

Using the GCFE to study the “ringing” of a Schwarzschild black hole

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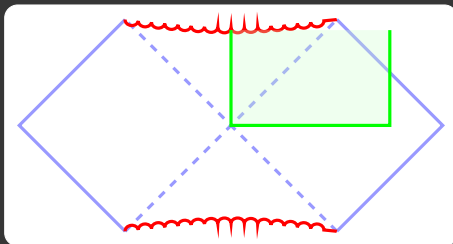
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Conformal field equations

- ▶ implement Penrose's **conformal approach**, developed by H. Friedrich
- ▶ system of equations equivalent to standard Einstein equations
- ▶ determine a space-time in a **conformally covariant** way
- ▶ equations for
 - ▶ frame
 - ▶ (Weyl) connection
 - ▶ Ricci tensor
 - ▶ Weyl tensor
- ▶ conformal Gauß gauge:
 - ▶ based on congruence of time-like **conformal geodesics**
 - ▶ yields mostly advection equations along conformal geodesics
 - ▶ spin-2 equation for Weyl tensor is **symmetric hyperbolic**

Setup

Schematic diagram of computational domain



- ▶ Schwarzschild initial data (note, not spatially compactified)
- ▶ left boundary inside horizon
 - ▶ isometry condition
- ▶ finite outer boundary
 - ▶ data with $\Psi_0 \neq 0$, (axi-symmetrically) 'distorted' black hole
- ▶ reach the singularity and null-infinity on the same grid
- ▶ read off Ψ_4 on \mathcal{I} and compute the Bondi mass

Results and issues

Results

- ▶ constraints propagate (numerical observation)
- ▶ IBVP is (numerically) well-posed
- ▶ reproduce Schwarzschild (and S-dS and S-AdS)
- ▶ distorted black hole loses energy, Bondi mass loss

Issues

- ▶ $t = \text{const}$ hyper-surfaces tilt upwards, become null outside \mathcal{I}
- ▶ variable number of incoming modes at the outer boundary
- ▶ 'ringing' not (yet) seen
 - ▶ lack of resolution (?)
 - ▶ magnitude (?)
 - ▶ inner boundary condition (?)
- ▶ work in progress

Outlook

- ▶ approach seems to be ideal to study **global properties** of space-times
- ▶ currently work on improving the code to run higher resolutions
- ▶ can be used also when initial data surface is compactified
- ▶ allows the setup of a **scattering problem**
 - ▶ ingoing wave data given on \mathcal{I}^- read
 - ▶ outgoing wave data read off on \mathcal{I}^+
- ▶ should also be applicable to **Kerr black holes**