde{G}(f, f') = \int dt e^{2i\pi f' t} e^{-2i\pi f d(t)} F(t)$$

Separation of timescales: if  $F, d$  have only frequencies  $\ll f$ , local convolution - expand  $h(f-f')$  in  $f'$

Convolution with frequency-dependent kernel

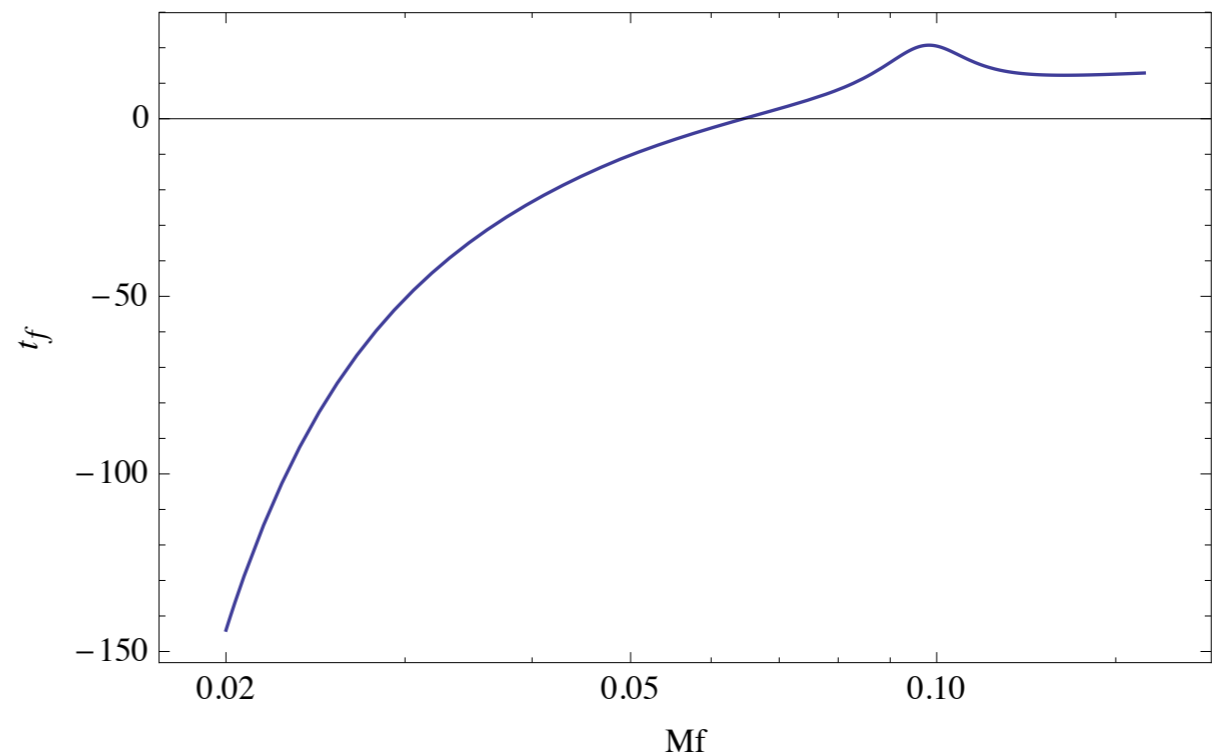
The leading order approximation

Keeping linear term in the phase:

$$t_f \equiv -\frac{1}{2\pi} \frac{d\Psi}{df}$$

$$\tilde{s}(f) = H(f)\tilde{h}(f)$$

$$H(f) = G(f, t_f)$$



# Delays and modulations in Fourier domain

## Higher-order corrections

$$\tilde{s}(f) = H(f)\tilde{h}(f) \quad \text{Leading order: } H(f) = G(f, t_f)$$

Phase (quadratic term):

$$H(f) = \sum \frac{1}{p!} \left( \frac{i}{8\pi^2} \frac{d^2\Psi}{df^2} \right)^p \partial_t^{2p} G(f, t_f)$$

Amplitude:

$$H(f) = \sum \frac{1}{(2i\pi)^p p!} \frac{1}{A} \frac{d^p A}{df^p} \partial_t^p G(f, t_f)$$

Frequency-dependence:

$$H(f) = \sum \frac{1}{(2i\pi)^p p!} \partial_f^p \partial_t^p G(f, t_f)$$

## Separation of timescales

(e)LISA:

$$\partial_t G \sim 2\pi f_0 G$$

$$f_0 = 1/\text{yr} = 3.10^{-8} \text{Hz} \ll f$$

Precessing binaries:  $G = F(t)$

Inspiral:  $\partial_t^2 F \sim \Omega_{\text{prec}}^2 \sim 2\text{PN}$

$$\frac{d^2\Psi}{df^2} \sim T_{\text{RR}}^2 \sim -2.5\text{PN} \quad + \text{Merger-ringdown ?}$$

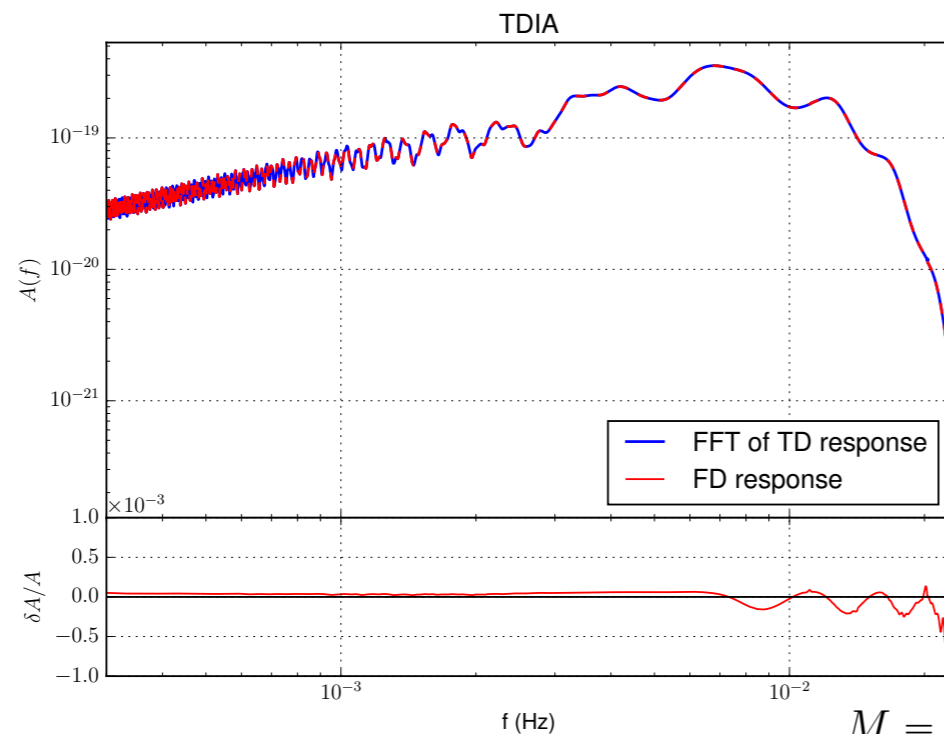
# Application to the (e)LISA response

## Magnitude of corrections

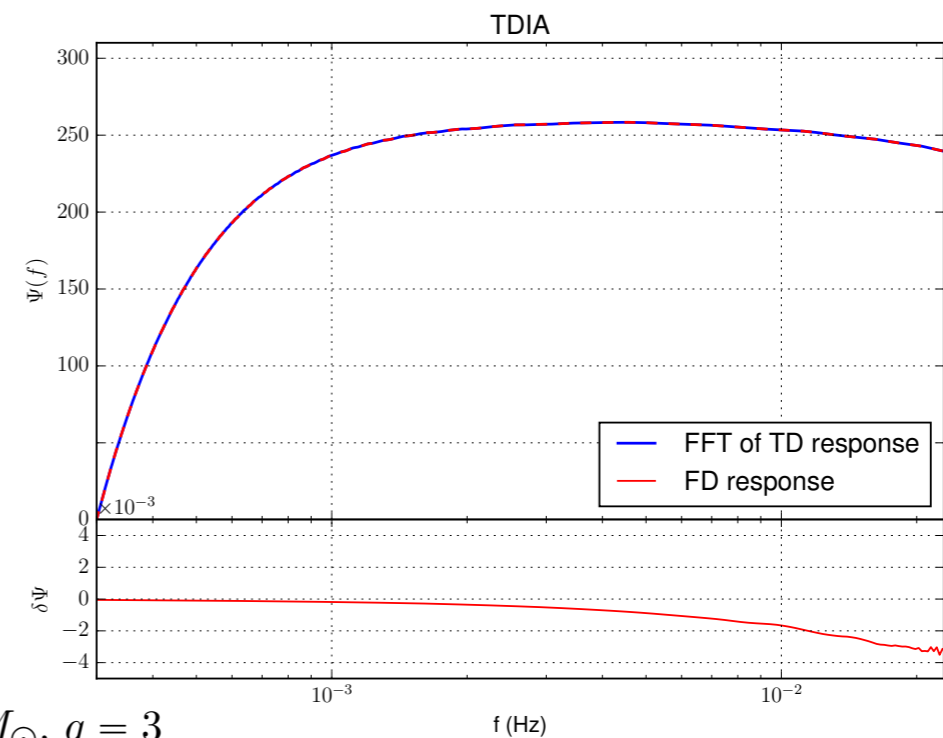
- Amplitude corrections  $< 10^{-3}$
- Phase corrections  $< 10^{-3}$
- High frequency:  $\partial_t \partial_f G \sim R^2 f f_0 / c^2 \sim 10^{-1}$  at 1Hz

## Example of errors

- TD response tested against SynthLISA
- Typical errors  $10^{-3}$



$$M = 2 \times 10^6 M_{\odot}, q = 3$$



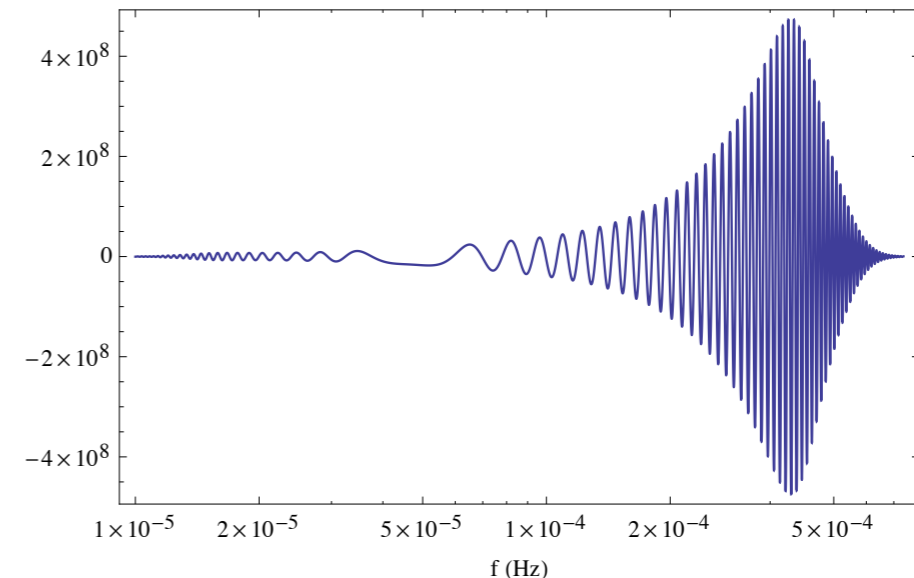
# Implementation

## Accelerated no-noise overlaps

- Amplitude/phase: splines on ~200 intervals
- Cost increases when including HM

$$(h_1|h_2) = 4\text{Re} \int df \frac{\tilde{h}_1(f)\tilde{h}_2^*(f)}{S_n(f)} \longrightarrow \int_{f_i}^{f_{i+1}} P(f) e^{i[af+bf^2]} \longrightarrow \int_{f_i}^{f_{i+1}} e^{i[af+bf^2]}$$

## Example of oscillatory integrand



## Performances

- Fourier-domain IMR sparse amplitude/phase waveforms (ROM) (w/o 21,33,44,55 modes)
- Accelerated overlaps for amplitude/phase (no noise)

Likelihood cost	Single mode	5 modes
Lin. overlaps	~10ms	~50ms
Acc. overlaps	~1ms	~10ms

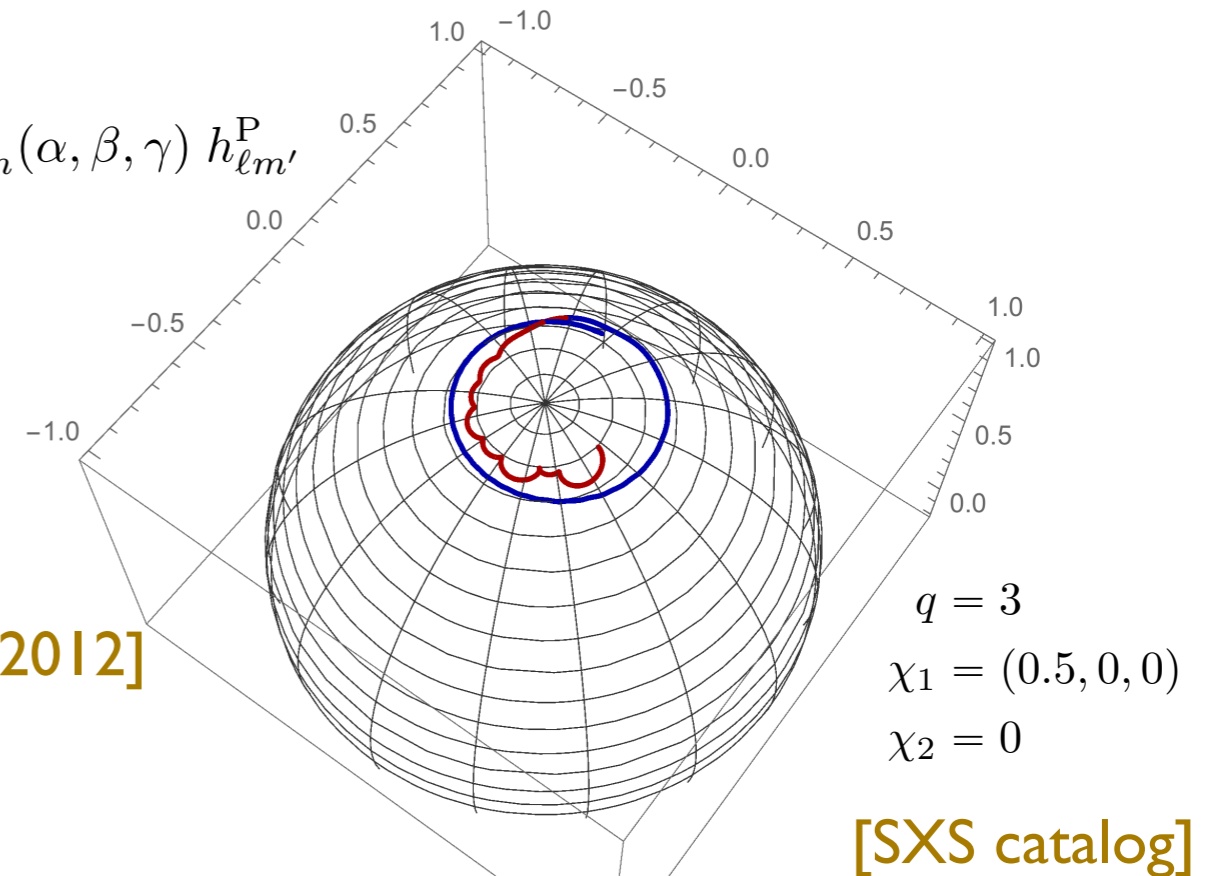
# Precessing waveforms: frame evolution

## Frame trajectory

$$h_{\ell m}^I = \sum_{m'} D_{m' m}^{\ell*}(\alpha, \beta, \gamma) h_{\ell m'}^P$$

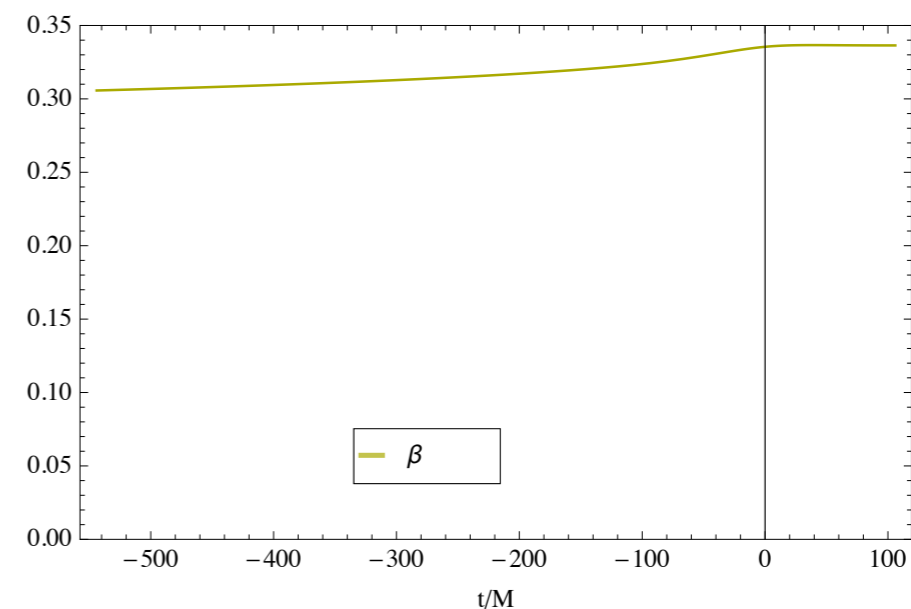
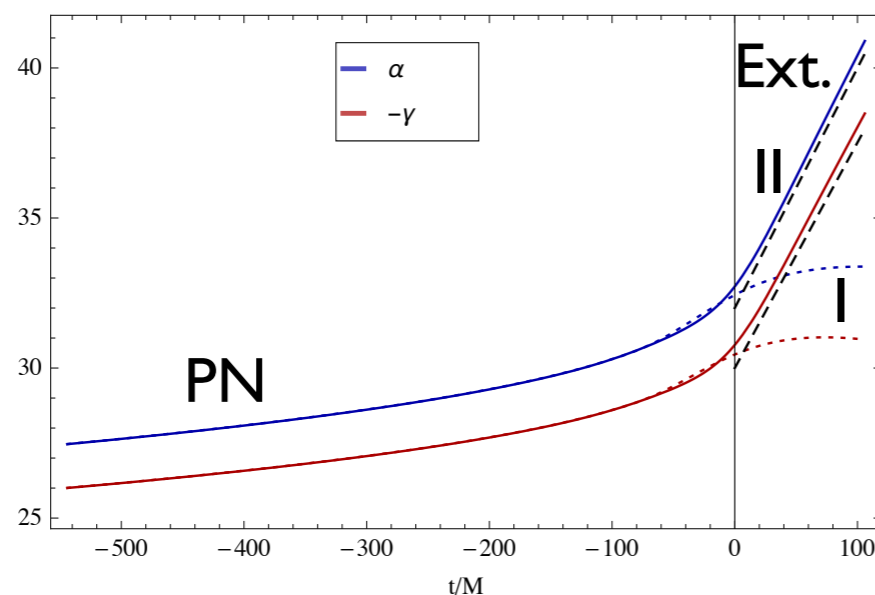
- PN dynamics:  $Z_{\text{frame}} = \hat{L}$
- Extracting frame from the waveform (IMR)  
[O'Shaughnessy&al 2011]
- Approximate behaviour post-merger:

$$\Omega_{\text{frame}} \sim \omega_{220}^{\text{QNM}} - \omega_{210}^{\text{QNM}} \quad [\text{O'Shaughnessy\&al 2012}]$$



## Pre- and post-merger frame: toy model

[Smoothness assumption]



# Relation to previous works

- Previous works:
- Leading order (different MR) [SpinTaylorF2, PhenomP]
  - Quadratic phase (SUA) [Klein&al 2014]

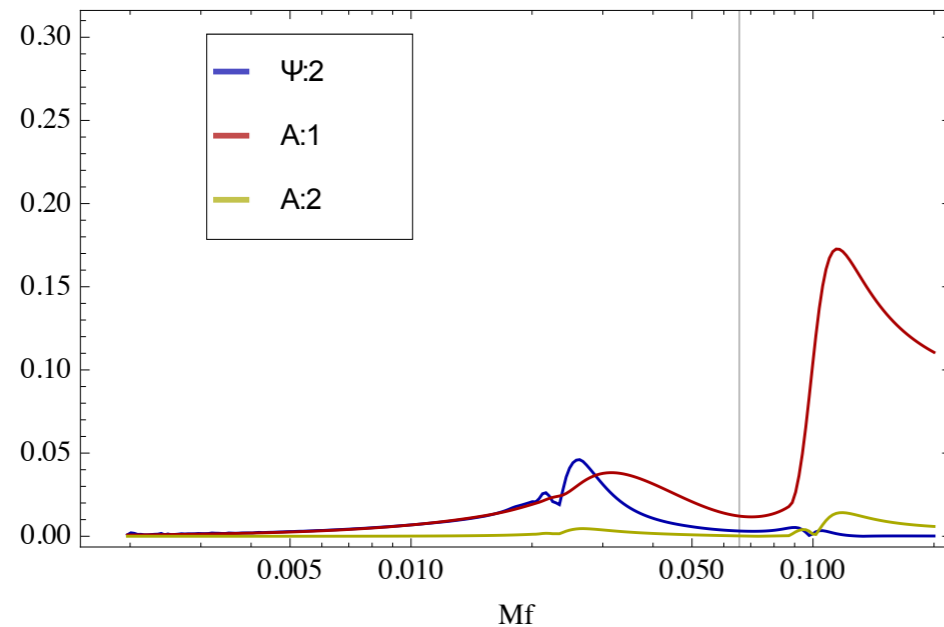
SPA/SUA	Fourier domain approach
$t_f : \omega(t_f) = \pi f \quad (\text{SPA})$	$t_f = -\frac{1}{2\pi} \frac{d\Psi}{df} \quad (\text{IMR})$
$T_f = \frac{1}{\sqrt{2\dot{\omega}(t_f)}} \quad \text{Rad. Reac. (SUA)}$	$T_f^2 = \frac{1}{4\pi^2} \left  \frac{d^2\Psi}{df^2} \right  \quad (\text{IMR})$
$\tilde{s}(f) = \tilde{h}(f) \sum \frac{(-i)^p}{2^p p!} T_f^{2p} \partial_t^{2p} F \quad (\text{SUA})$	$\tilde{s}(f) = \tilde{h}(f) \sum \frac{(-i)^p}{2^p p!} T_f^{2p} \partial_t^{2p} F \quad \begin{array}{l} \text{Taylor FD} \\ \text{Quad. phase} \end{array}$
$\tilde{s}(f) = \tilde{h}(f) \sum a_k F(t_f \pm kT_f) \quad (\text{Resum.})$	$\tilde{s}(f) = \tilde{h}(f) \int dt \exp \left[ -\frac{i}{2} \left( \frac{t - t_f}{T_f} \right)^2 \right] F(t)$

- New corrections:
- Higher-order amplitude corrections  $d^p A/df^p$
  - Local convolution approach for post-merger

# Precession: magnitude of corrections

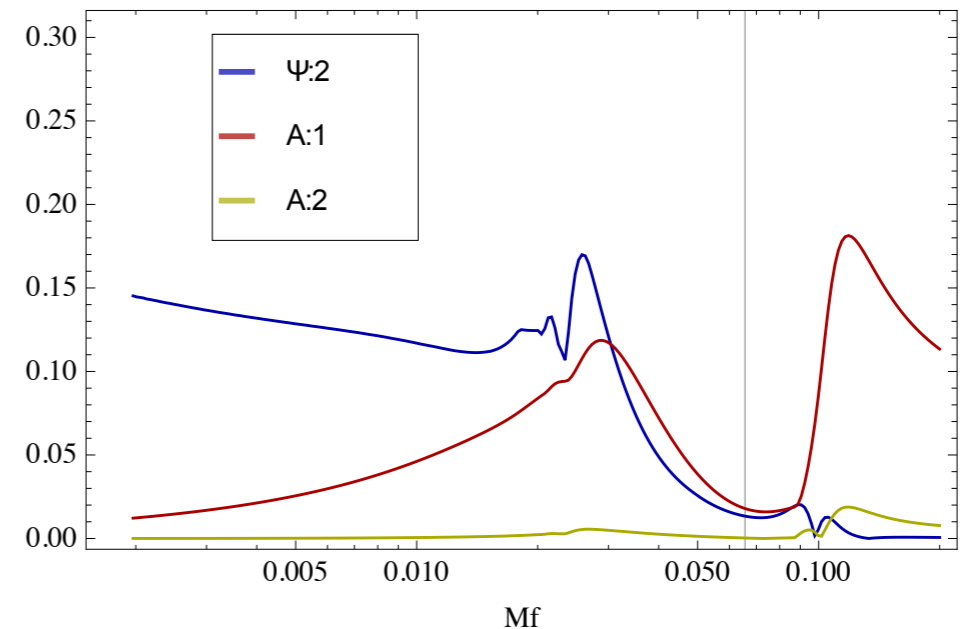
Example:  $q = 3$ ,  $\chi_1 = (-0.3, 0.5, 0.7)$ ,  $\chi_2 = (0.3, -0.2, -0.5)$

$$h_{22}^P \rightarrow h_{22}^I$$

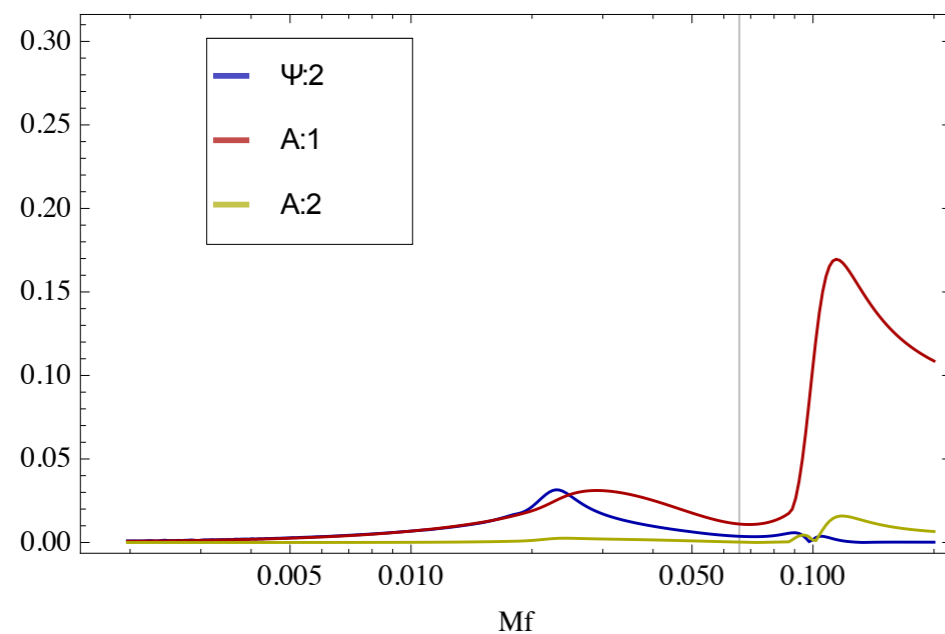


Case I

$$h_{22}^P \rightarrow h_{21}^I$$

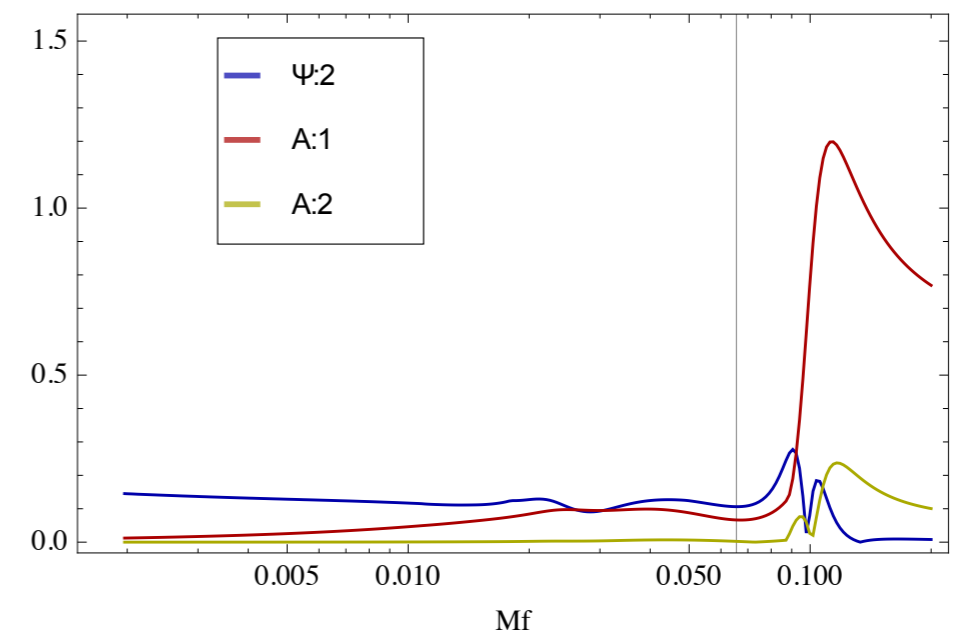


$$h_{22}^P \rightarrow h_{22}^I$$



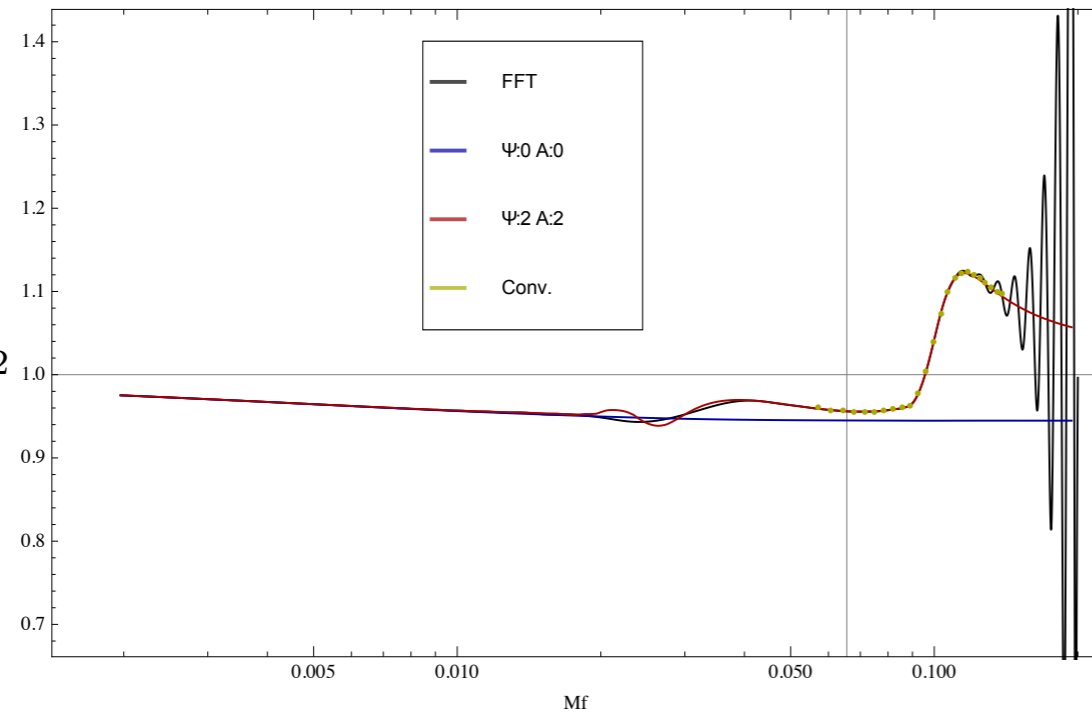
Case II

$$h_{22}^P \rightarrow h_{21}^I$$

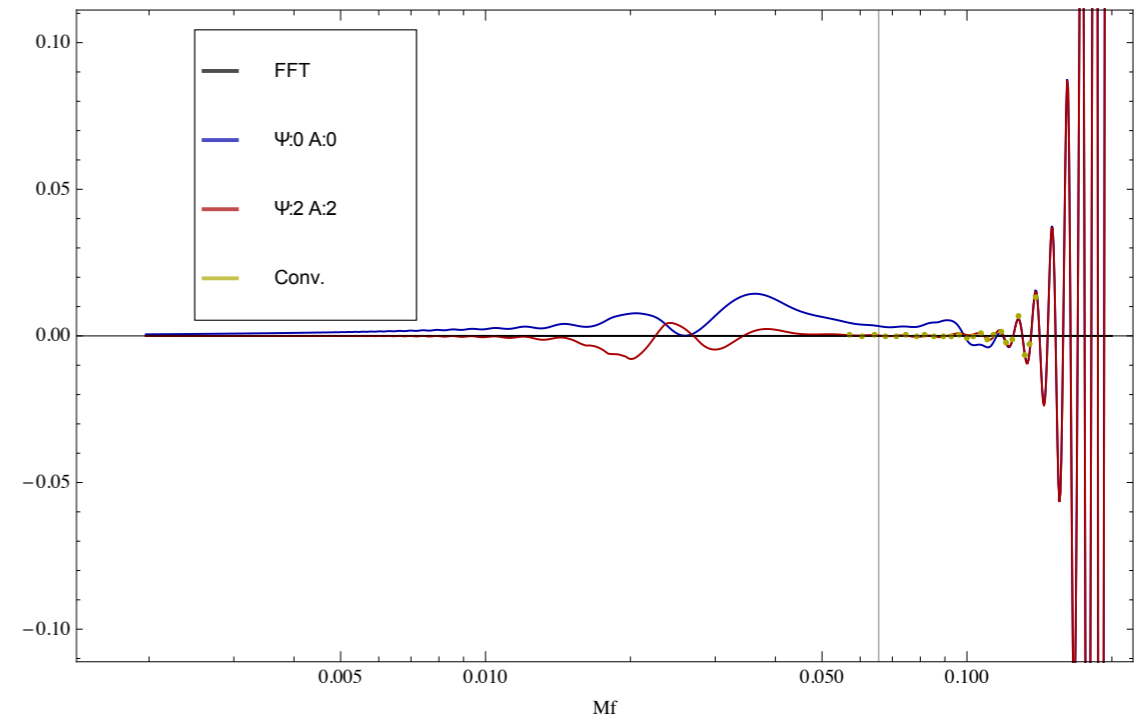


# Precession: errors - case I

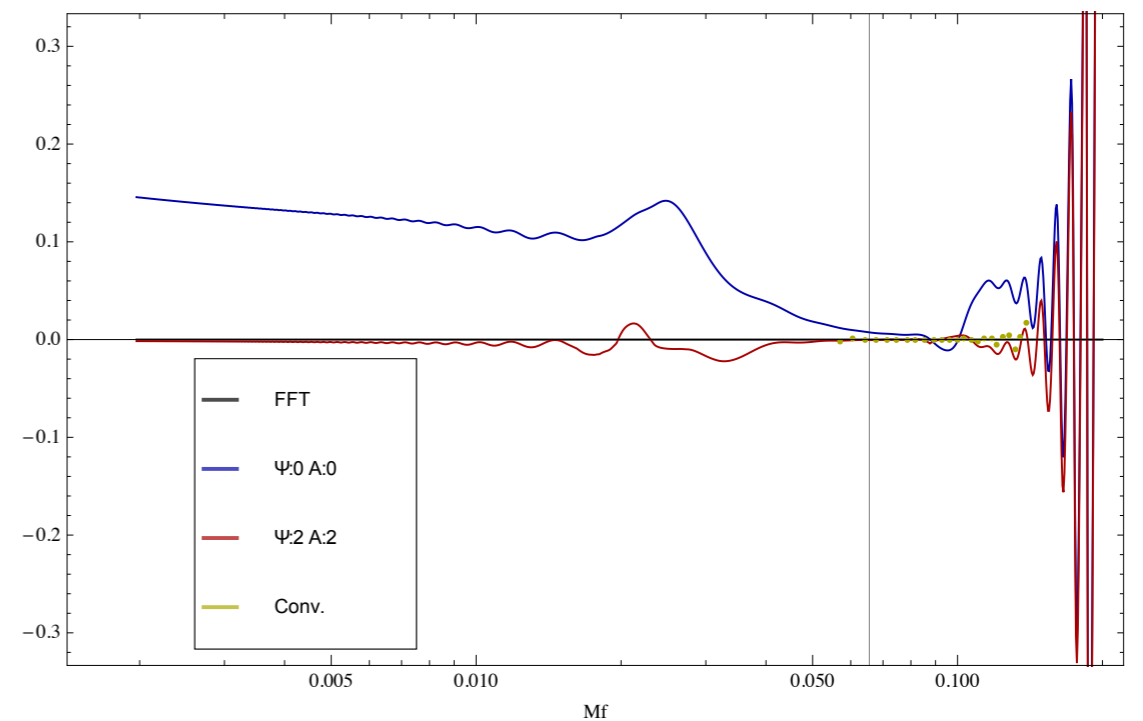
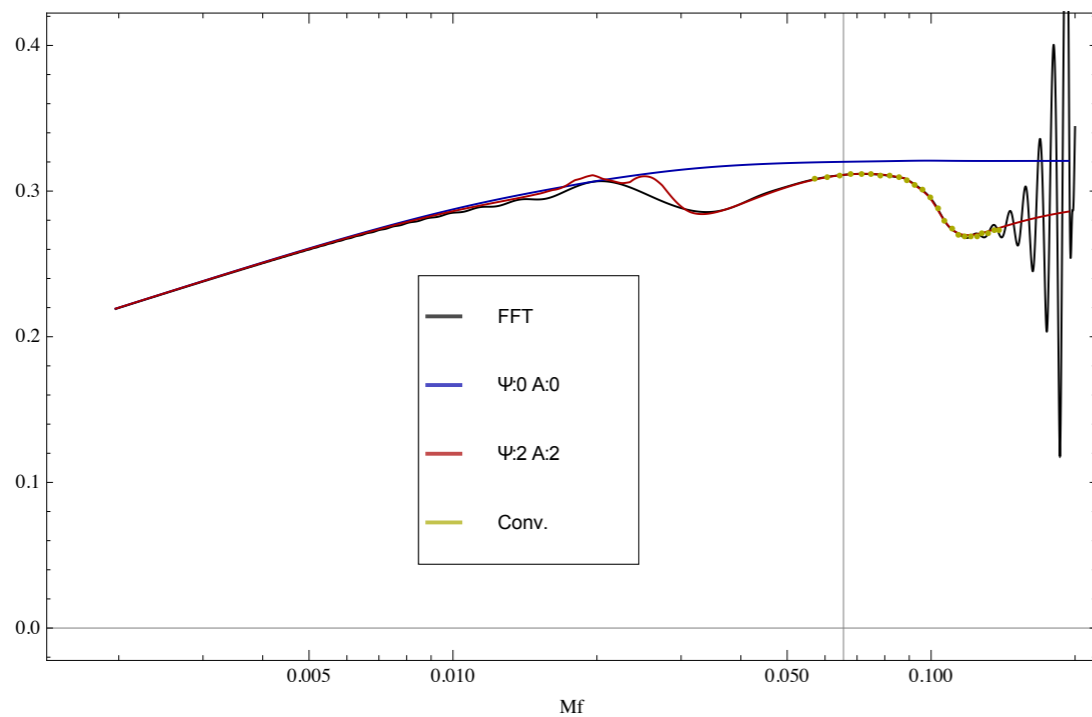
Amplitude relative to h22P



Phase difference

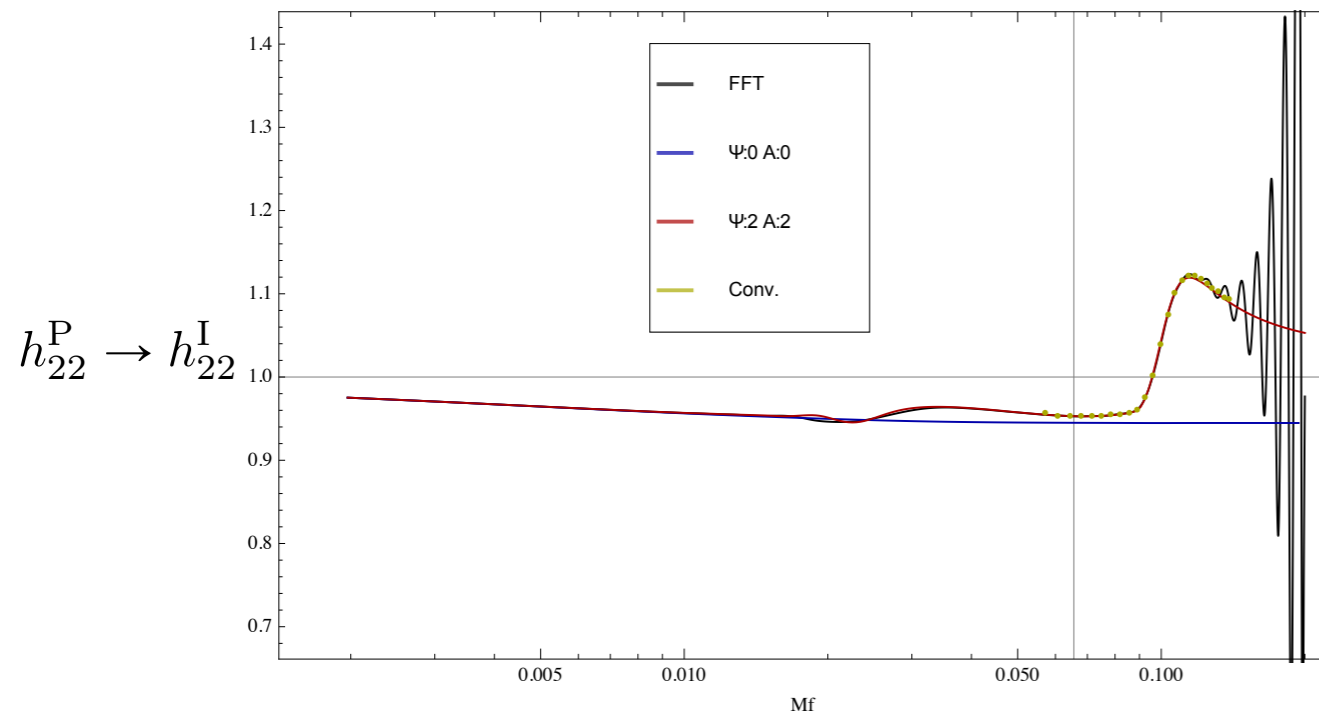


$h_{22}^P \rightarrow h_{21}^I$

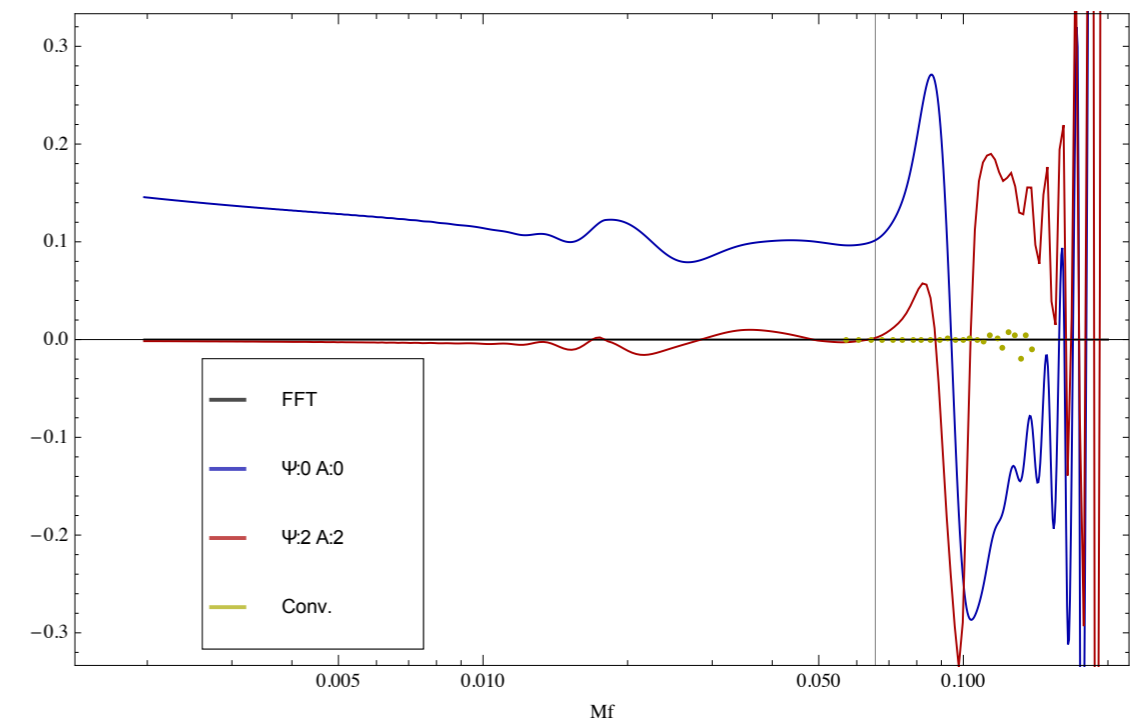
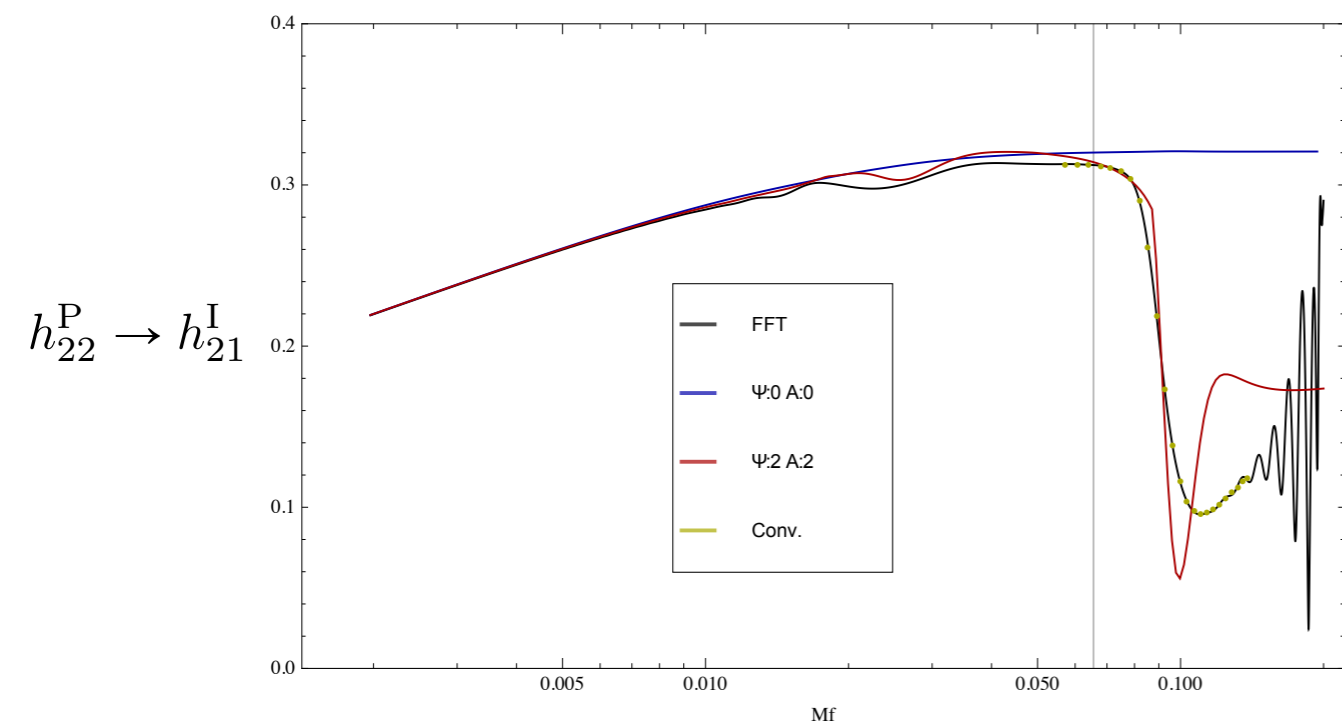
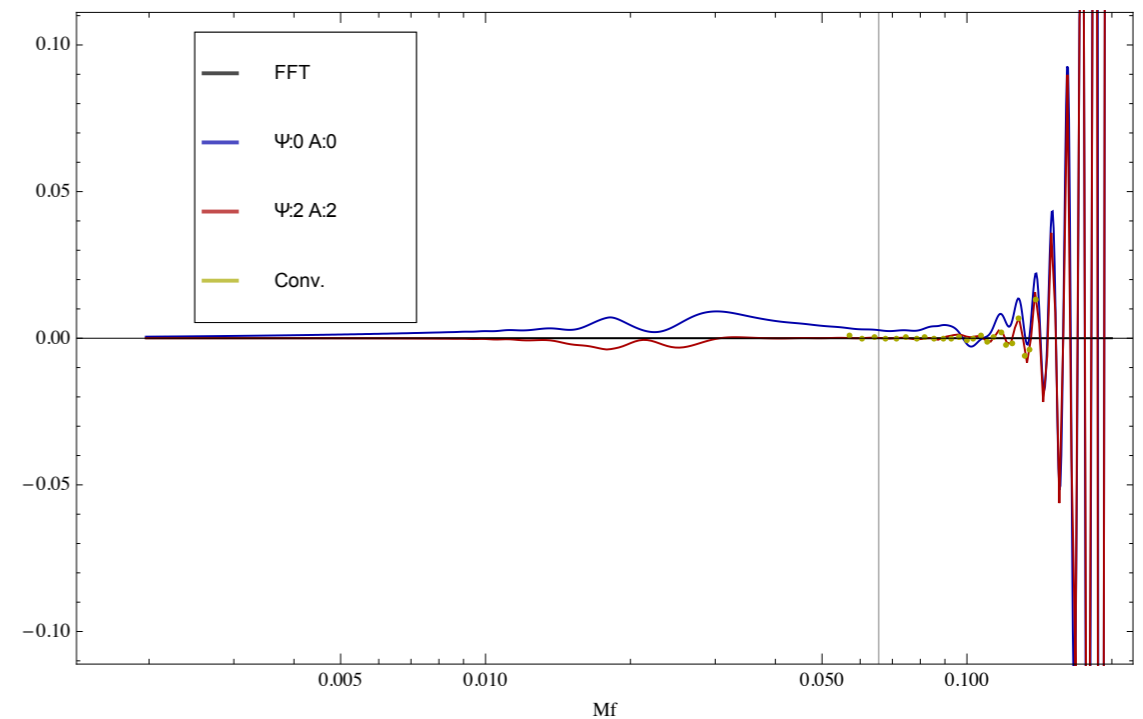


# Precession: errors - case II

Amplitude relative to h22P



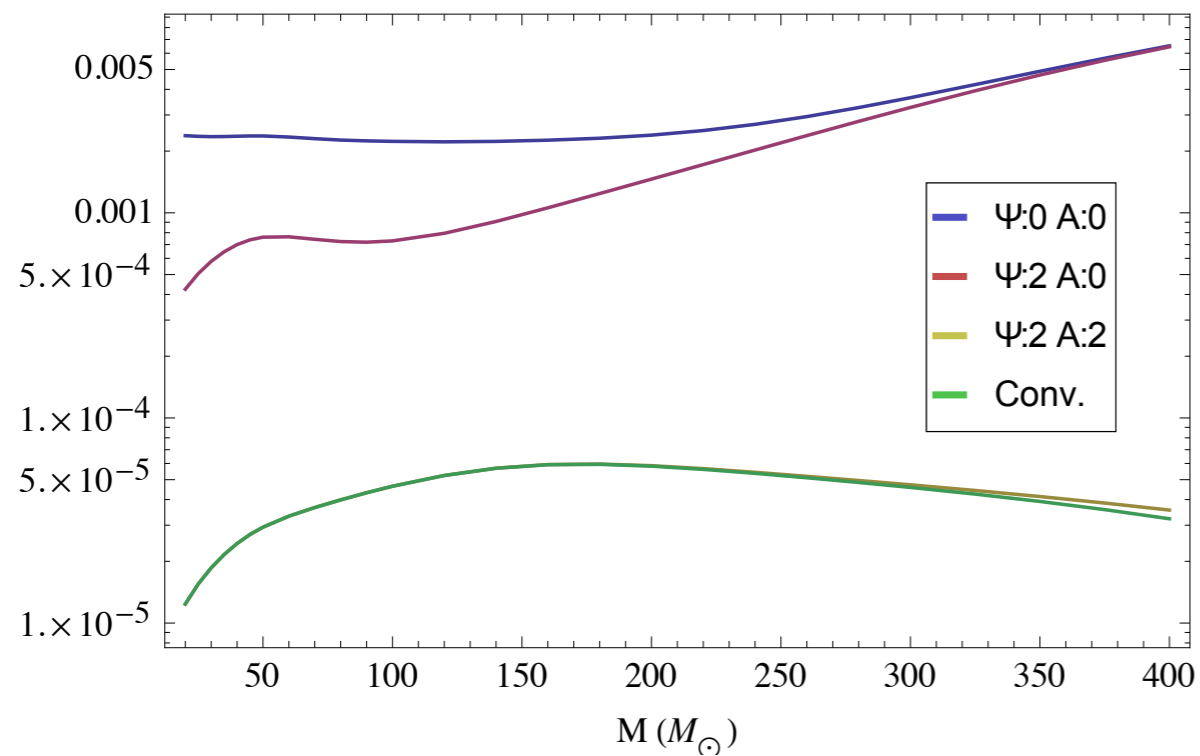
Phase difference



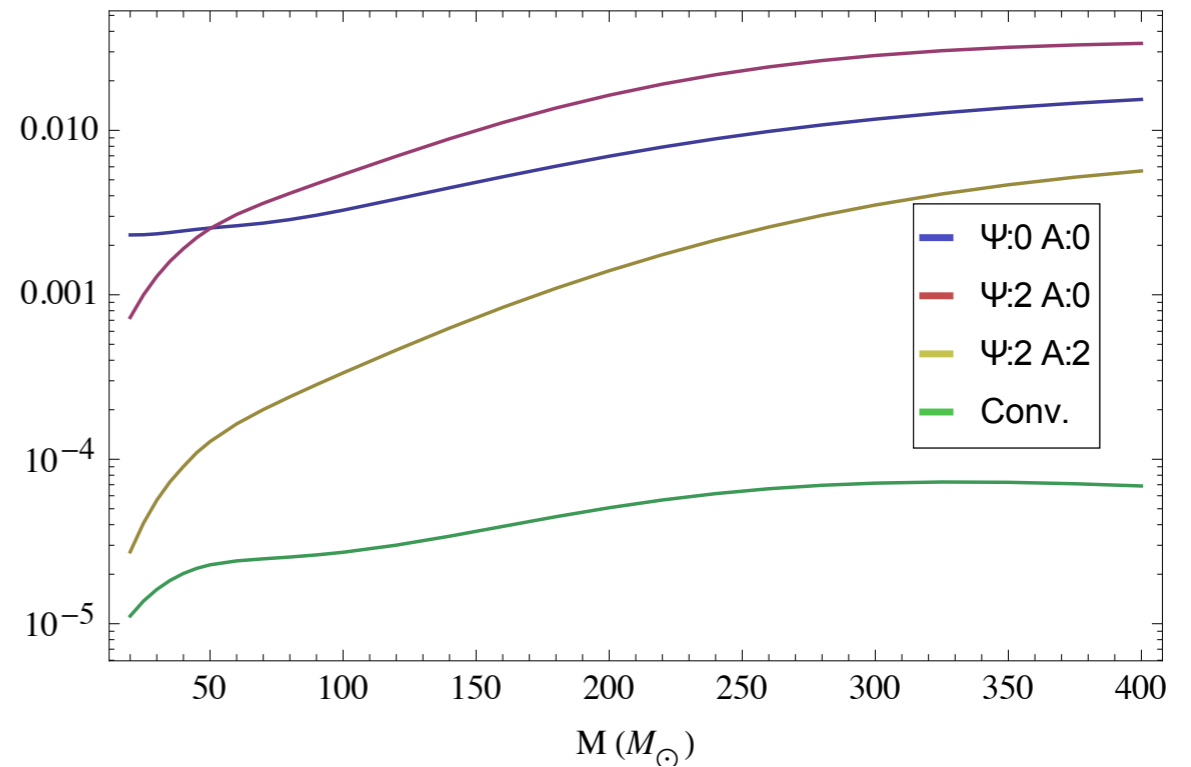
# Precession: mismatches

[Preliminary]  
[Broader exploration of parameter space needed]

Case I



Case II



$$\iota = \pi/2$$

# Summary

## (e)LISA prospective parameter estimation

- Fourier-domain processing through the response of the instrument using  $t(f)$  correspondence
- Higher-order corrections available
- Applicable to all FD IMR waveform models: aligned spins (SEOBNRv2, PhenomD), precession (PhenomP), compact FD amplitude/phase representation
- Implementation using accelerated no-noise overlaps: few ms/likelihood
- Ongoing and future exploration: impact of MR, HM, spins and of instrument design

## Modeling waveforms from precessing binaries

- Formally recover and extend previous results on FD frame rotation
- Inclusion of FD amplitude corrections, direct convolution approach for post-merger
- ‘Mild’ frame rotation: captured by amplitude corrections
- ‘Fast’ frame rotation: requires direct convolution
- New corrections are local to MR, and mostly in amplitude — small mismatches !
- More systematic exploration of parameter space is needed



# Example: sky position degeneracies

SNR 20

Parameters:

(artificially distant)

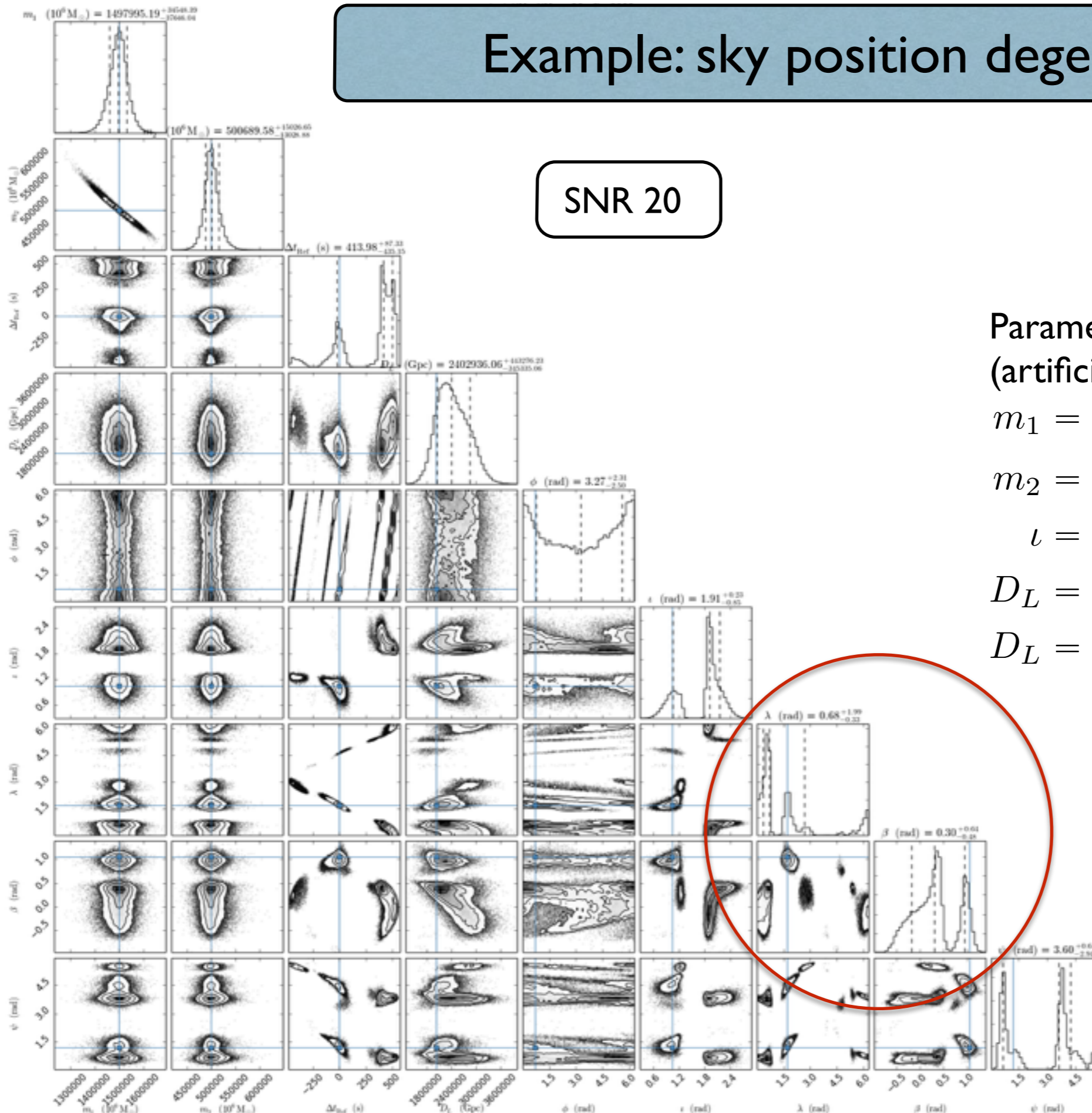
$$m_1 = 1.5 \times 10^6 M_\odot$$

$$m_2 = 0.5 \times 10^6 M_\odot$$

$$\iota = \pi/3$$

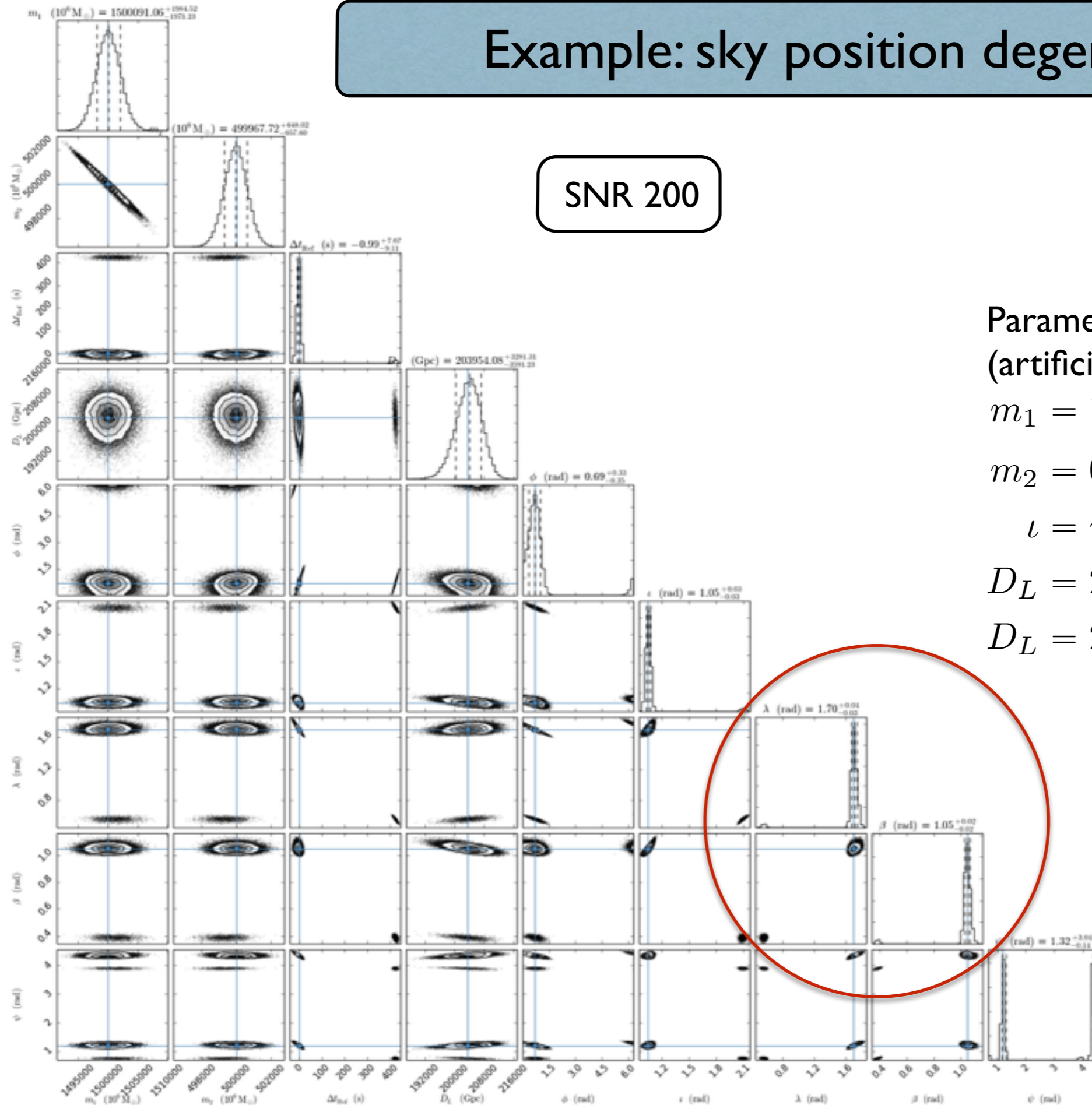
$$D_L = 2036 \text{ Gpc (SNR 20)}$$

$$D_L = 203.6 \text{ Gpc (SNR 200)}$$



# Example: sky position degeneracies

SNR 200



Parameters:

(artificially distant)

$$m_1 = 1.5 \times 10^6 M_\odot$$

$$m_2 = 0.5 \times 10^6 M_\odot$$

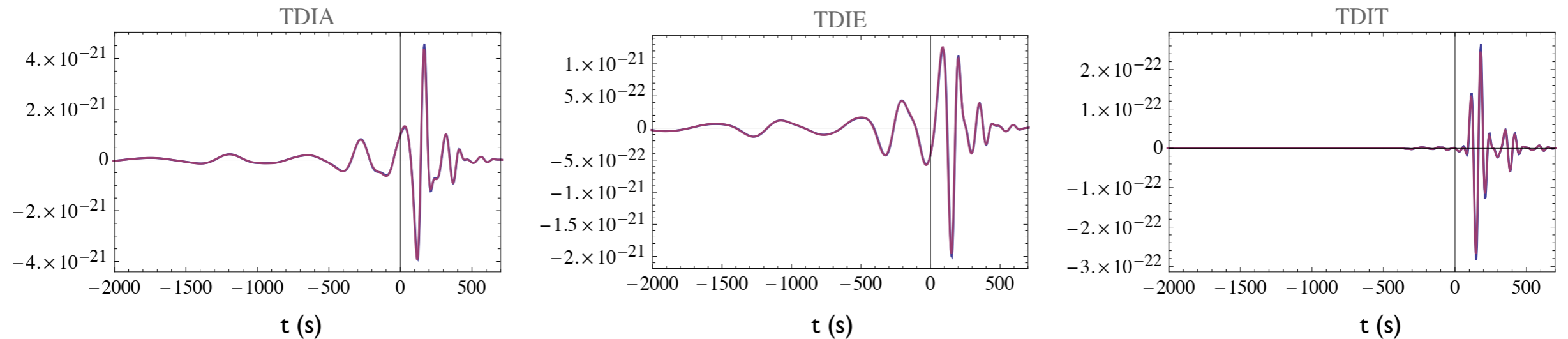
$$\iota = \pi/3$$

$$D_L = 2036 \text{ Gpc (SNR 20)}$$

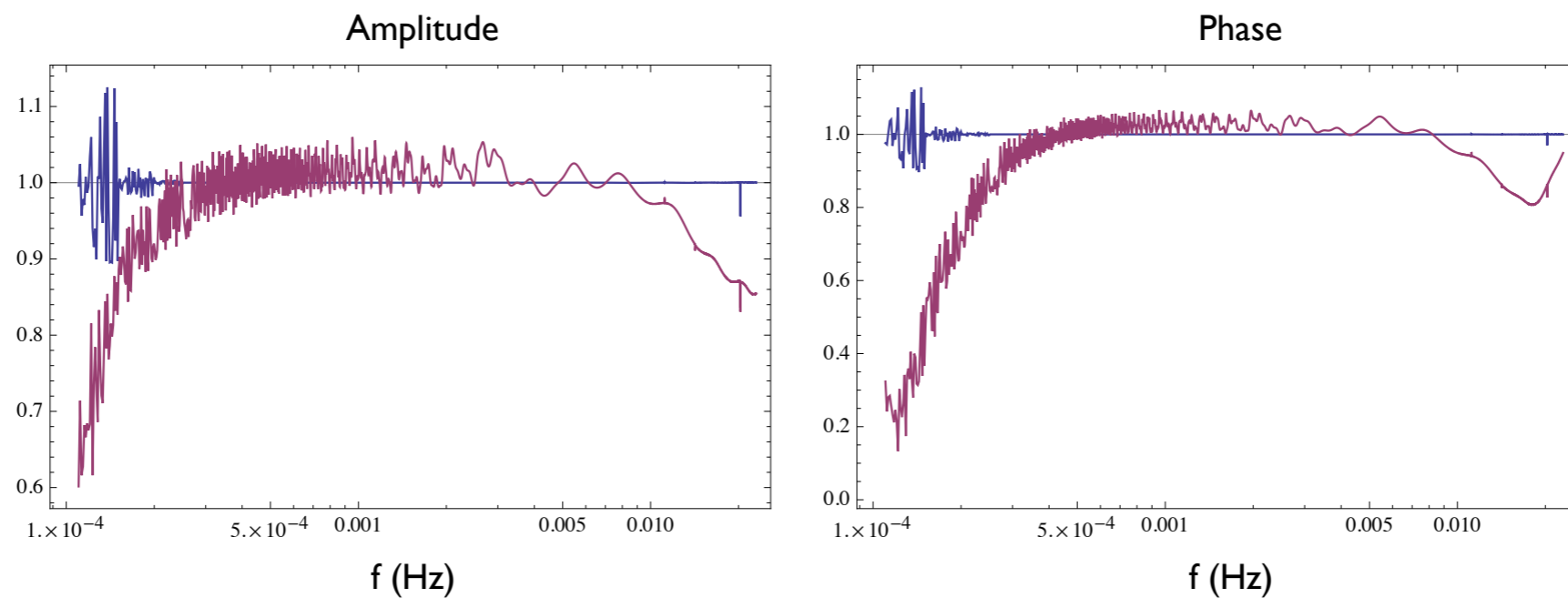
$$D_L = 203.6 \text{ Gpc (SNR 200)}$$

# Example: sky position degeneracies

## TD comparison to 2nd peak

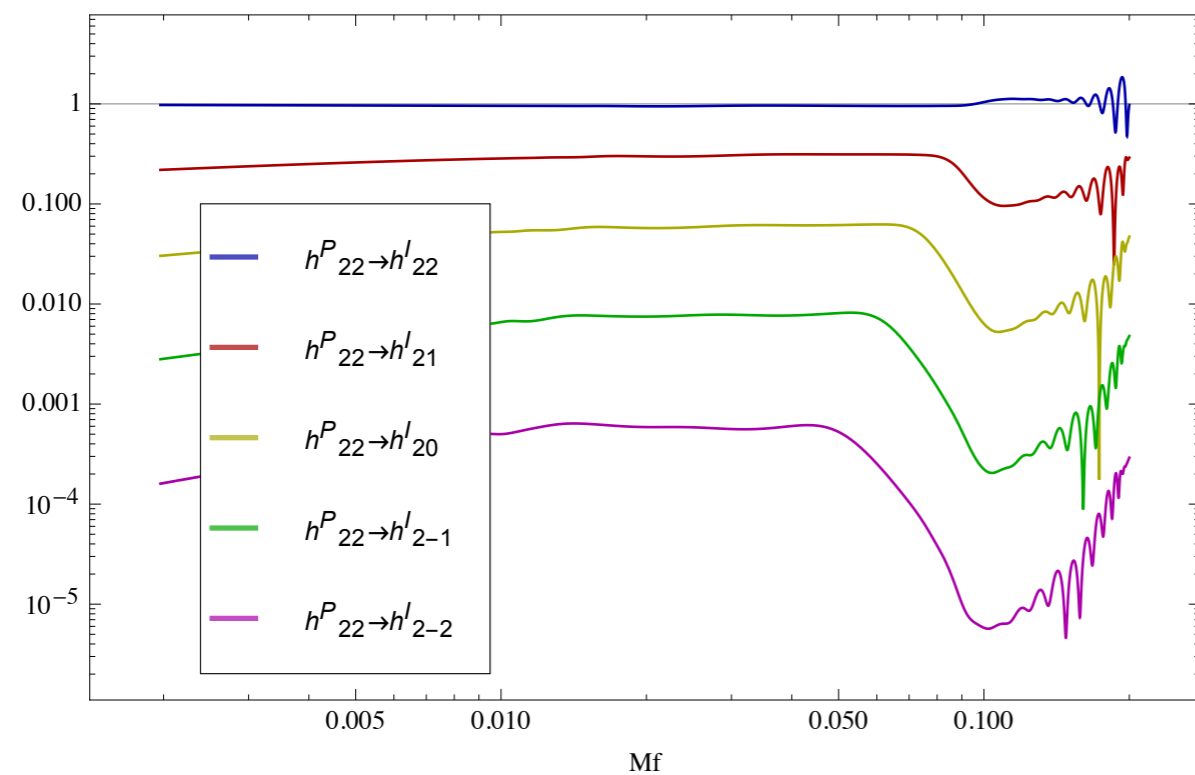


## FD comparison to 2nd peak

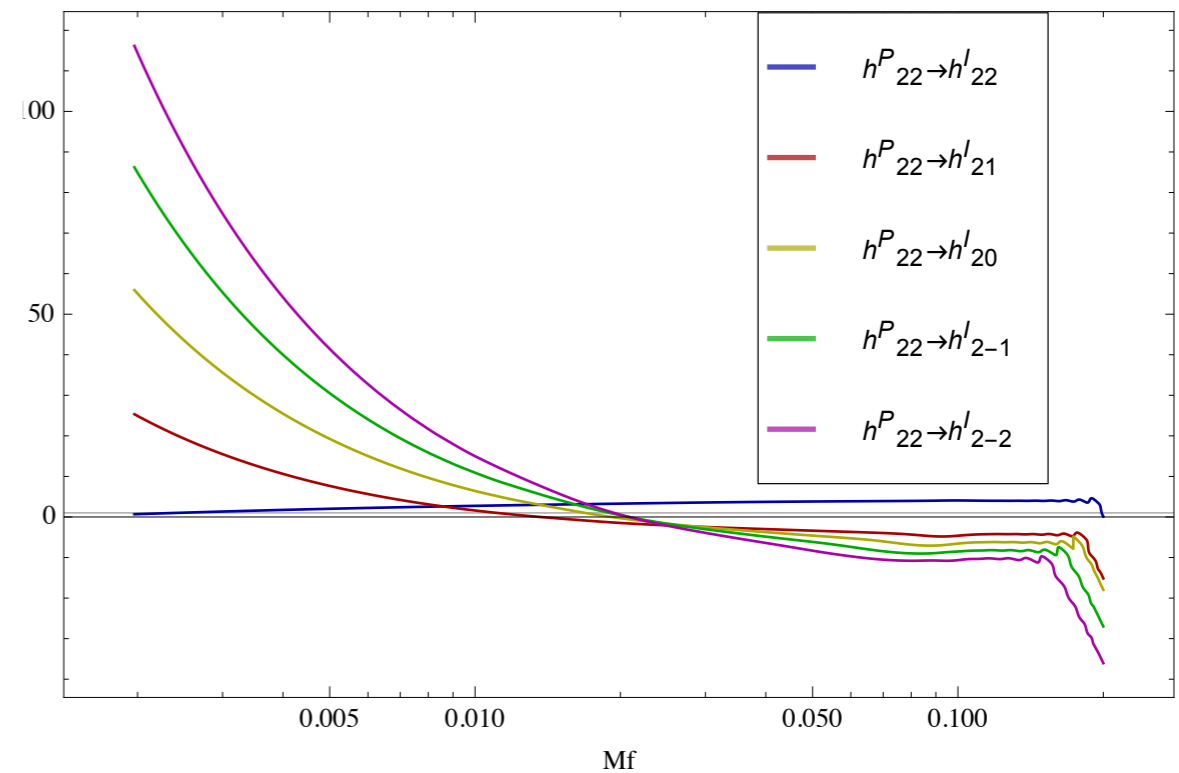


# FD transfer functions for different modes

## Normalized amplitude

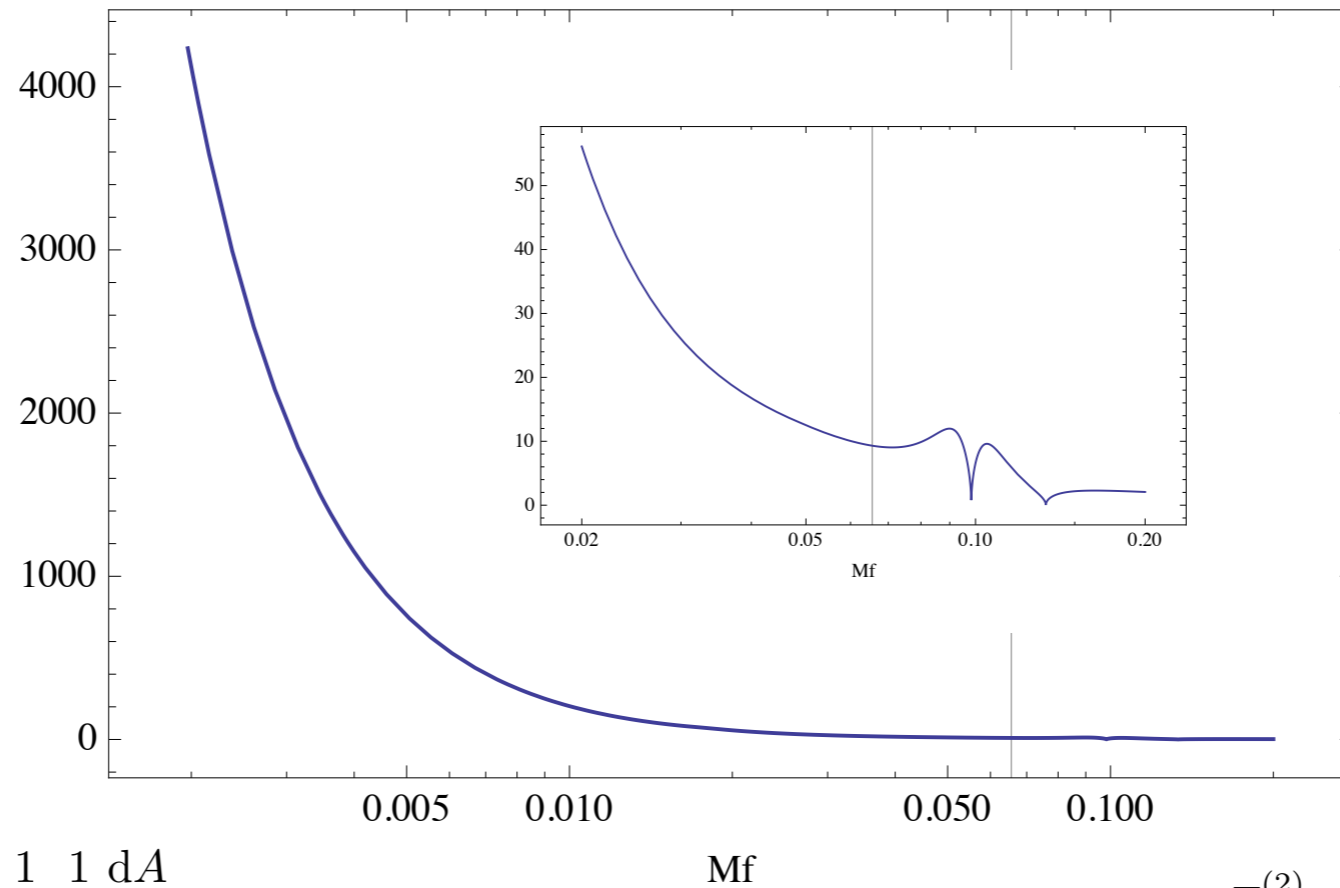


## Phase

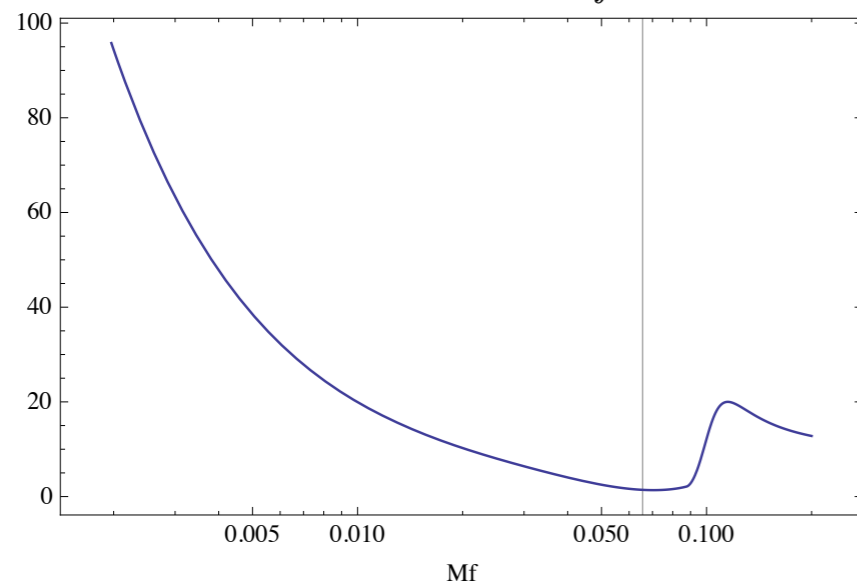


# FD timescales

$$T_f = \sqrt{\frac{1}{4\pi^2} \left| \frac{d^2\Psi}{df^2} \right|}$$



$$T_A^{(1)} = \frac{1}{2\pi} \frac{1}{A} \frac{dA}{df}$$



$$T_A^{(2)} = \sqrt{\frac{1}{4\pi^2} \frac{1}{A} \frac{d^2A}{df^2}}$$

