

# Second-order self-force: a progress report

Adam Pound

in collaboration with Wardell, Warburton, Miller, Barack; Moxon,  
Flanagan, Hinderer; Yamada, Tanaka, Isoyama

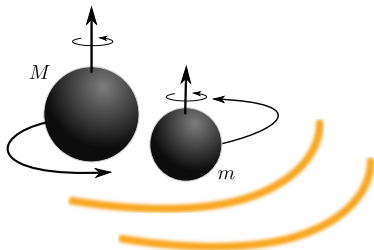
University of Southampton

14 July 2016

# Outline

- 1 Why second-order self-force?
- 2 Self-force theory: replacing an extended object with a puncture
- 3 First application: quasicircular orbits in Schwarzschild

# Gravitational waves and comparable-mass binaries

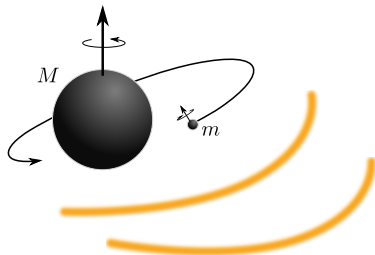


- ground-based detectors will continue to observe such systems
- but they will not be able to observe binaries with very unequal masses

- first directly detected waves were generated by inspiral and merger of two comparable-mass black holes

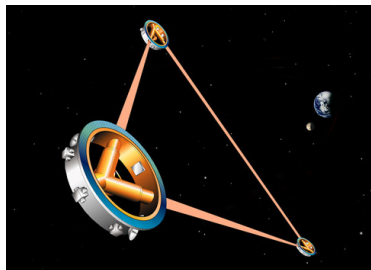


# Extreme-mass-ratio inspirals (EMRIs)

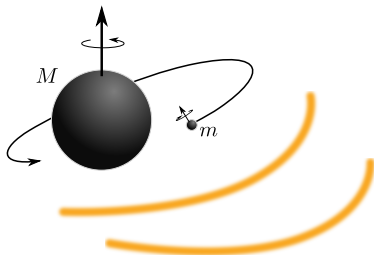


- small object spends  $\sim M/m \sim 10^5$  orbits near BH  
 $\Rightarrow$  unparalleled probe of strong-field region around BH
- measure deviations from GR (with  $\gtrsim 10$  times precision of other probes)

- space-based detector eLISA will observe extreme-mass ratio inspirals of stellar-mass BHs or neutron stars into massive BHs



# Modeling EMRIs: why second-order self-force?



- highly relativistic, strong fields
- disparate lengthscales
- long timescale: inspiral is slow, produces  $\sim \frac{M}{m} \sim 10^5$  wave cycles

- treat  $m$  as source of perturbation of  $M$ 's metric  $g_{\mu\nu}$ :

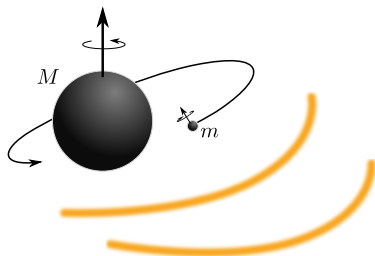
$$g_{\mu\nu} = g_{\mu\nu} + \epsilon h_{\mu\nu}^1 + \epsilon^2 h_{\mu\nu}^2 + \dots$$

- represent motion of  $m$  via worldline  $z^\mu$  satisfying

$$\frac{D^2 z^\mu}{d\tau^2} = \epsilon F_1^\mu + \epsilon^2 F_2^\mu + \dots$$

- scaling arguments show we need  $F_2^\mu$  for accurate model

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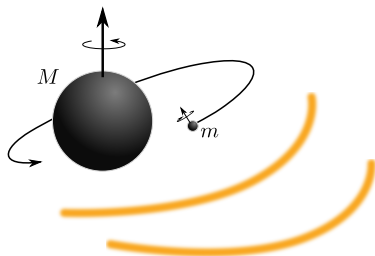
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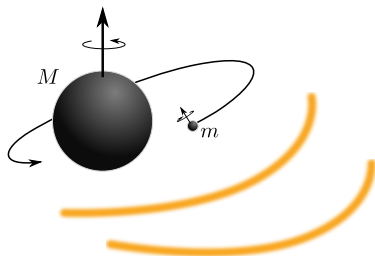
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 $\Rightarrow$  *need a model that is accurate over those  $\sim 10^5$  cycles*

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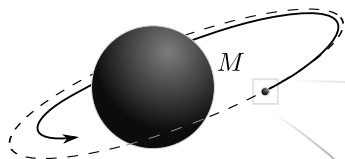
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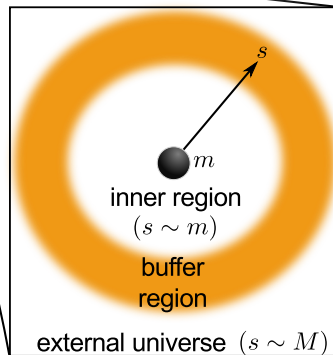
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# Matched asymptotic expansions

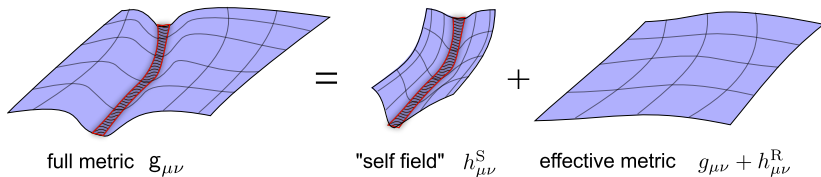


- equation of motion and form of field near object found by local analysis in buffer region



# Punctures and EOM [Detweiler & Whiting; Barack et al; AP; Gralla; Harte]

- split field into “self-field” and “effective field”
- locally replace self-field with singular field

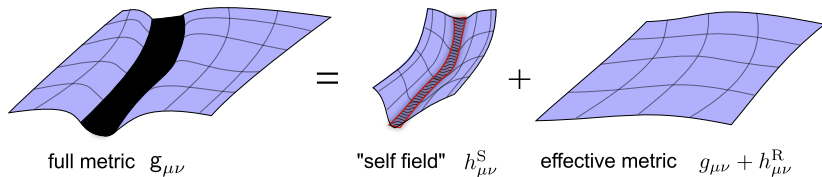


- replaces object with a *puncture*, a local singularity in the field, moving on  $z^\mu$ , equipped with the object's multipole moments
- $z^\mu$  geodesic in  $g_{\mu\nu} + h_{\mu\nu}^R$ :

$$\frac{D^2 z^\mu}{d\tau^2} = -\frac{1}{2} (g^{\mu\nu} + u^\mu u^\nu) (g_\nu{}^\rho - h_\nu{}^\rho) (2h_{\rho\sigma;\lambda}^R - h_{\sigma\lambda;\rho}^R) u^\sigma u^\lambda + O(m^3)$$

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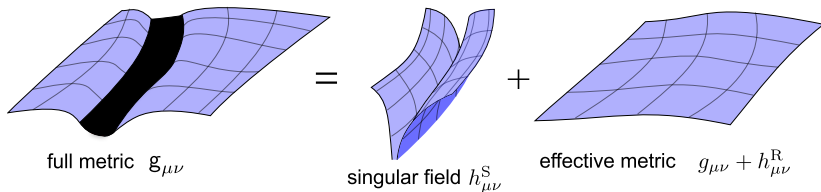


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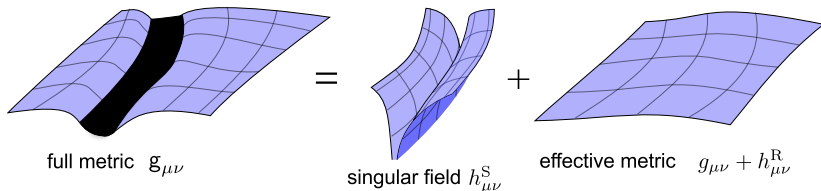


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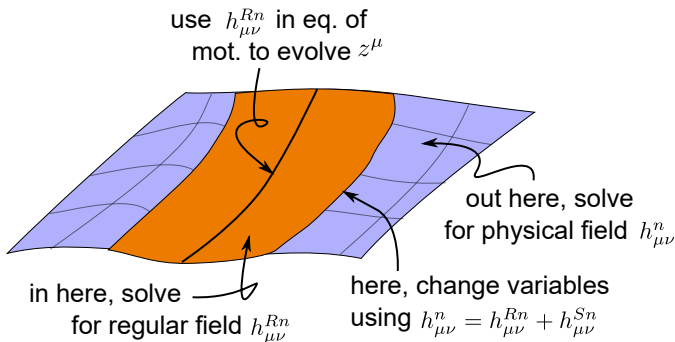


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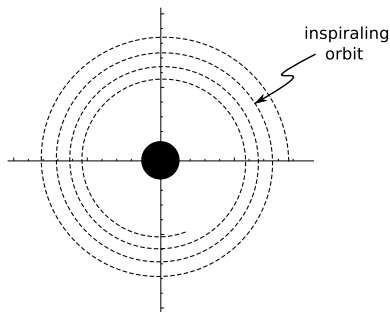
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# How you replace an object with a worldline

- in region near  $z^\mu$ , move  $h_{\mu\nu}^S$  to RHS of field equation, solve for  $h_{\mu\nu}^R$



# Quasicircular orbits [AP, Wardell, Warburton, Miller, Barack]



Introduce slow time  $\tilde{t} \sim \epsilon t$

- multiscale expansion of the worldline:

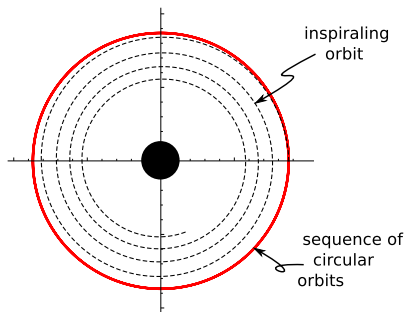
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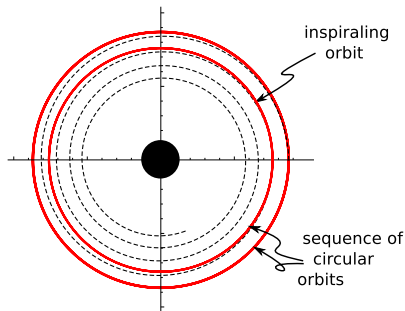
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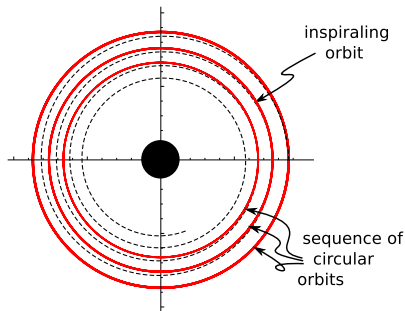
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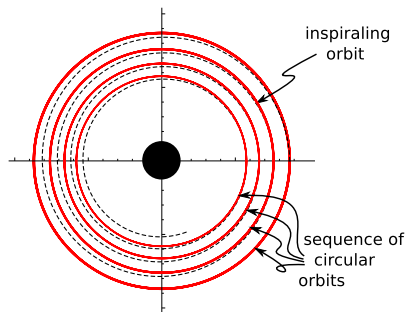
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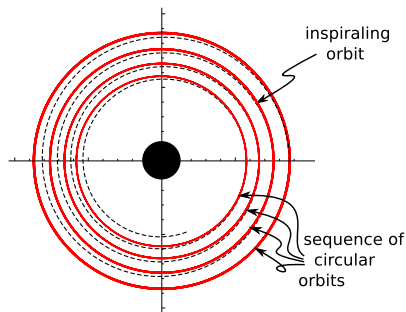
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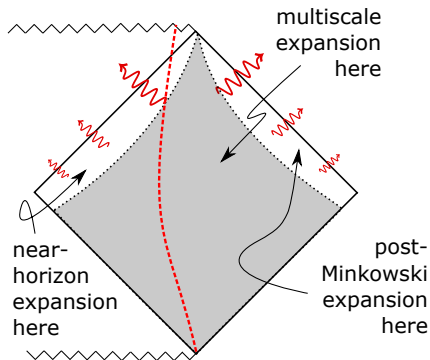
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# Matched expansions [AP,Moxon,Flanagan,Hinderer,Yamada,Isoyama,Tanaka]



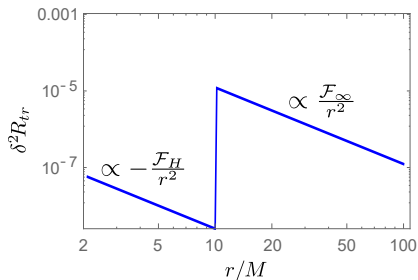
- multiscale expansion works on large scales, but not globally
- so use different expansions in different regions

- post-Minkowski and near-horizon expansions provide BCs  $h_{\mu\nu}^{\infty}$  and  $h_{\mu\nu}^{\mathcal{H}}$  for multiscale expansion
- we subtract the BCs from the field, numerically solve ODEs for residual  $h_{\mu\nu}^{\mathcal{R}} = h_{\mu\nu} - h_{\mu\nu}^S - h_{\mu\nu}^{\infty} - h_{\mu\nu}^{\mathcal{H}}$ ,

$\ell = 0$ , dissipative sector

- field equation:

$$\begin{aligned}\partial_r^2 h_{tr}^{\mathcal{R}2} &\sim \delta^2 R_{tr}[h^1, h^1] \\ &\quad - \partial_r^2 h_{tr}^{S2} \\ &\quad - \partial_t h_{tt}^1\end{aligned}$$



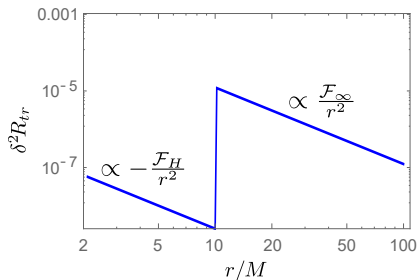
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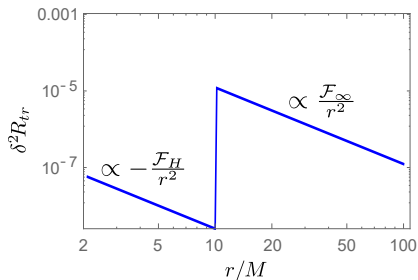
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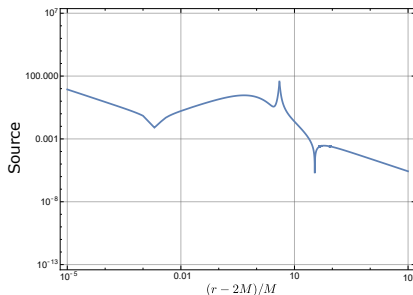
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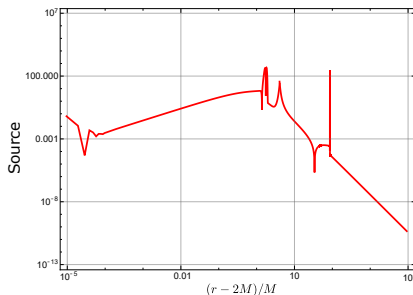
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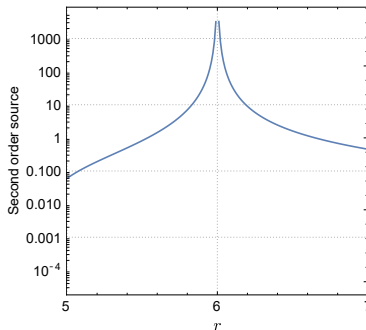
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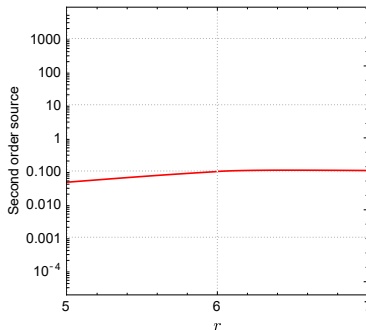
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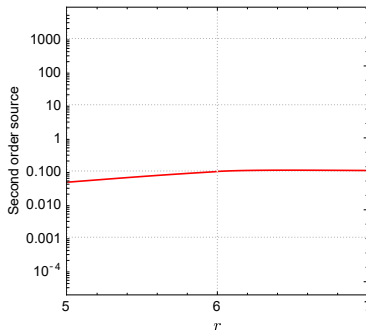
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# Conclusion

## Status of formalism

- basic formalism in place
- practical multiscale expansion under development

## Status of implementations

- everything is working properly for quasicircular orbits in Schwarzschild—just more labour needed